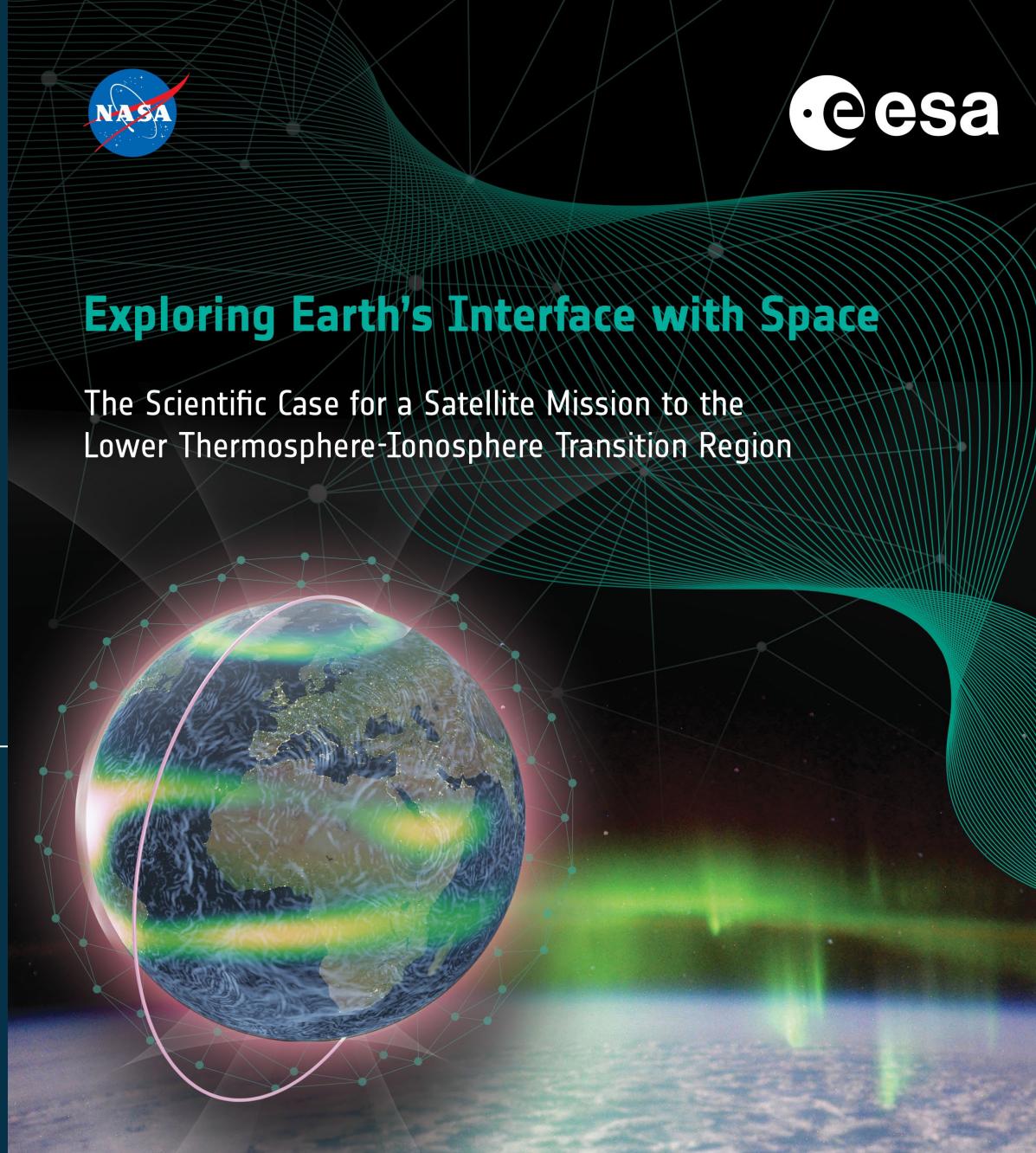


ENLoTIS: ESA-NASA Lower Thermosphere Ionosphere Science Report

Jeff Thayer
University of Colorado

2024-09-24

ESA/NASA (2024). Exploring Earth's Interface with Space – The Scientific Case for a Satellite Mission to the Lower Thermosphere-Ionosphere Transition Region, European Space Agency (ESA), Noordwijk, The Netherlands and National Aeronautics and Space Administration (NASA), Washington DC, United States of America, ESA-EOPSM-ELTI-RP-4592, 128 pp., <http://doi.org/10.5270/ESA-NASA.LTI-SC.2024-07-v1.0>



We Live in an Electrified World

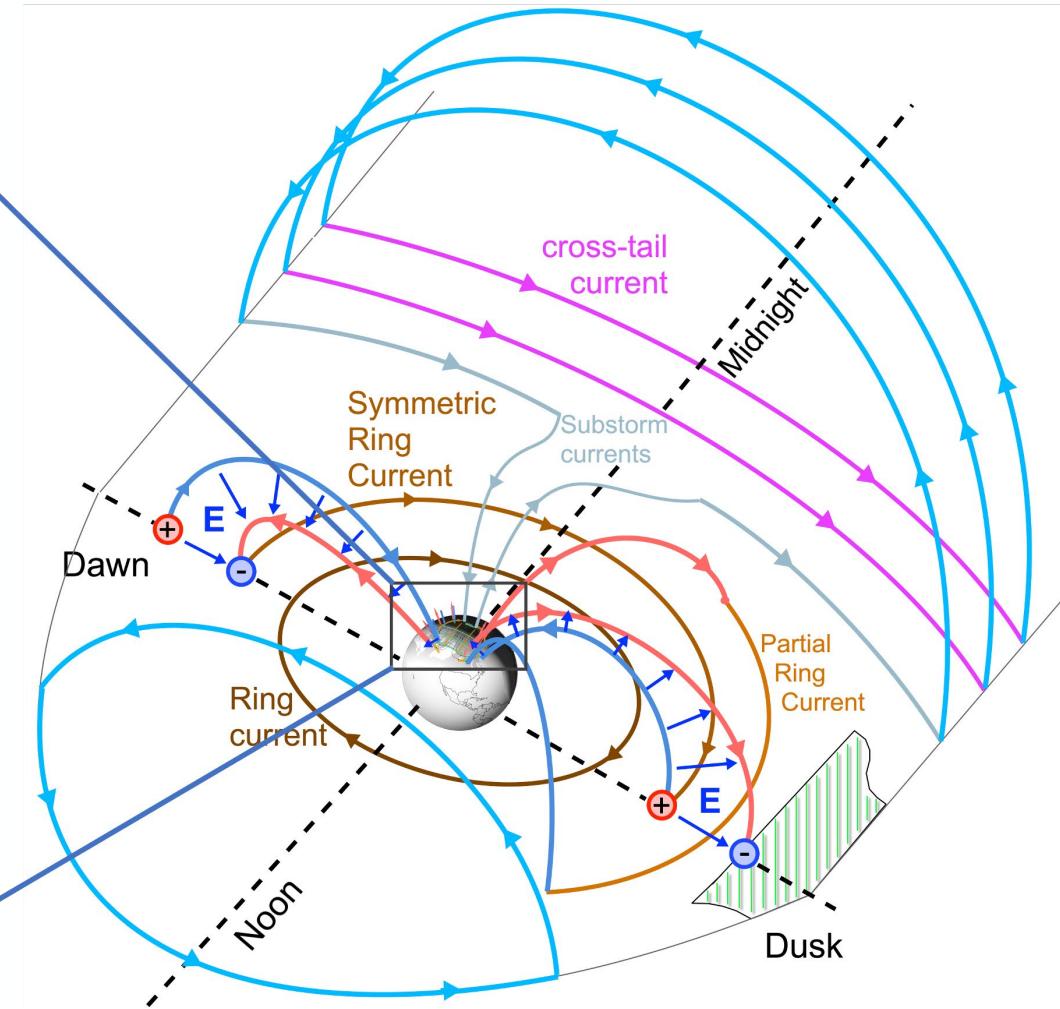
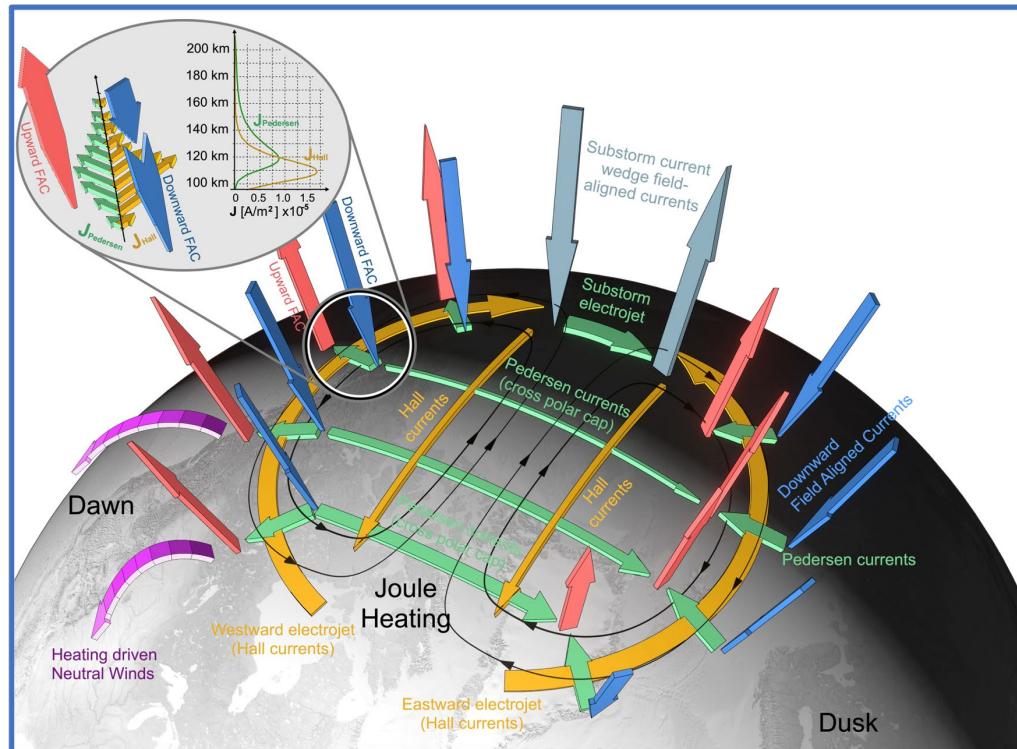


Part of a Global Electric Circuit



Part of a Planetary Electric Circuit

Electrified Geospace: A Planetary Electric Circuit



What is ENLoTIS?

https://science.nasa.gov/science-news/NASA_and_ESA_Exploring_New_Joint_Satellite_Mission_Concepts

