Snowmass Cosmic Frontier: Science Highlights

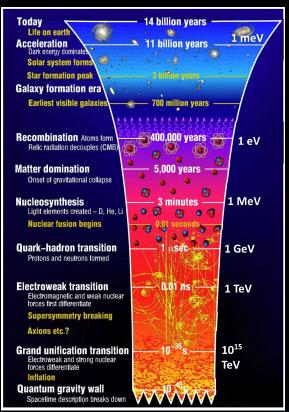
Tim M.P. Tait

University of California, Irvine





EPP: Progress & Promise November 29, 2022



The Cosmic Frontier seeks to understand the fundamental physics that governs the behavior of the Universe and its constituents.



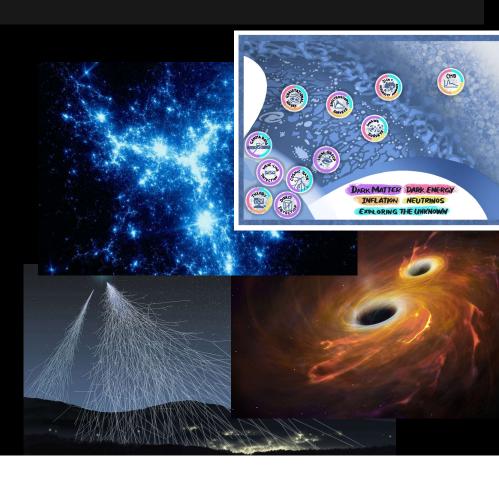




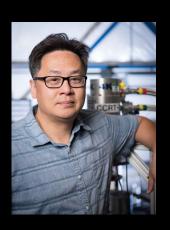
Cosmic Frontier

- The Cosmic Frontier hosts a wide range of research, spanning new cutting-edge detectors to novel analyses of data that touch on astroparticle physics and astronomy.
- It encompasses the interface between particle physics and cosmology and astrophysics.
- With the exception of neutrino masses, all of the current incontravertible evidence for physics beyond the Standard Model (dark matter, cosmic inflation, dark energy, and the baryon asymmetry of the Universe) are the direct result of Cosmic Frontier research.

Captions here, please



Cosmic Frontier Conveners



Aaron Chou Fermilab



Marcelle Soares-Santos University of Michigan



Tim Tait UC Irvine

Topical Subgroups & Conveners

Captions here, please

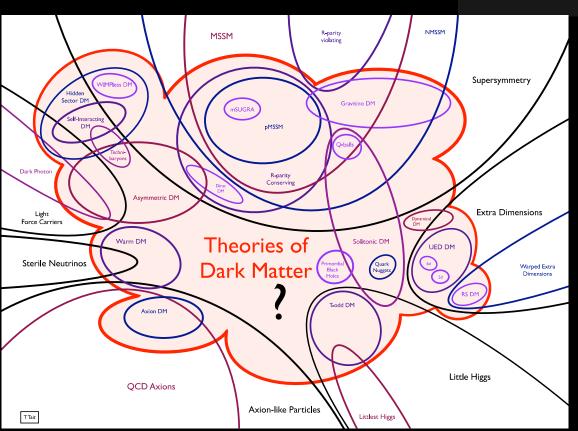
Topical Group	Co-conveners				
CF1: Particle DM	Jodi Cooley SMU	Tracy Slatyer MIT	Hugh Lippincott UCSB	Tien-Tien Yu Oregon	
CF2: Wavelike DM	Joerg Jackel Heidelberg	Gray Rybka UW	Lindley Winslow MIT		
CF3: DM Astro Probes	Alex Drlica-Wagner Fermilab	Chanda Prescod-Weinstein UNH	Haibo Yu UC Riverside		
CF4: DE & CA The Modern Universe	James Annis Fermilab	Jeff Newman Pittsburgh	Anze Slosar BNL		
CF5: DE & CA Cosmic Dawn & Before	Clarence Chang ANL	Laura Newburgh Yale	Deirdre Shoemaker Georgia Tech		
CF6: Dark Energy Complimentarity	Vivian Miranda Stony Brook	Brenna Flaugher Fermilab	David Schlegel LBNL		
CF7: Cosmic Probes	Rana Adhikari Caltech	Luis Anchordoqui CUNY	Ke Fang UW Madison	B.S. Sathyaprakash Penn State	Kirsten Tollefson MSU

Big Questions

Captions here, please

- What is the fundamental nature of the dark matter? How does it fit in with the Standard Model and what would we learn by detecting it?
 - Does it manifest as individual quanta (CFI)...or as collective waves (CF2)?
 - Can we further refine our understand of its properties based on cosmic observations (CF3)?
- What is the nature of dark energy and cosmic acceleration (CF4 & CF5)?
 - Is the dark energy dynamical? What is the physics of cosmic inflation? Are there other cosmological transitions whose existence we can infer?
 - Can we constrain or discover ultra-weakly interacting or super-heavy components of the Universe?
 - How can we use our existing and planned facilities to extract information that is more than the sum of the individual parts (CF6)?
- How can we use cosmic probes to learn about fundamental physics (CF7)?

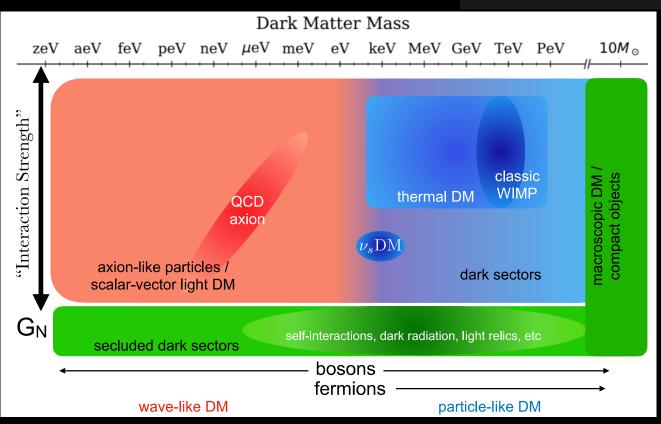
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The space of dark matter models encompasses a dizzying array of possibilities, representing many orders of mass and couplings.