ENLoTIS Working Group Members

Jean-Jacques Berthelier	CNRS / LATMOS, France
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Astrid Maute	Univ. Colorado/CIRES, USA
Minna Palmroth	Univ. of Helsinki, Finland
Noora Partamies	Univ. Centre in Svalbard, Norway
Gareth Perry	New Jersey Inst. of Technology, USA
Robert Pfaff	Goddard Space Flight Center, USA
Theodoros Sarris	Democritus Univ. of Thrace, Greece
Claudia Stolle	Leibniz Inst. of Atm. Phys., Germany
Jeff Thayer	Univ. of Colorado at Boulder, USA
Sarah Vines	Applied Physics Lab/JHU, USA

- ❑ EN-LoTIS Working Group explores agency cooperation on future lower thermosphere-ionosphere (LTI) satellite mission concepts, targeting *in situ observations* that advance understanding of neutral-ion interactions from 100 - 200 km altitude and the ionospheric E region.
- ❑ Concept of low-flying LTI mission poses unique scientific & technical challenges. Joint ESA/NASA collaboration proposed to help address these challenges.
- ❑ Initial phase of WG provides information via science study report to help agencies plan possible future joint mission development.
- ❑ **Report now available.**

ENLoTIS Steering Committee: **John McCormack, Alex Hoffmann**

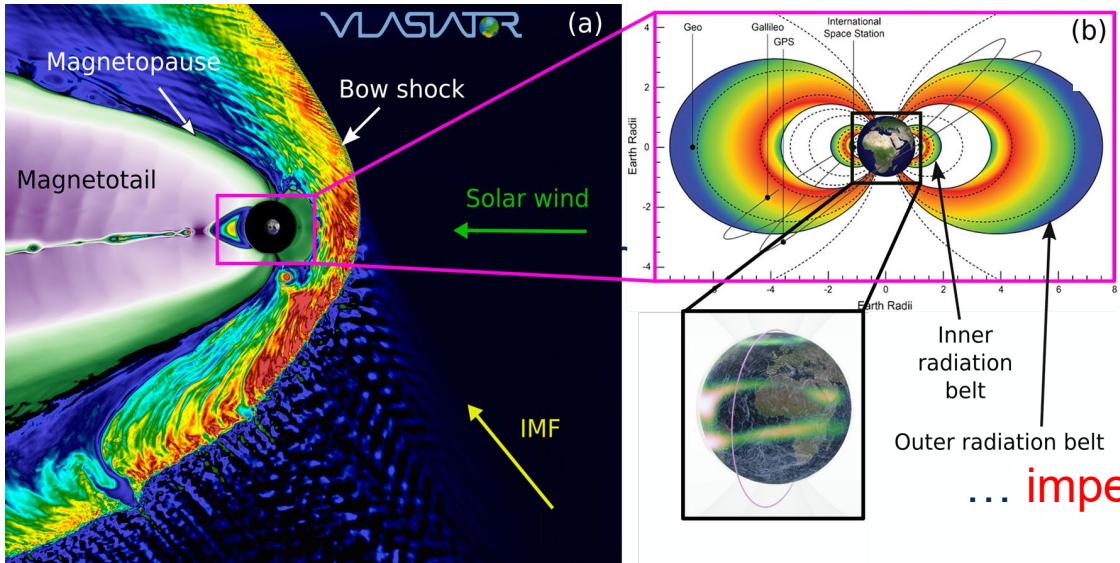
David Cheney, Larry Kepko, Anja Stromme, Matt Taylor

Why ENLoTIS?

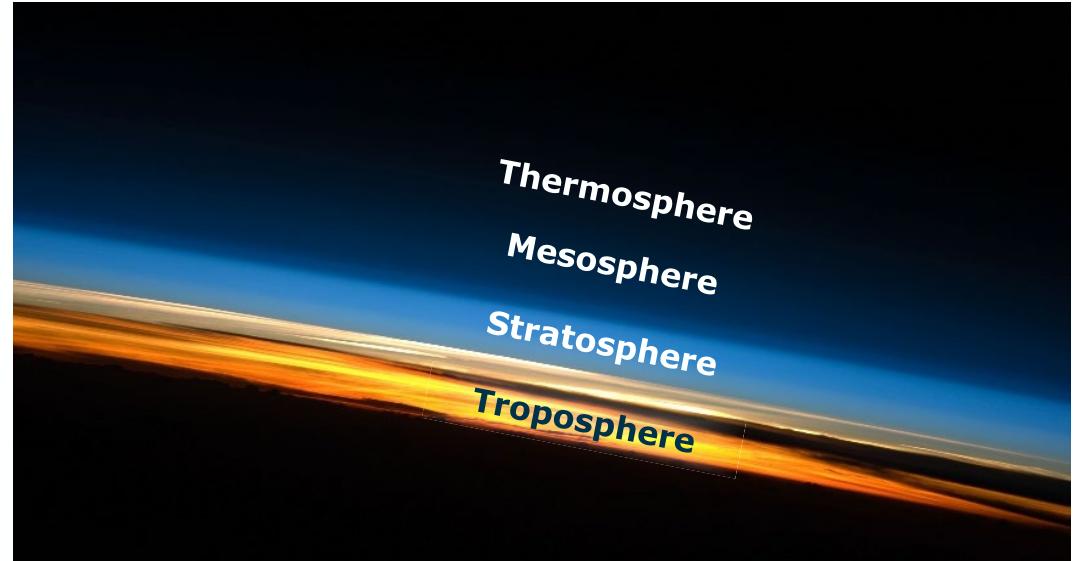


Inadequate knowledge of the transition from atmosphere to space ...

... prevents understanding of how Earth's atmosphere and space are linked



... impedes efforts to understand the Sun-Earth system as a whole



... inhibits accurate predictions of the behavior of humanity's space-faring and space-reliant systems

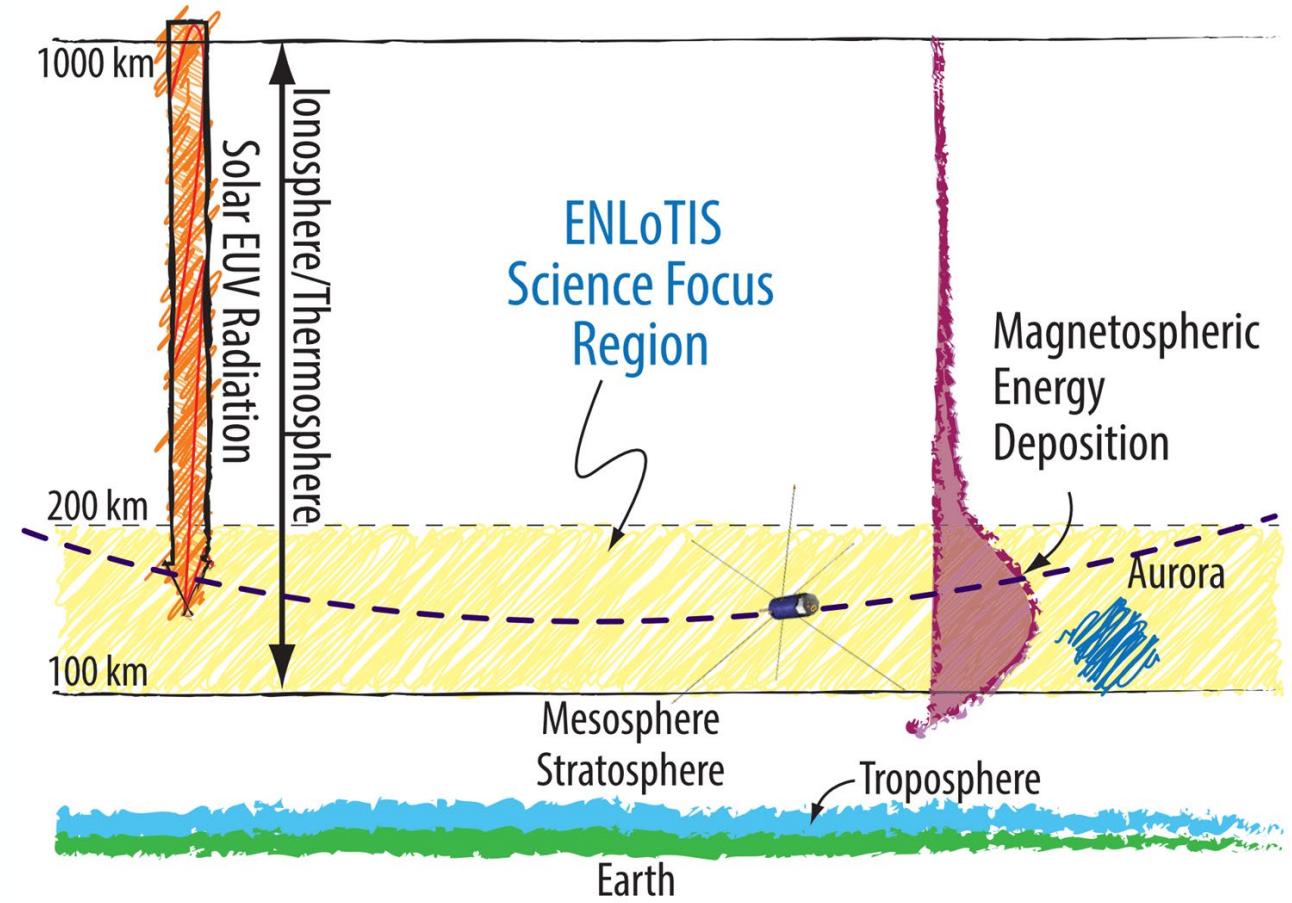
EN-LoTIS WG is studying the first systematic, comprehensive *in situ* exploration of a collision-dominated, neutral-plasma space environment.

LTI behavior consists of interactions among commingled matter and fields:

- Neutral gas (thermosphere)
- Plasma (ionosphere)
- Electric and magnetic fields

Frequent collisions between neutral and charged particles results in emergent behavior* not present in simpler systems

Global understanding of LTI behavior has been inhibited by lack of *in-situ*, multi-property measurements

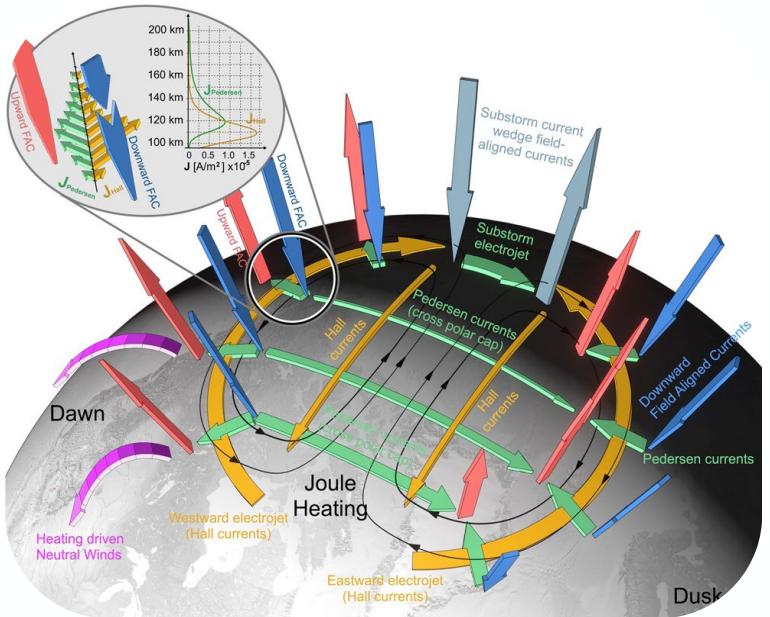


*Emergent behavior: Behavior that arises out of the interactions between parts of a system and which cannot easily be predicted or extrapolated from the behavior of those individual parts.