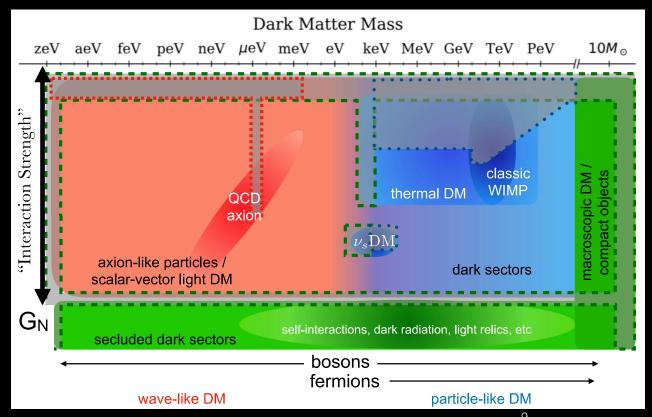
Snowmass 2013 organized the space of possibilities in terms of particle physics models, many of which have common features.

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Since Snowmass 2013, there has been something of a reorganization of the way we think about models of dark matter, to classify them more based on their common shared predictions for phenomena.

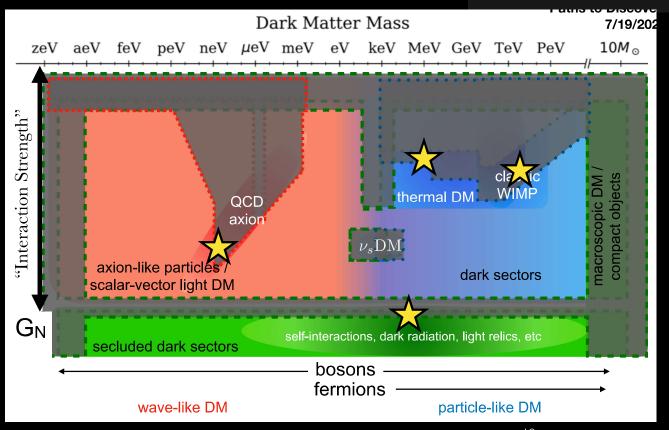
This 'cartoon' encapsulates a broad brush parameter space of DM mass versus coupling to the Standard Model.



The current situation has seen excellent exploration of some of the best-motivated candidates, including portions of parameter space corresponding to WIMPs, sterile neutrinos, and QCD axions.

Interesting parameter space in these well-defined models remains unexplored, and there are also vast unexplored regions.

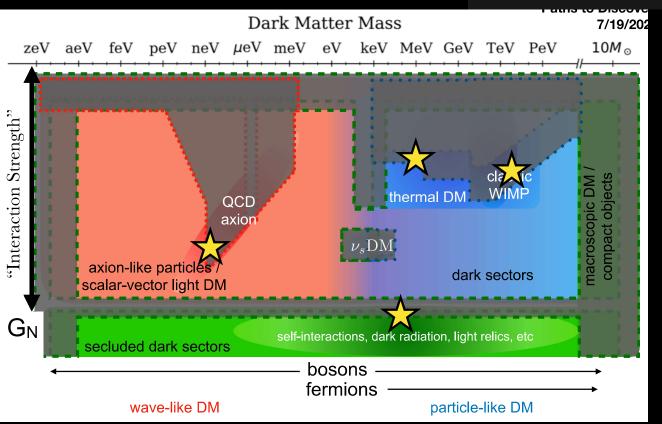
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And we have a plan!

'Delve Deep, Search Wide'
employs a range of direct searches
for WIMPs interacting with targets
on Earth, indirect searches for
annihilation products, and cosmic
probes based on structure, to
scrutinize priority targets such as
WIMPs and QCD axions, while
broadly scanning parameter space,
leaving no stone unturned.

Captions here, please



The next 10 years, including future Generation 3 direct searches for WIMPs and axions, combined with future indirect observatories, a program of smaller scale experiments, and key inputs from cosmic probes, results in broad coverage.

...or transformational discoveries!

Cosmological Probes

Captions here, please

Thanks to investments recommended by the previous P5, we are on the threshold of a dramatic expansion of available cosmological data, able to inform the deepest mysteries of fundamental physics.

Cosmic Expansion History
Cosmic Microwave Background
Growth of Structure
Gravitational Waves



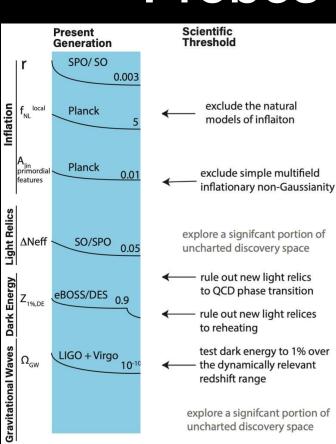






Cosmic Explorer

Cosmological Probes



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Current observations have revealed key properties of the Universe and its fundamental constituents. But there is still much left to learn.