The LTI: Where Newton and Maxwell meet

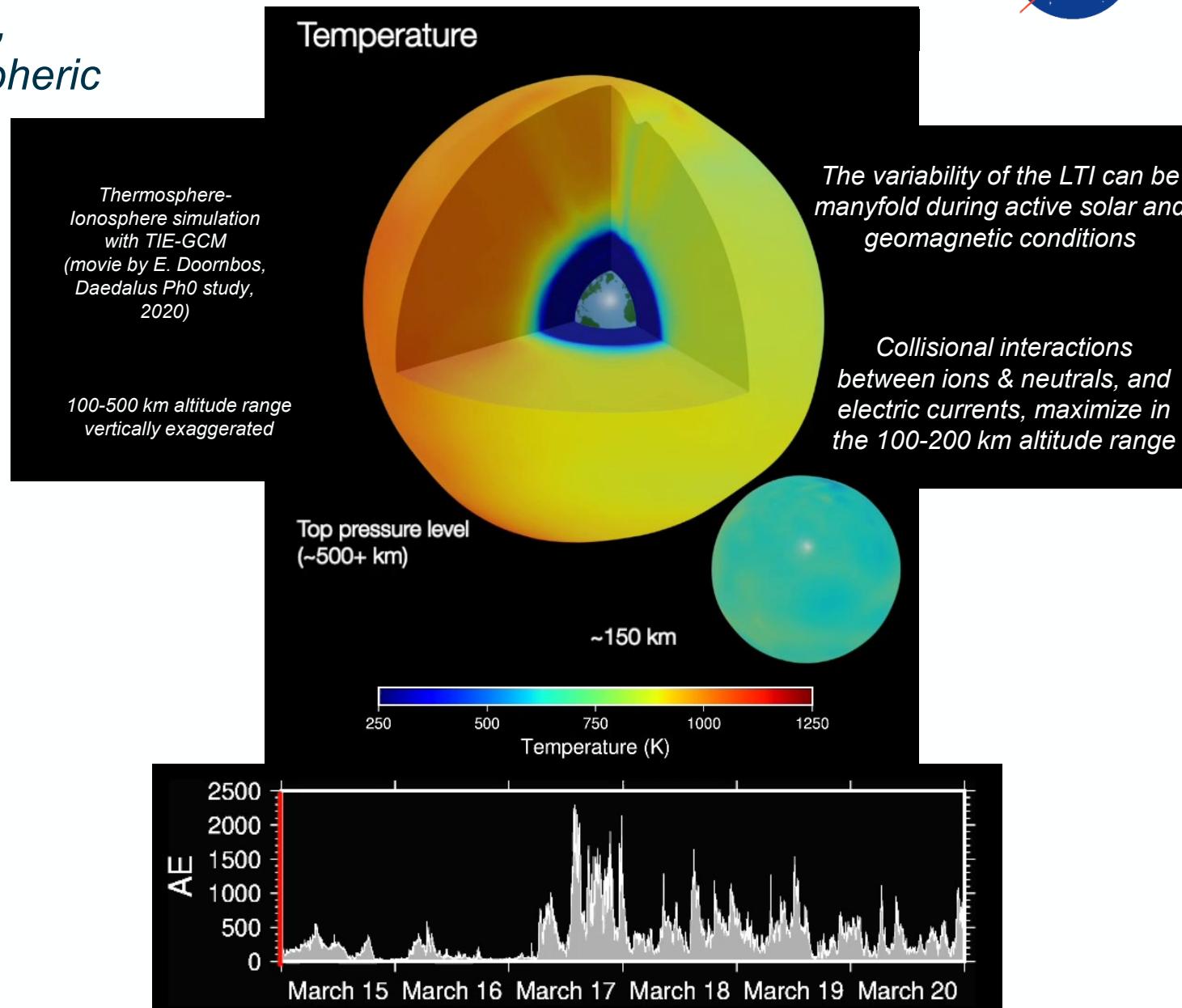
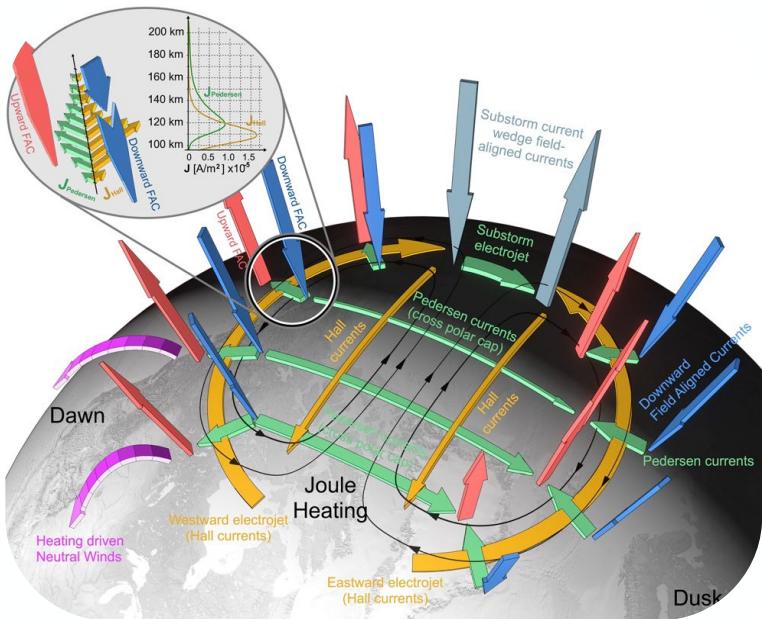


*Electric currents close in the LTI,
coupling the Earth to magnetospheric
electrodynamics*



The LTI: Where Newton and Maxwell meet

*Electric currents close in the LTI,
coupling the Earth to magnetospheric
electrodynamics*