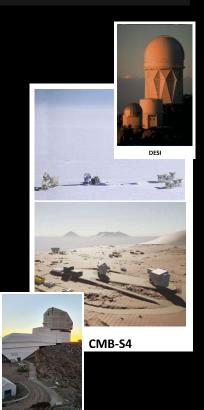
Cosmological **Probes**

Scientific Present **Ready This Threshold** Generation Decade SPO/SO 0.003 CMB-S4 0.0005 Inflation exclude the natural **Planck** models of inflaiton CMB-S4 S5SF **Planck** exclude simple multifield CMB-S4 + DESI features inflationary non-Gaussianity 0.002 **Light Relics** explore a significant portion of ΔNeff SO/SPO uncharted discovery space CMB-S4 0.03 rule out new light relics Dark Energy \$5SF eBOSS/DES 0.9 to QCD phase transition DESI/LSST 1.5 rule out new light relices to reheating S5SF **Gravitational Waves** test dark energy to 1% over LIGO + Virgo the dynamically relevant 05 redshift range explore a significant portion of uncharted discovery space

'Search Wide, Aim High'

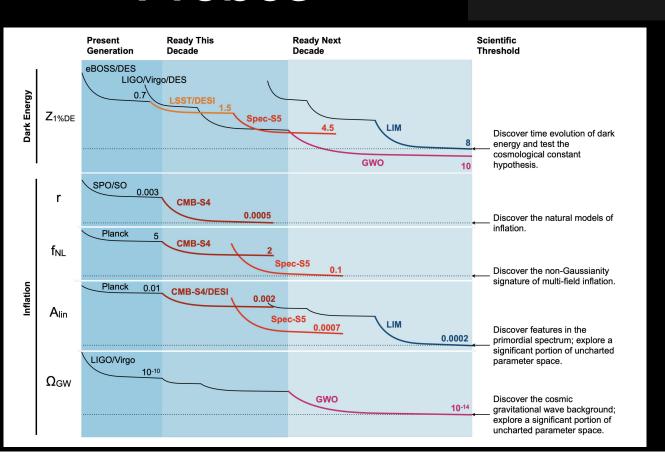
The near term future will be collecting the data from DESI, completing and executing Rubin/LSST, and constructing CMB-S4, currently on track with CD0.

This program will make unparalleled progress toward understanding the dynamics of cosmic inflation and search for new physics. 14



Cosmological Probes

Captions here, please

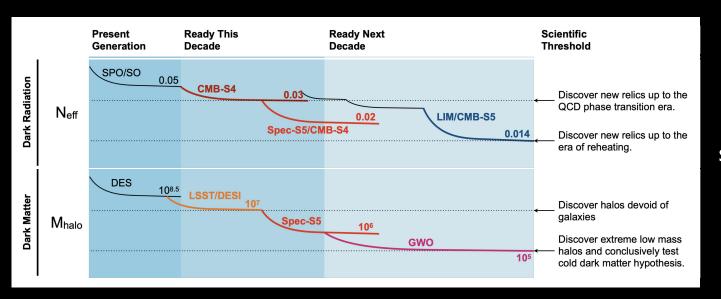


Longer term priorities are the roadmap to a future Stage V Spectroscopic Facility, and small projects and pathfinders toward new opportunities such as gravitational waves, 21cm, and Line-Intensity Mapping.

They offer unparalleled views on dark energy and inflation.

Cosmological Probes

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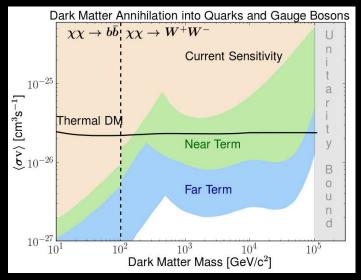
This program also offers unique opportunities to search for new physics and to understand the physics of the dark sector.

2014 P5 Science Drivers



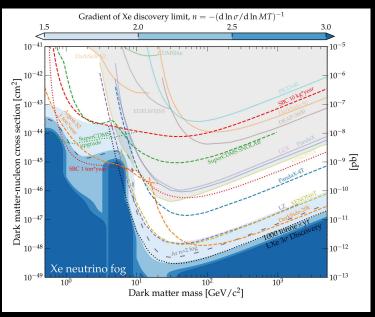


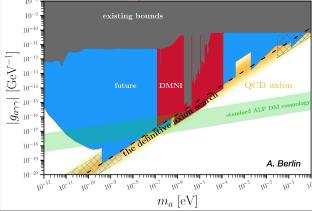
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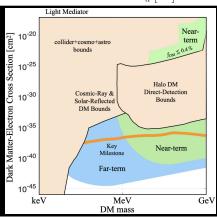


The Cosmic Frontier is required to verify that any candidate new physics we discover is actually making up the dark matter we see.

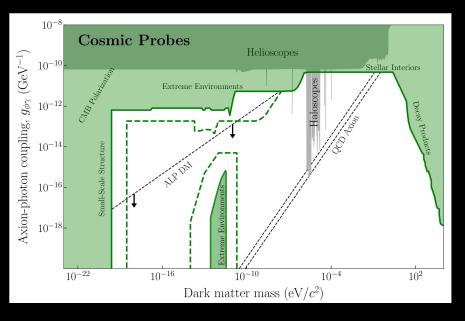
Delve deep, search wide!