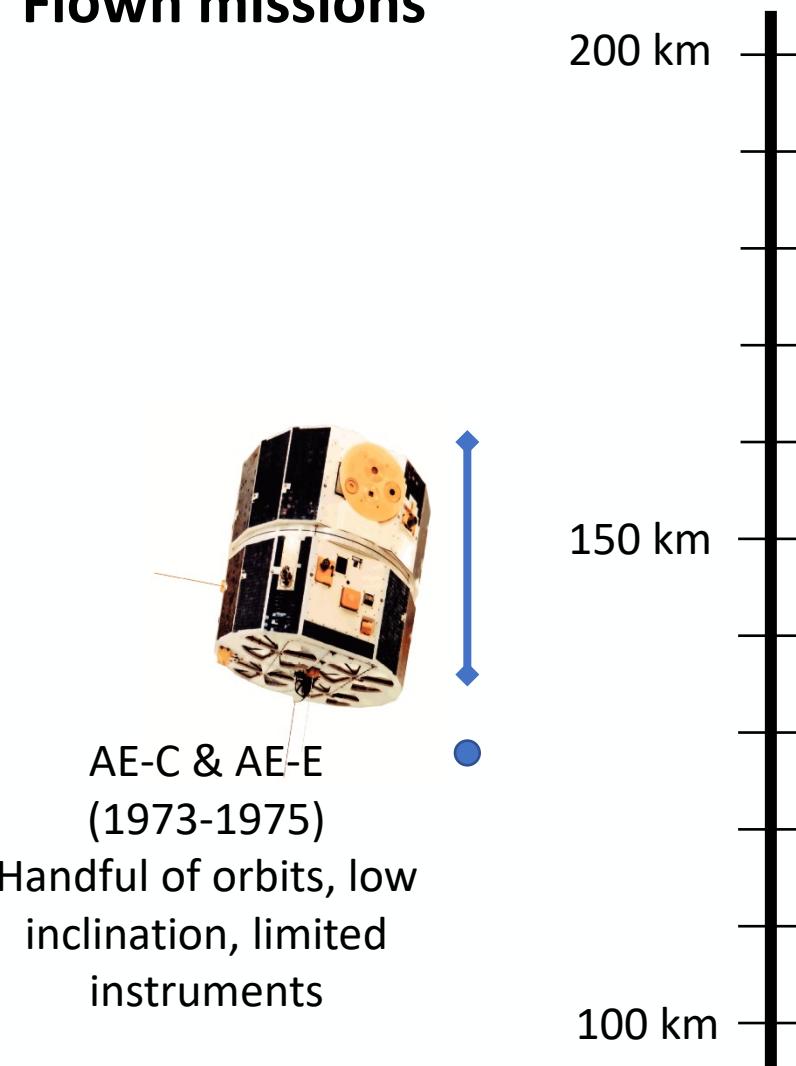


Programmatic context

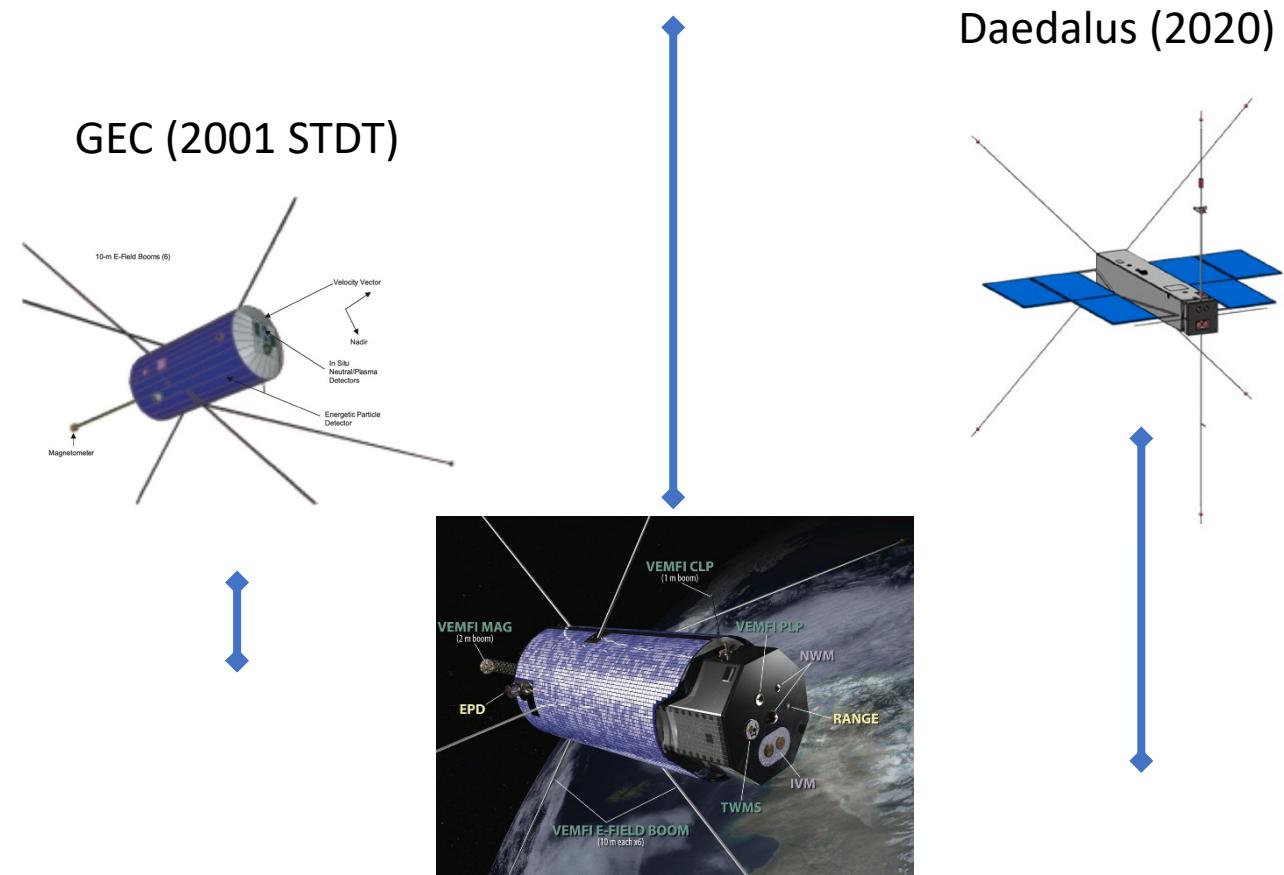
The past, present and future of LTI missions



Flown missions

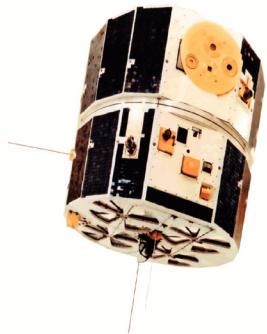


Proposed missions

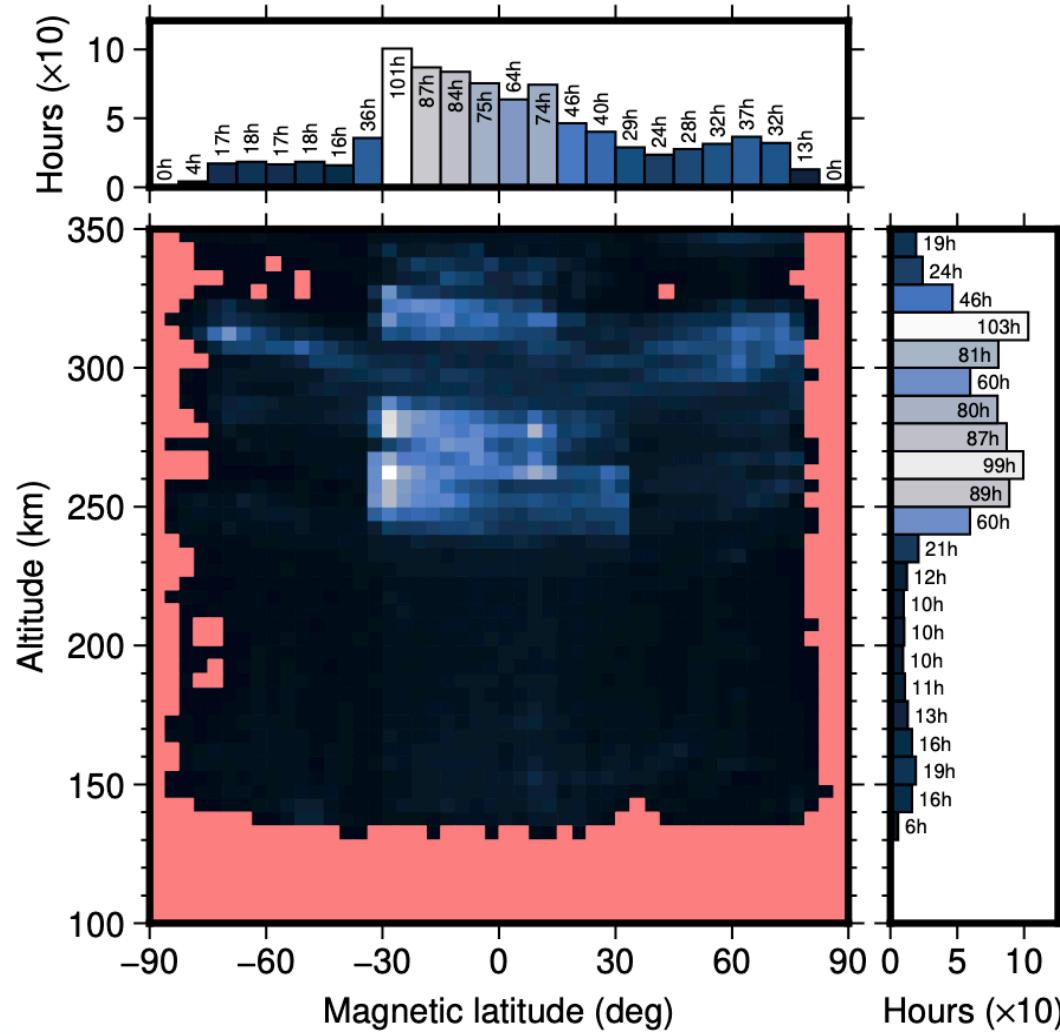


LOPEX, Dipper, MAX, ASTRE,
ASTRE2 (1993-2011)

LTI In-Situ Measurements of the Past

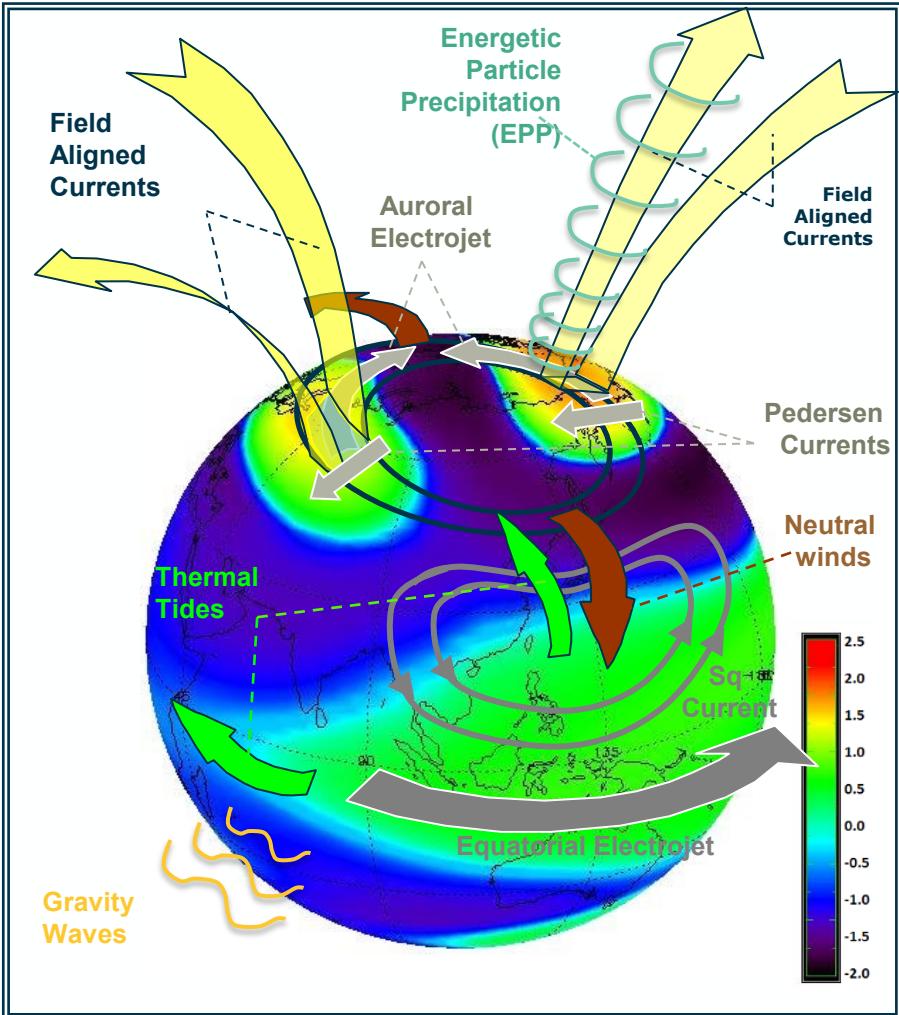


AE-C & AE-E
(1973-1976)