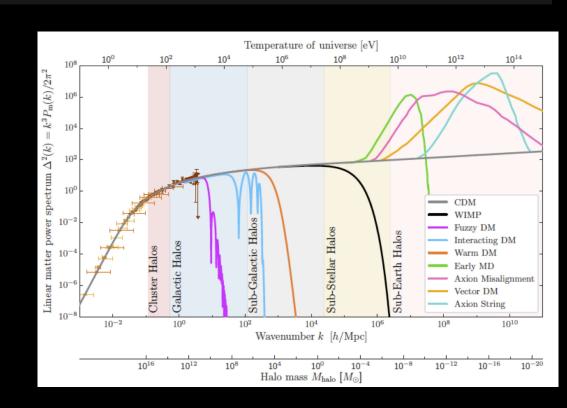




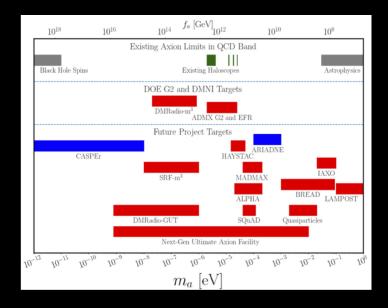


At the lightest allowed masses, ultralight bosonic dark matter is expected to manifest coherently as waves. The smallest scales of structure provide important tests of its mass and production mechanism.

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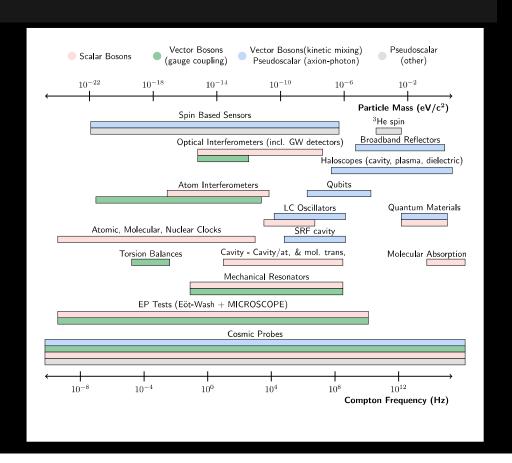




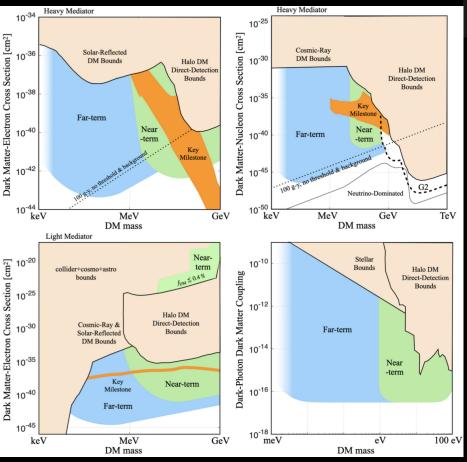


A variety of experiment can search for waves of dark matter. The QCD axion provides an important target, via both its coupling to photons and to gluons.

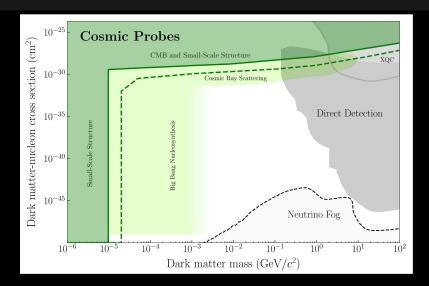
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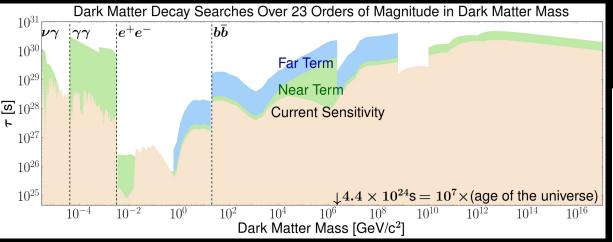
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Light particle dark matter (up to ~ GeV) is challenging to detect because of its limited energy deposition. Electron scattering and scattering off of collective excitations are key.

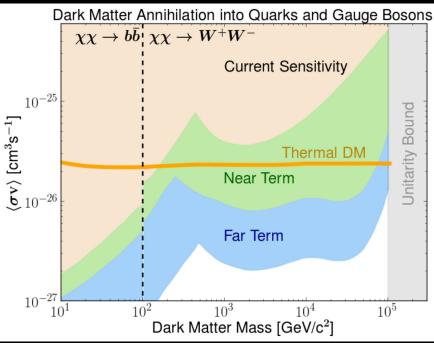


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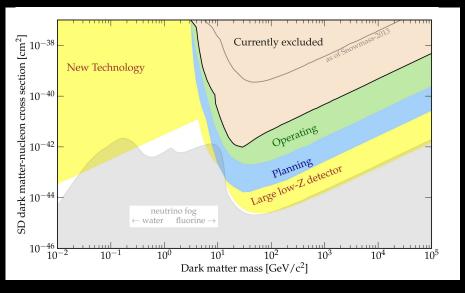


Indirect searches for dark matter decay products constrain the lifetime of the dark matter over a wide range of masses.

Annihilation into heavy particles results in key probes of the WIMP paradigm.

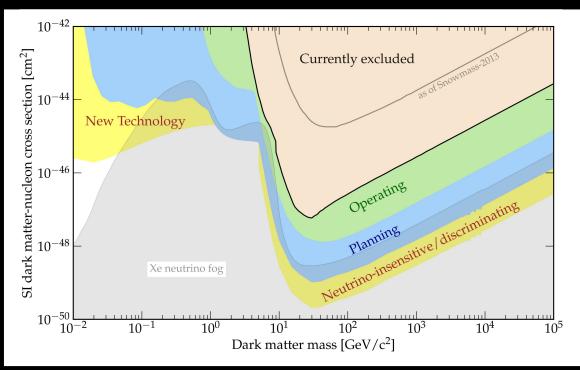






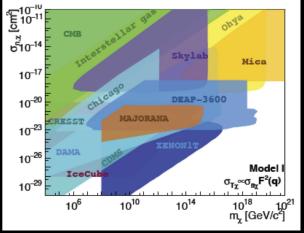
Direct searches for WIMPs have made incredible strides over the past two decades, and are now approaching the `neutrino fog'.