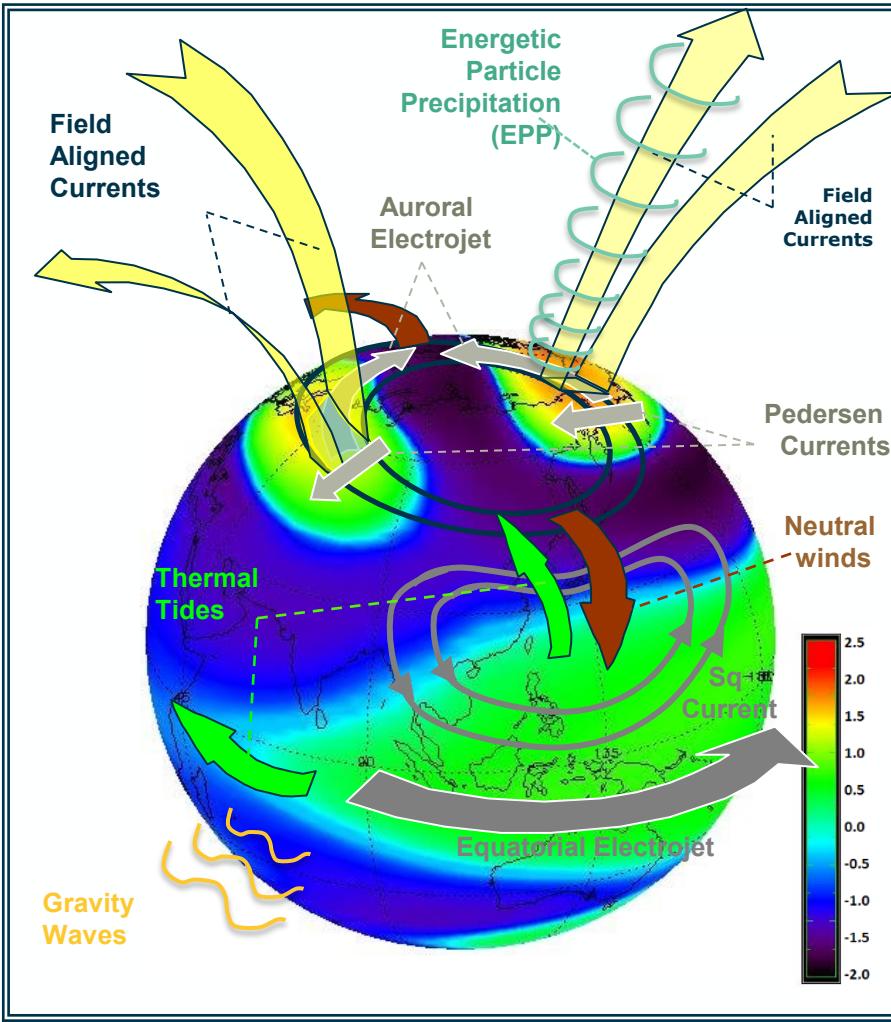
Total observation time < 200 km:
~60 hours

ENLoTIS Report: Scientific Goal & Objectives



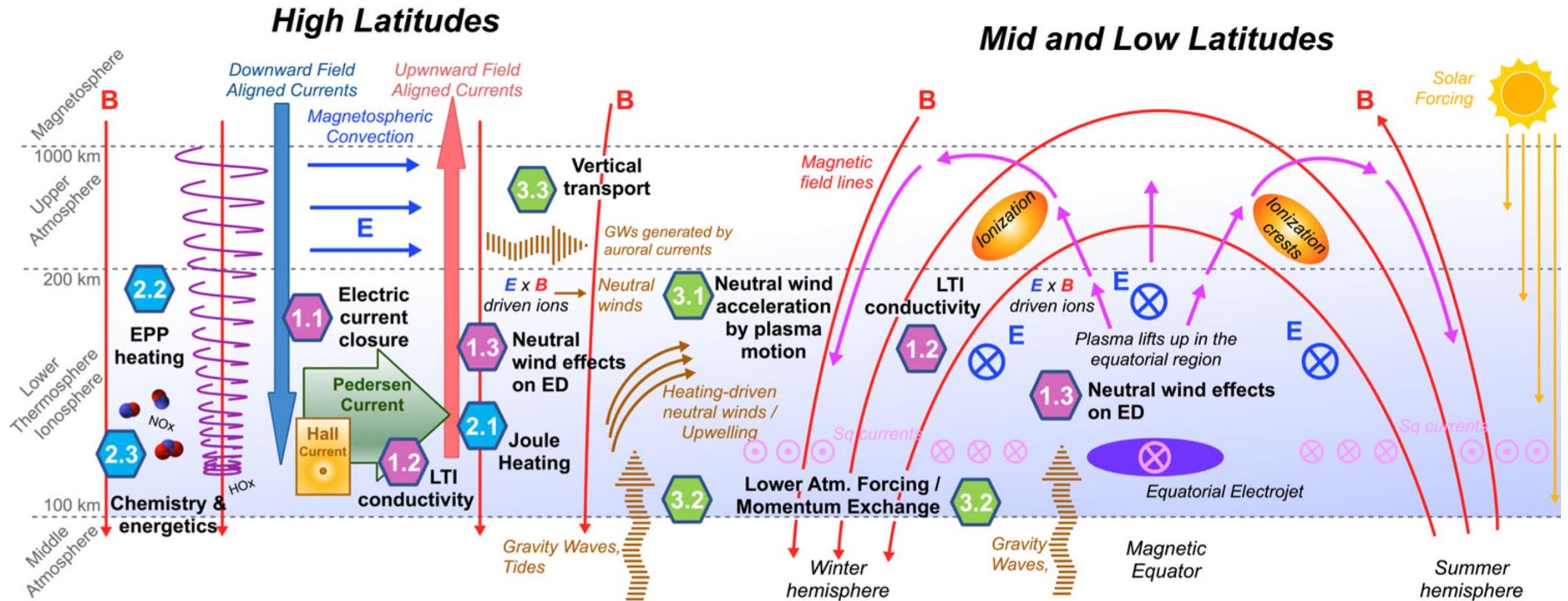
“To understand how the transition from Earth’s atmosphere to space is governed by the fundamental plasma-neutral interactions that are intrinsic to the lower thermosphere-ionosphere (LTI) currents.”

ENLoTIS Report: Scientific Goal & Objectives

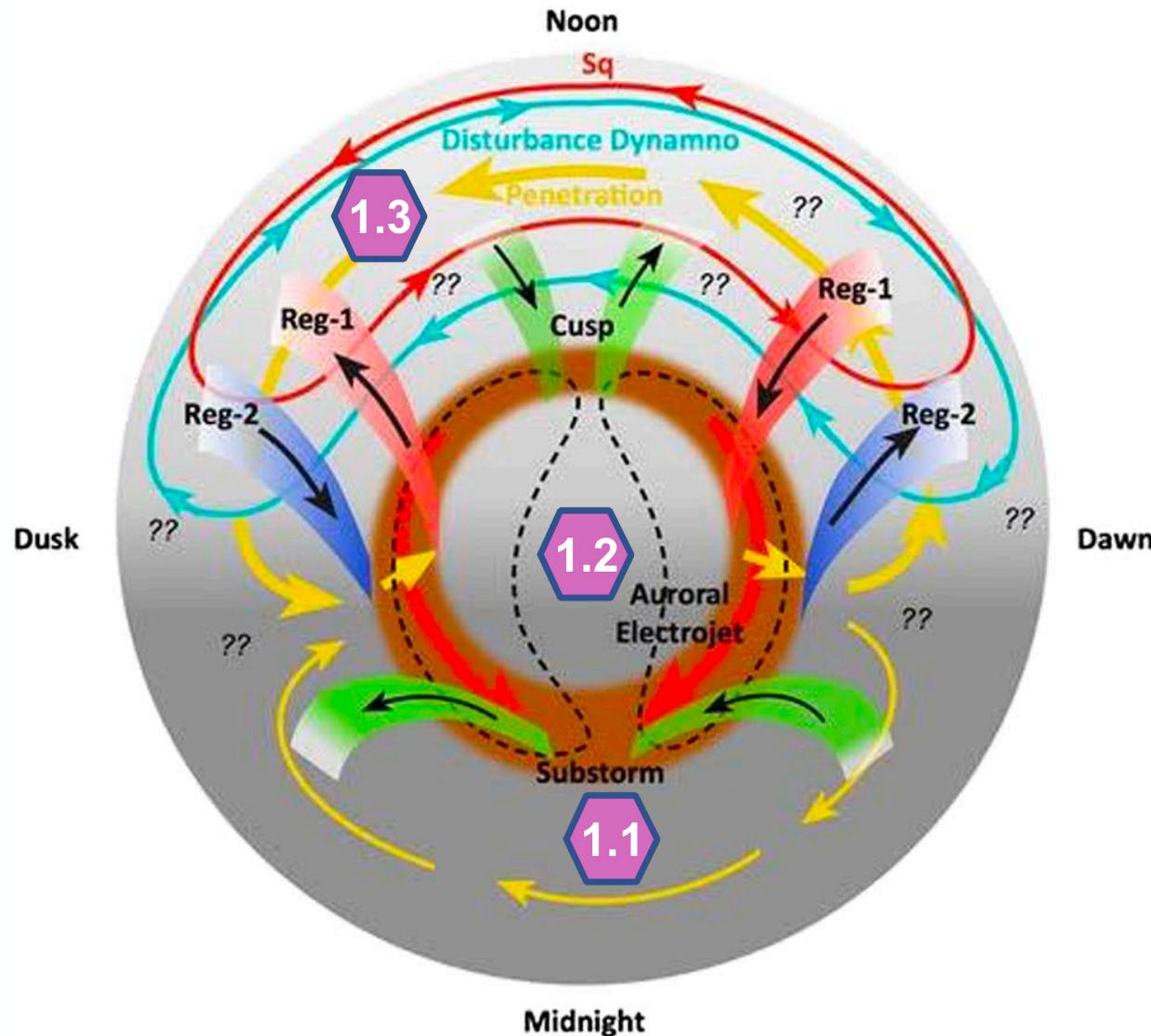


SO1: Collisional Electrodynamics J	SO2: Collisional Energetics $J \cdot E$	SO3: Collisional Dynamics $J \times B$
<p>1.1 Determine how electric currents flow and close in the LTI, and thereby couple to the mag/spheric electrodynamics</p>	<p>2.1 Determine how Joule (frictional) heating in the LTI depends on scale size, altitude and neutral winds.</p>	<p>3.1 Determine how neutral winds in the LTI are accelerated by plasma motions via ion-neutral collisions</p>
<p>1.2 Understand how the various LTI processes act to determine the Hall & Pedersen conductivity</p>	<p>2.2 Determine how energy from energetic precipitating particles (EPP) directly heats the LTI.</p>	<p>3.2 Discover how the exchange of momentum across scales by means of lower atmospheric forcing manifest in the LTI.</p>
<p>1.3 Determine the effect of the neutral winds on the LTI electrodynamics.</p>	<p>2.3 Determine how plasma-neutral collisions cause chemical changes that affect the energetics of the LTI.</p>	<p>3.3 Determine how collisional processes drive vertical transport & cause composition changes</p>