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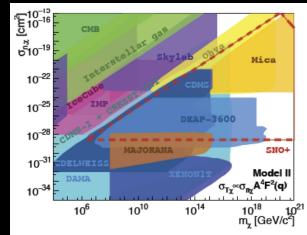




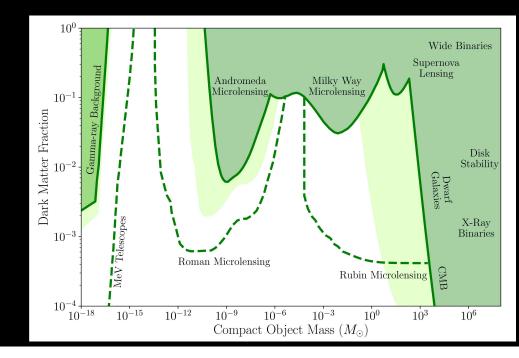
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At the largest masses, dark matter could take the form of dark blobs or primordial black holes. Novel detection strategies are required.

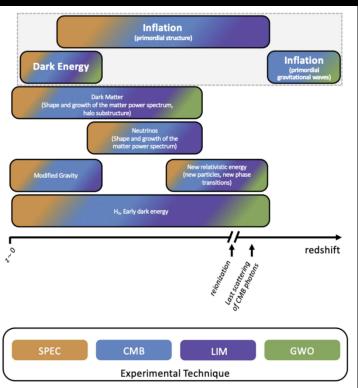


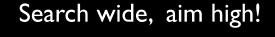
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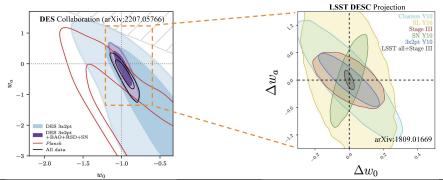
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Precision cosmology

The discovery of dark energy led to a precision measurement program to understand its physics.

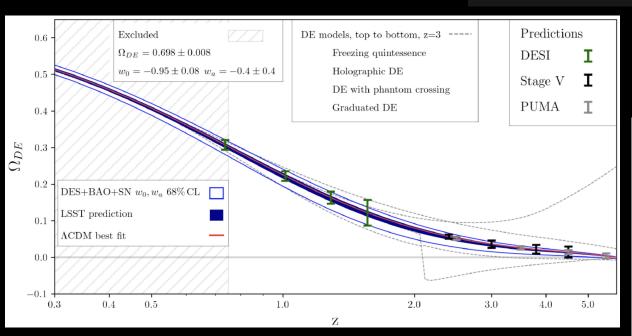


The Cosmic Frontier offers unique handles on the physics of inflation and to search for ultra-weakly interacting particles.

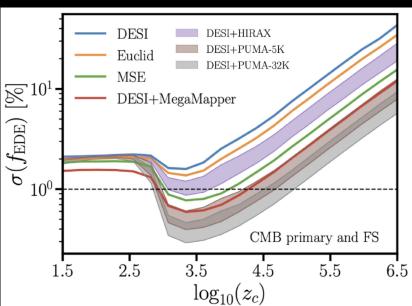
CMB-S4



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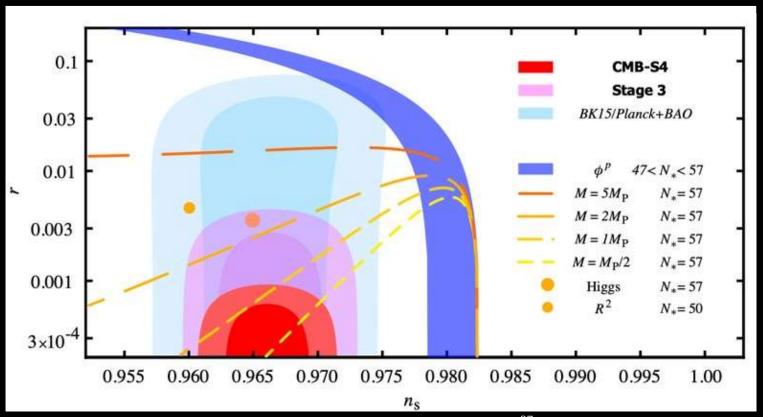


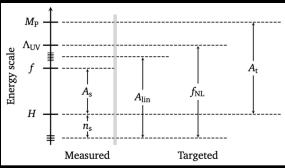
Precise measurements of dark energy in different epochs could shed light on models of Early Dark Energy aimed at resolving the Hubble tension.





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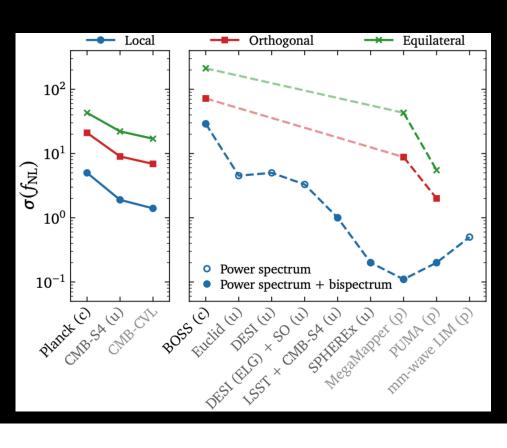


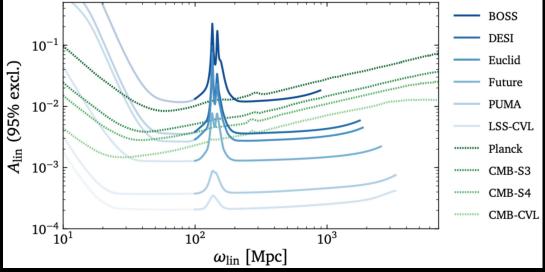


CMB-S4 offers a unique opportunity to access the Physics of inflation.



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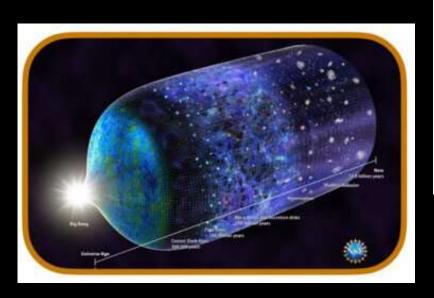


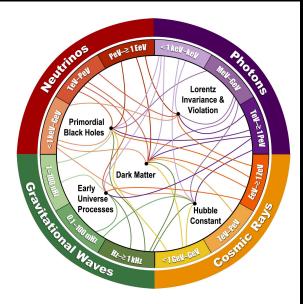
Features in the matter power spectrum, as revealed by template shapes or sharp local features can reveal new high energy dynamics probed via inflation.

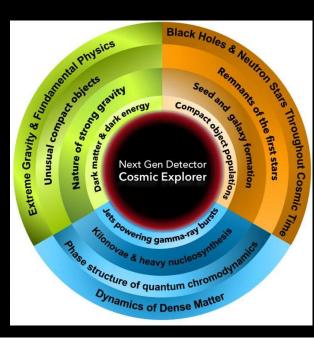


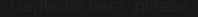
Explore the unknown: new particles, interactions, and physical principles.

Captions here, please



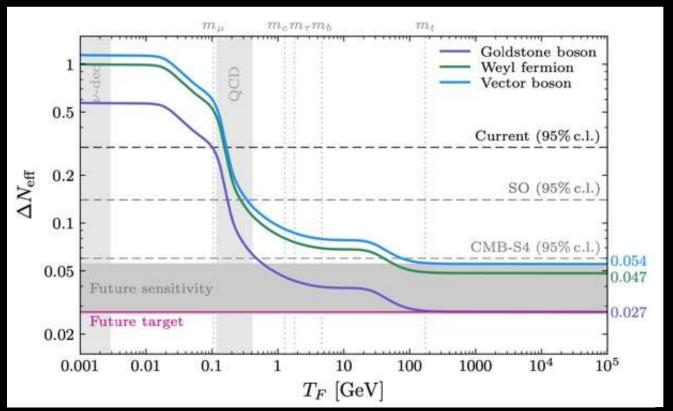








Explore the unknown: new particles, interactions, and physical principles.

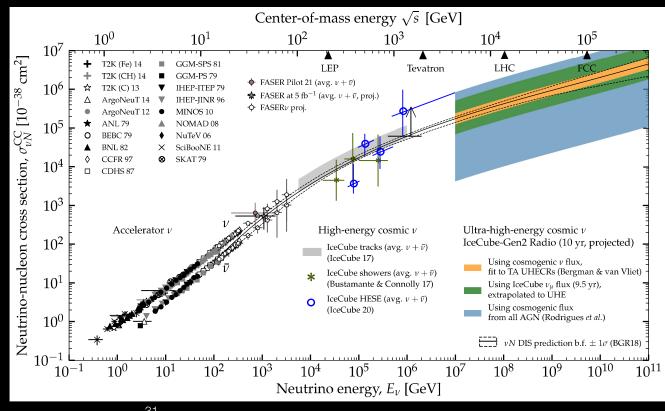


Cosmological probes offer unique ways to infer the existence of particles too weakly interacting for terrestrial probes, but present in the early Universe.



Explore the unknown: new particles, interactions, and physical principles.

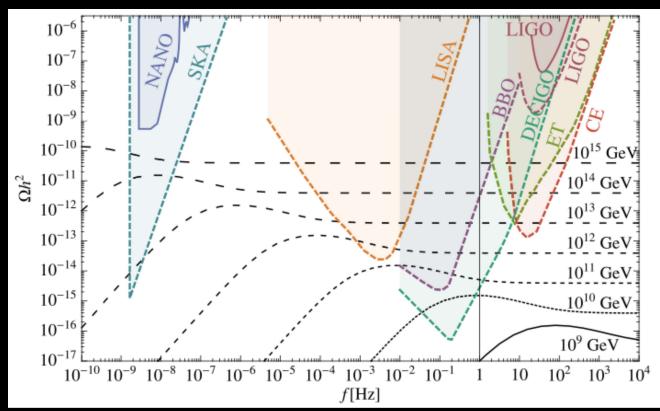
Natural high energy cosmic accelerators provide access to particles and interactions at the very highest energies.





Explore the unknown: new particles, interactions, and physical principles.

Captions here, please



The background of stochastic gravitational waves provides a window sensitive to early phase transitions.

E.g. Leptogenesis & Seesaw Dror, Hiramatsu, Kohri, Murayama, White 1908.03227 & PRL



Use the Higgs boson as a new tool for discovery.

Captions here, please

Higgs in Space!

C. B. Jackson^a, Géraldine Servant^b, Gabe Shaughnessy^{a,c}, Tim M.P. Tait^{a,c,d} and Marco Taoso ^{b,e}

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 ^b CERN Physics Department, Theory Division, CH-1211 Geneva 23, Switzerland
 ^c Northwestern University, 2145 Sheridan Road, Evanston, IL 60208, USA
 ^d Department of Physics and Astronomy, University of California, Irvine, CA 92697. USA
 ^e IFIC (CSIC-Universitat de València), Ed.Instituts, Apt. 22085, 46071 Valencia, Spain

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 g-shaughnessy@northwestern.edu, marco.taoso@ific.uv.es

The Higgs may not be the 'go to' tool for disovery in the Cosmic Frontier, but it has its important role to play!