Pictorial Form of SubObjectives

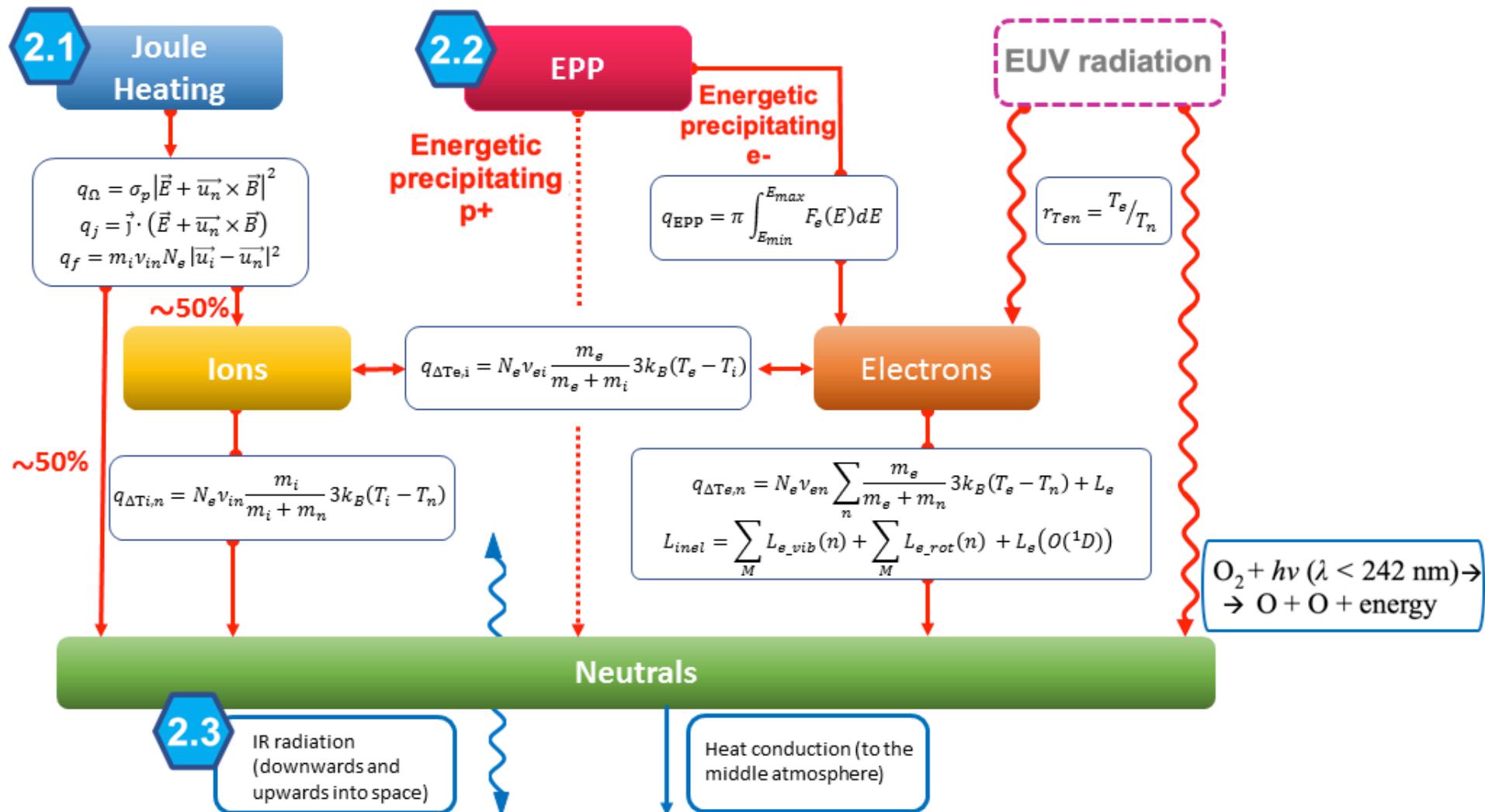


Collisional Electrodynamics

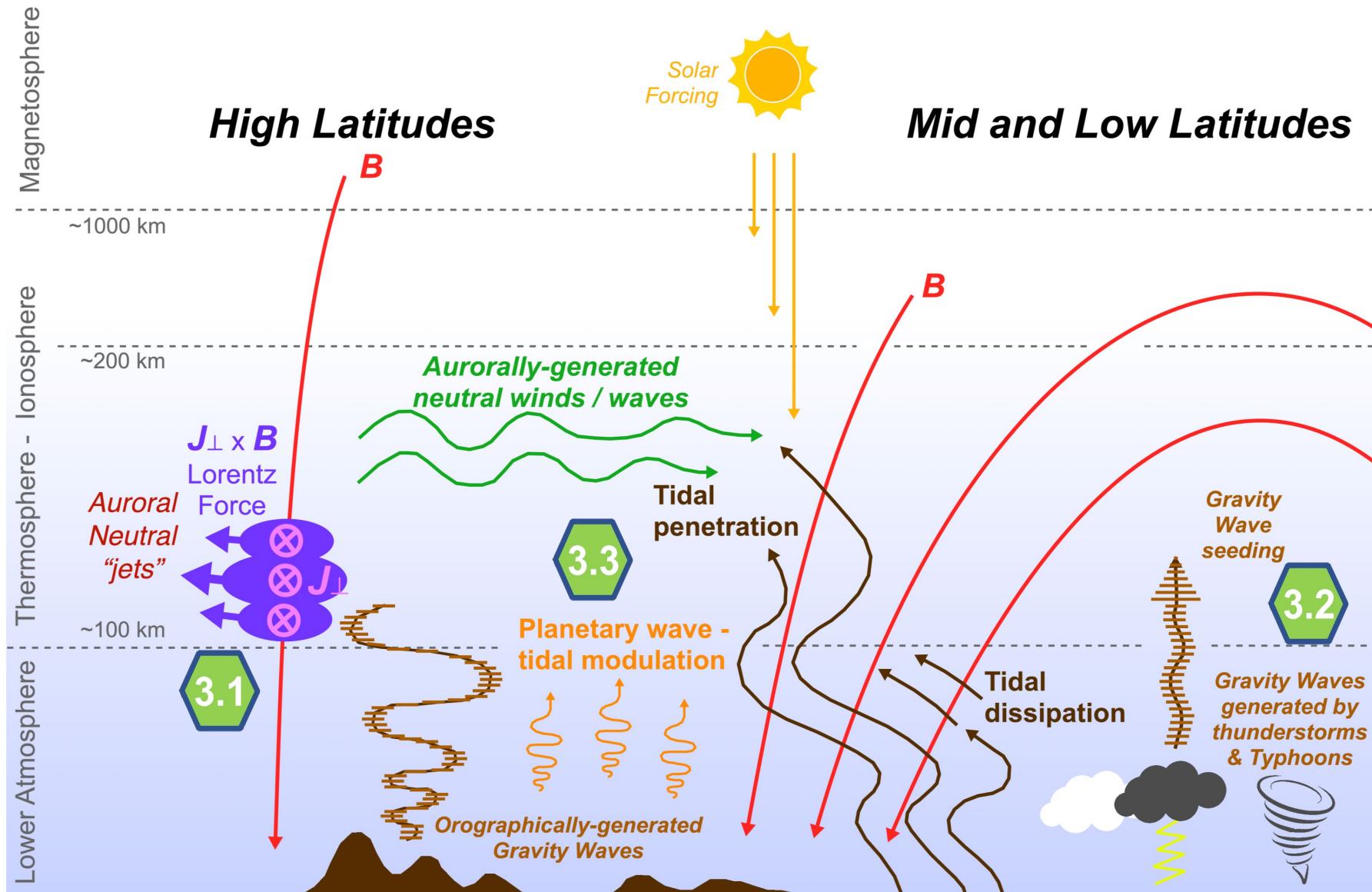


Collisional Energetics

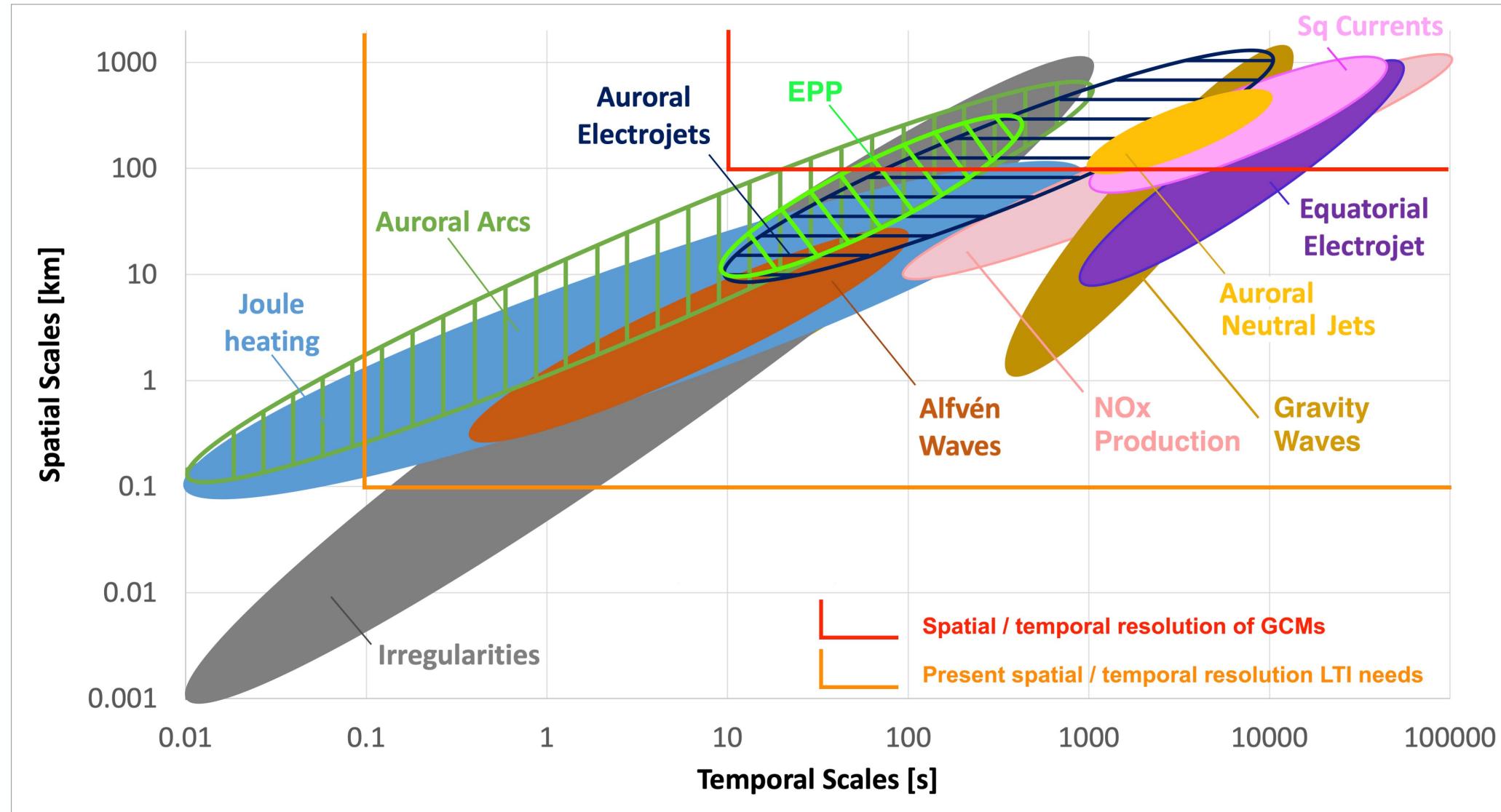
Energetics and Energy Flow in the LTI



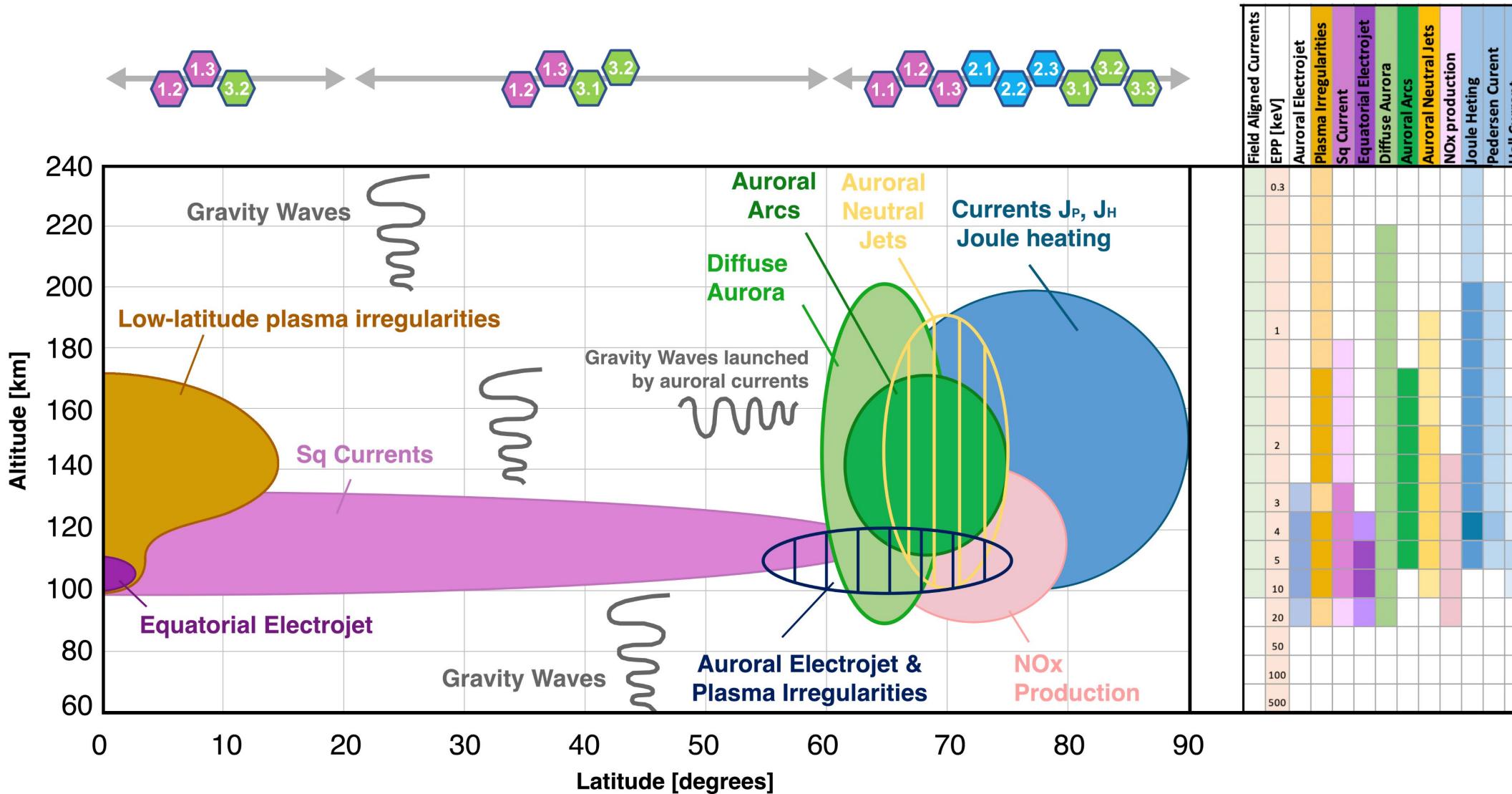
Collisional Dynamics