Higgs inflation

Javier Rubio

Institut für Theoretische Physik, Ruprecht-Karls-Universität Heidelberg, Philosophenweg 16, 69120 Heidelberg, Germany

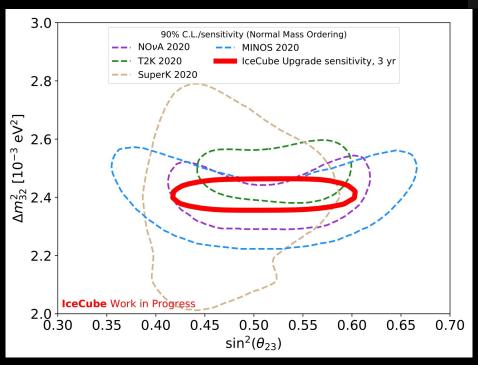
Abstract

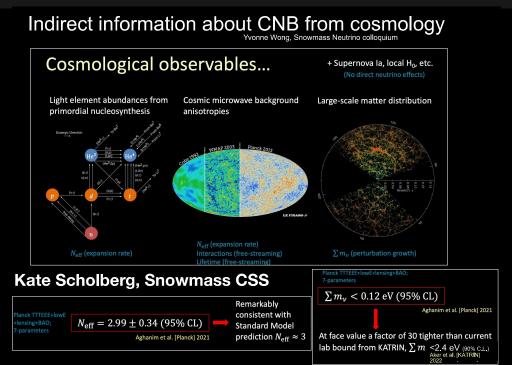
The properties of the recently discovered Higgs boson together with the absence of new physics at collider experiments allows us to speculate about consistently extending the Standard Model of particle physics all the way up to the Planck scale. In this context, the Standard Model Higgs non-minimally coupled to gravity could be responsible for the symmetry properties of the Universe at large scales and for the generation of the primordial spectrum of curvature perturbations seeding structure formation. We overview the minimalistic Higgs inflation scenario, its predictions, open issues and extensions and discuss its interplay with the possible metastability of the Standard Model vacuum.



Pursue the physics associated with neutrino mass.

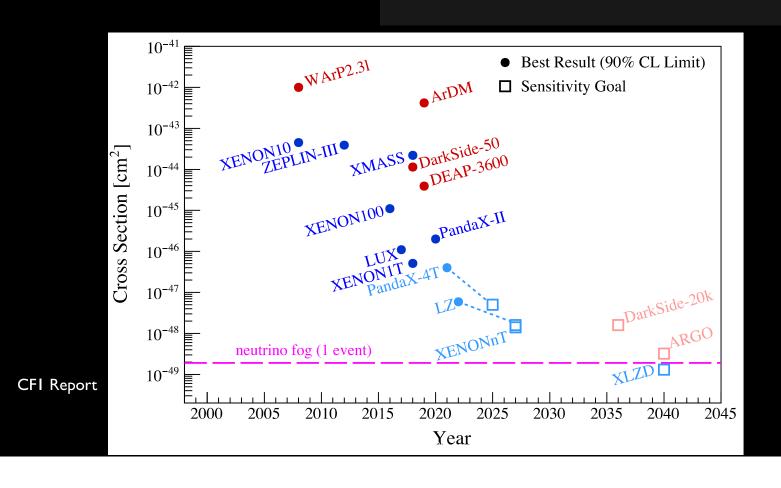
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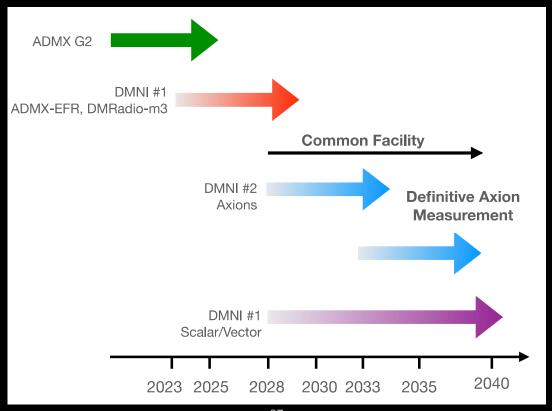
Cosmic Frontier observations offer unique and important perspectives on the parameters describing neutrino masses which can inform terrestrial searches.