Spatial and Temporal Scales of LTI Phenomena



Focus Regions of LTI Phenomena



ENLoTIS Report: Measurement Requirements

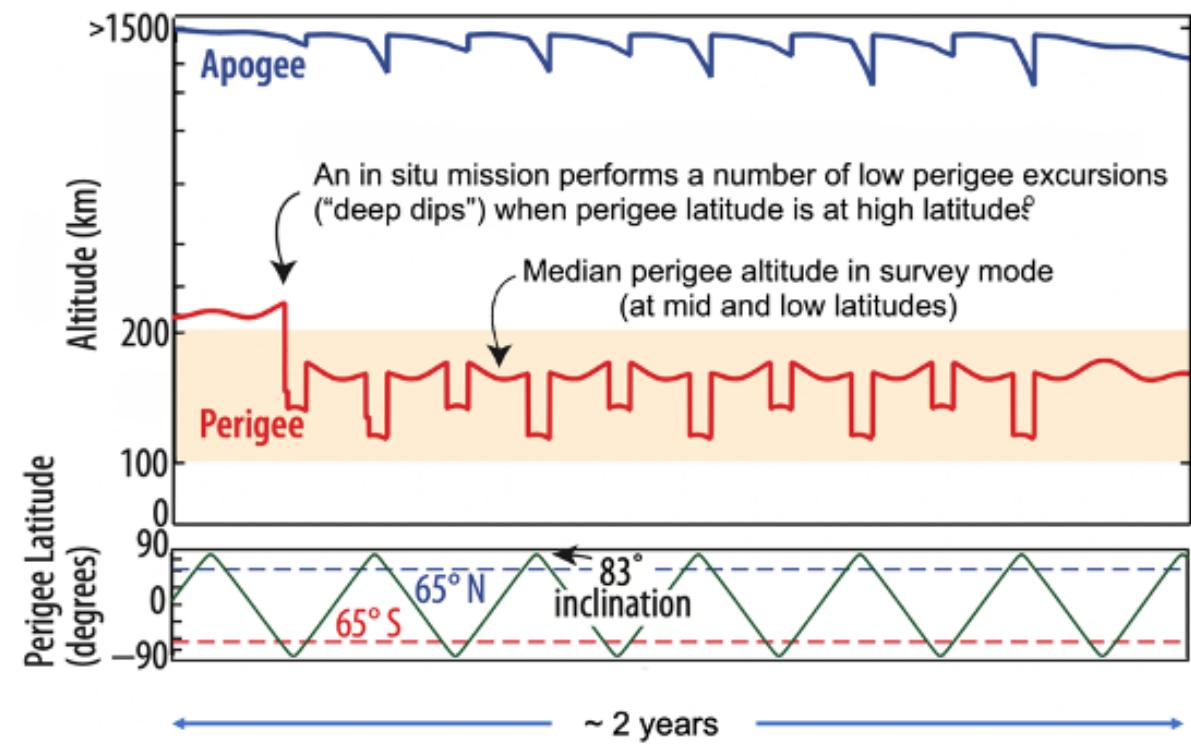
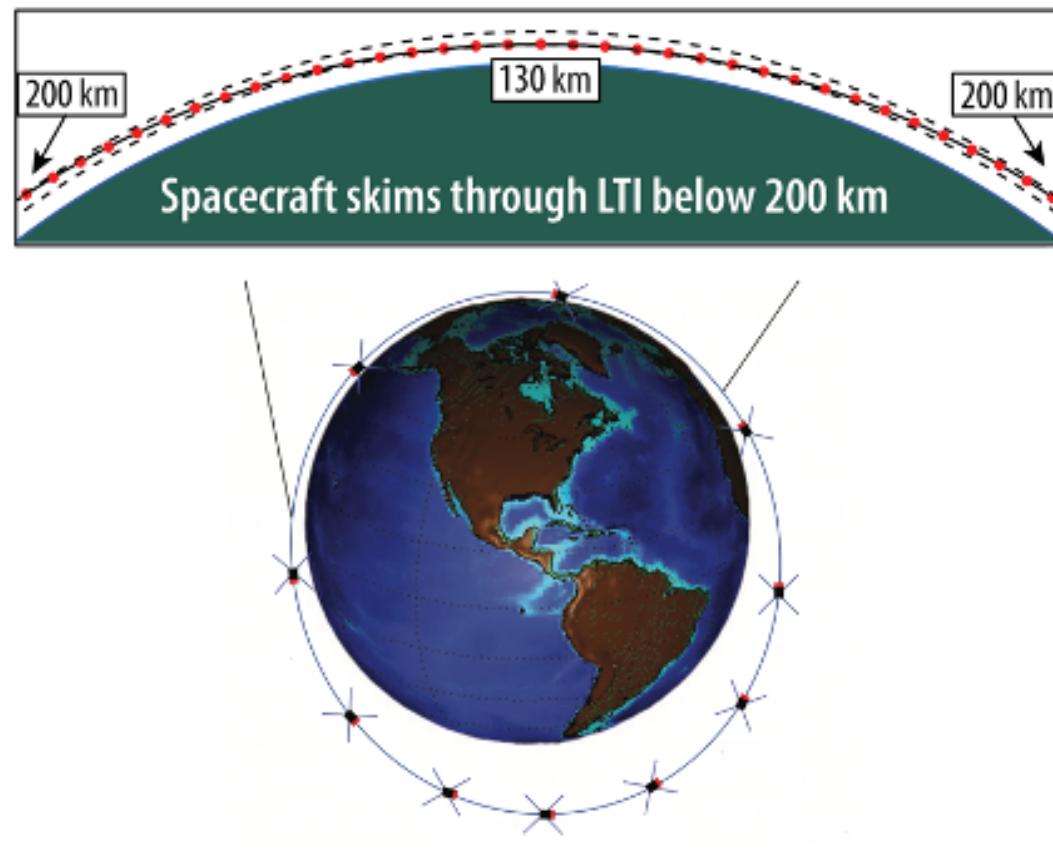
Co-temporal, co-spatial measurements of all parameters involved are needed, in the regions of interest (high and low latitudes, altitudes within 100-200 km)