Timelines

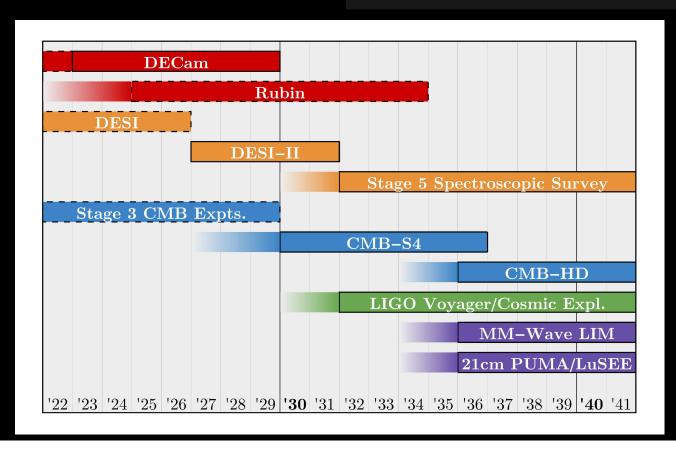


Liquid nobel DM Searches

CF2 Report

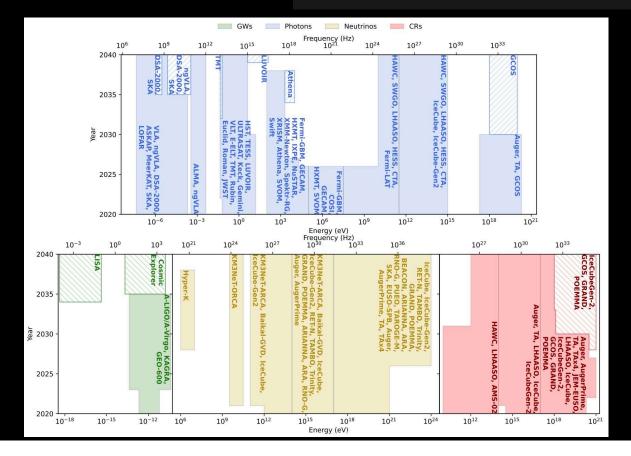


Wave DM Searches



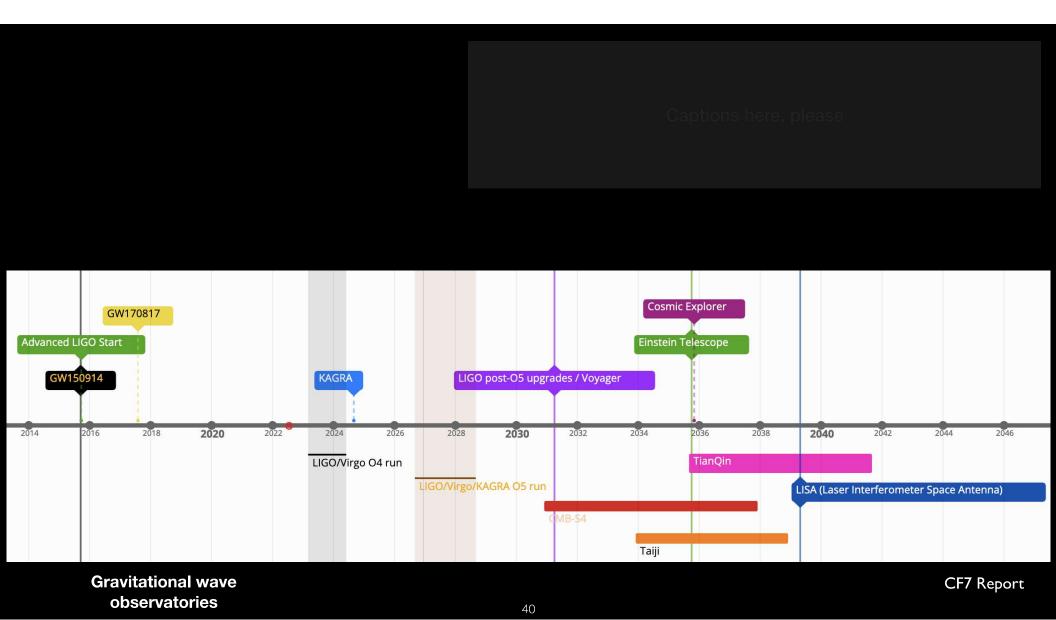
Cosmic Surveys

CF6 Report



High energy Cosmic particle Detectors

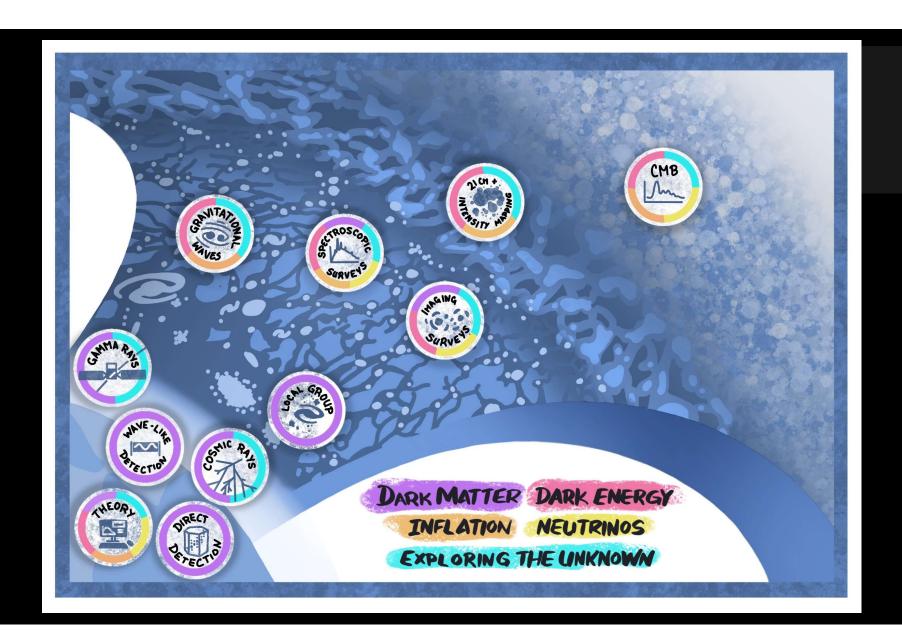
CF7 Report



Delve Deep. Search Wide. Aim High.

The Cosmic Frontier contributes unique opportunities to the pursuit of fundamental physics.

It has strong synergies with the Energy, Neutrino, and Rare Processes Frontiers, benefits from essential connections with the Theory, Computational, Instrumentation, and Underground Facilities Frontiers. It also offers interesting opportunities for the Community Engagement Frontier.



THANK YOU