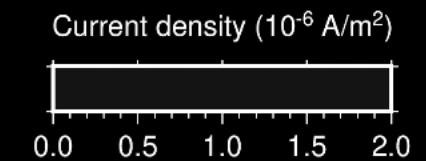
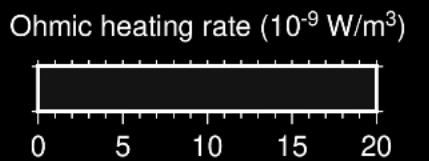
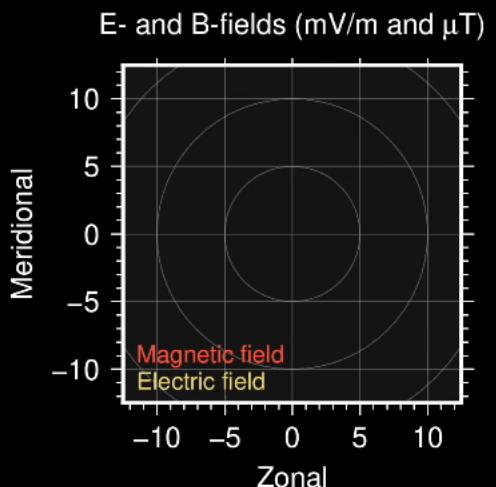
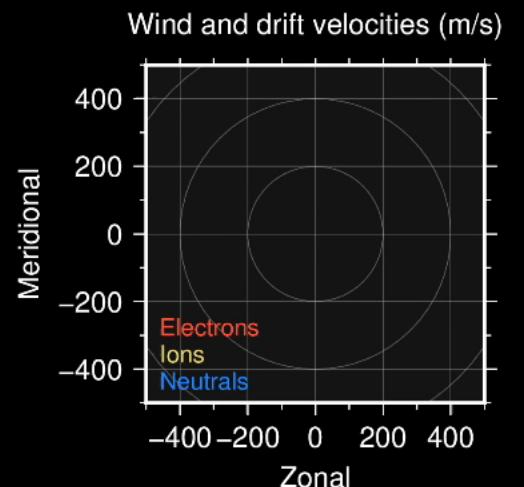
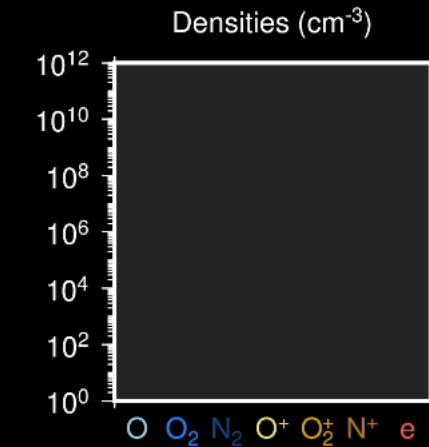
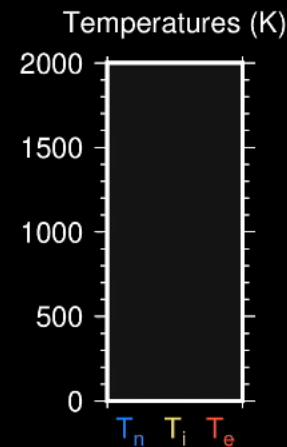
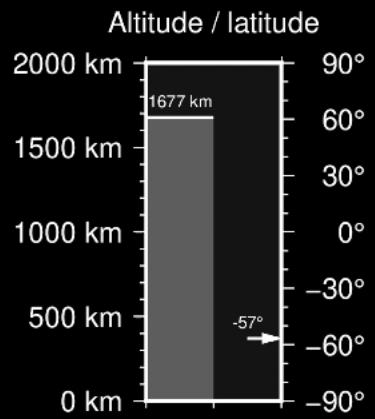
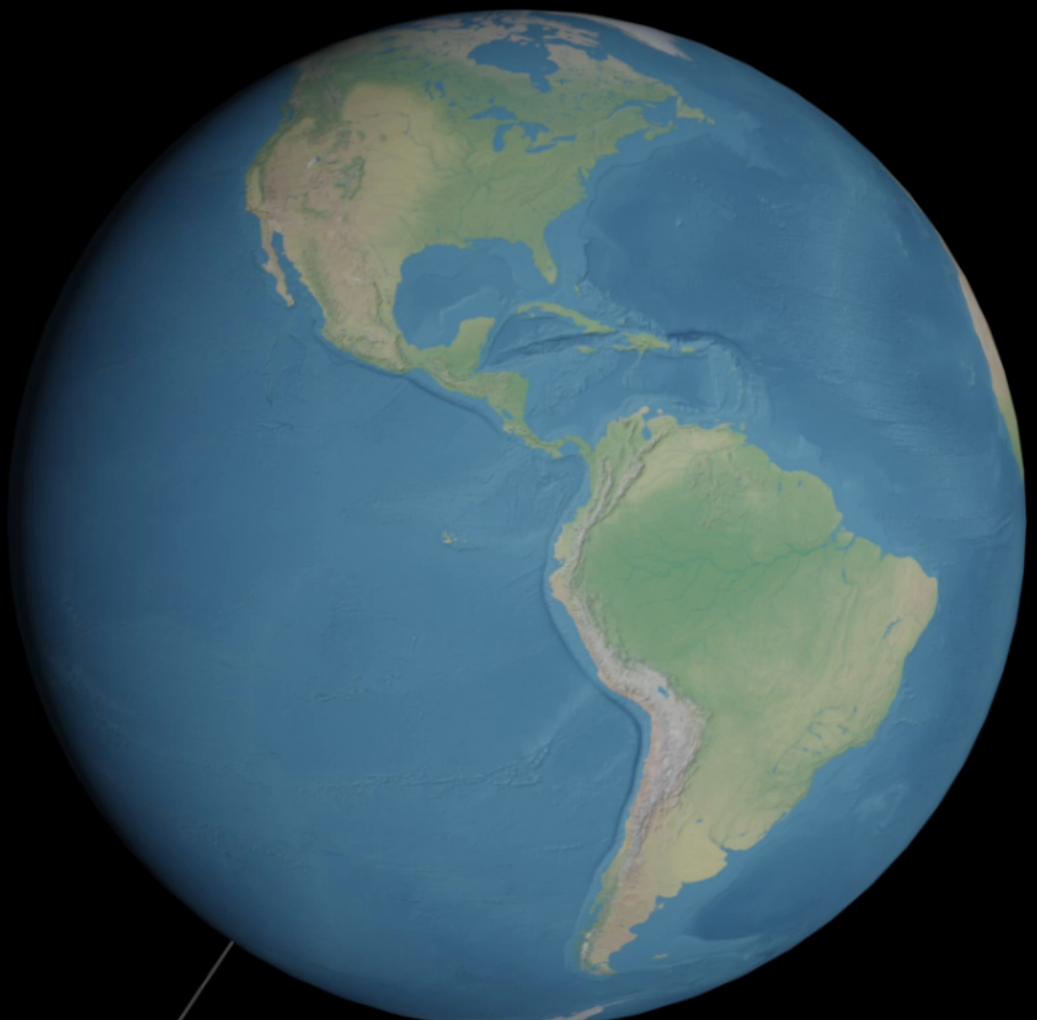
→ *In situ mission with comprehensive instrumentation*

	Abbreviation	Geophysical Observable
ionosphere	\vec{v}_i	Ion Drift velocity
	T_i	Ion Temperature
	T_e	Electron Temperature
	N_i	Ion Number Density
	N_e	Electron Number Density
	n_{ix}	Ion Composition
thermosphere	\vec{u}_n	Neutral Wind Velocity
	N_n	Neutral Number Density
	ρ	Neutral Mass Density
	a_{ng}	Non-gravitational Acceler.
	T_n	Neutral Temperature
fields	\vec{B}	Magnetic Field
	\vec{E}	Electric Field
	$F_{le}, F_{he}, F_{li}, F_{le}$	Energetic Precipitating Particles (ions, electrons)

SO	Derived product / reference	Theoretical equation for estimation
SO1: Collisional Electrodynamics	<ul style="list-style-type: none"> Perpendicular current (via \mathbf{u}_i, \mathbf{u}_e) (Richmond & Thayer 2013, eq. 10) Perpendicular, Pedersen, and Hall currents (Richmond and Thayer 2013, eq. 13) Field-aligned currents (Ampere's Law, e.g., Lühr et al., 2019, eq. 6.3) Convective plasma drift 	<ul style="list-style-type: none"> $\vec{j}_\perp = eN_e(\vec{v}_{i,\perp} - \vec{E} \times \vec{B}/B^2)$ $\vec{j}_\perp = \vec{j}_P + \vec{j}_H = \sigma_p(\vec{E} + \vec{u}_n \times \vec{B}) + \sigma_H \hat{b} \times (\vec{E} + \vec{u}_n \times \vec{B})$ $\vec{j} = (\vec{\nabla} \times \vec{B})/\mu_0 ; j_{ } = (dB_y/v_{sc,x}dt)/\mu_0$ $\vec{v} = \vec{E} \times \vec{B}/B^2$
	<ul style="list-style-type: none"> Pedersen conductivity (Schunk and Nagy 2009, eq. 5.117) Hall conductivity (Schunk and Nagy 2009, eq. 5.118) Parallel conductivity (Schunk and Nagy 2009, eq. 5.125a) 	<ul style="list-style-type: none"> $\sigma_p = \sum_i \frac{N_i e^2}{m_i v_{in} v_{in}^2 + \Omega_i^2} \frac{v_{in}^2}{m_e v_{en} v_{en}^2 + \Omega_e^2}$ $\sigma_H = - \sum_i \frac{N_i e^2}{m_i v_{in} v_{in}^2 + \Omega_i^2} \frac{v_{in} \Omega_i}{m_e v_{en} v_{en}^2 + \Omega_e^2}$ $\sigma_{ } = \frac{N_e e^2}{m_e (v_{en} + v_{ei})}$
	<ul style="list-style-type: none"> Poynting Flux (e.g., Richmond and Thayer 2013, eq. 23) Joule heating (Strangeway 2012, eq. 38) Ohmic heating (Lu et al. 1995, eq. 3) Frictional heating (Strangeway 2012, eq. 28) Energetic precipitating particle-associated heating 	<ul style="list-style-type: none"> $\vec{S} = (\vec{E} \times \delta \vec{B})/\mu_0$ $q_j = eN_e(\vec{v}_{i,\perp} - \vec{u}_{n,\perp}) \cdot (\vec{E} + \vec{u}_n \times \vec{B})$ $q_\Omega = \sigma_p \vec{E} + \vec{u}_n \times \vec{B} ^2$ $q_f = m_i v_{in} N_e$ $q_{EPP} = \sum_{n,s} \sigma_n(E) N_n F_{EPP,s}(E)$
	<ul style="list-style-type: none"> Magnetic forcing (Richmond and Thayer 2013, eq. 20) Gravity Wave Forcing - Momentum and Heat Flux Vertical Transport 	<ul style="list-style-type: none"> $J \times B$ $F_M = \frac{1}{2} \langle u' w' \rangle$ $F_H = \frac{1}{2} \langle w' T' \rangle$

Skeleton Mission Concept



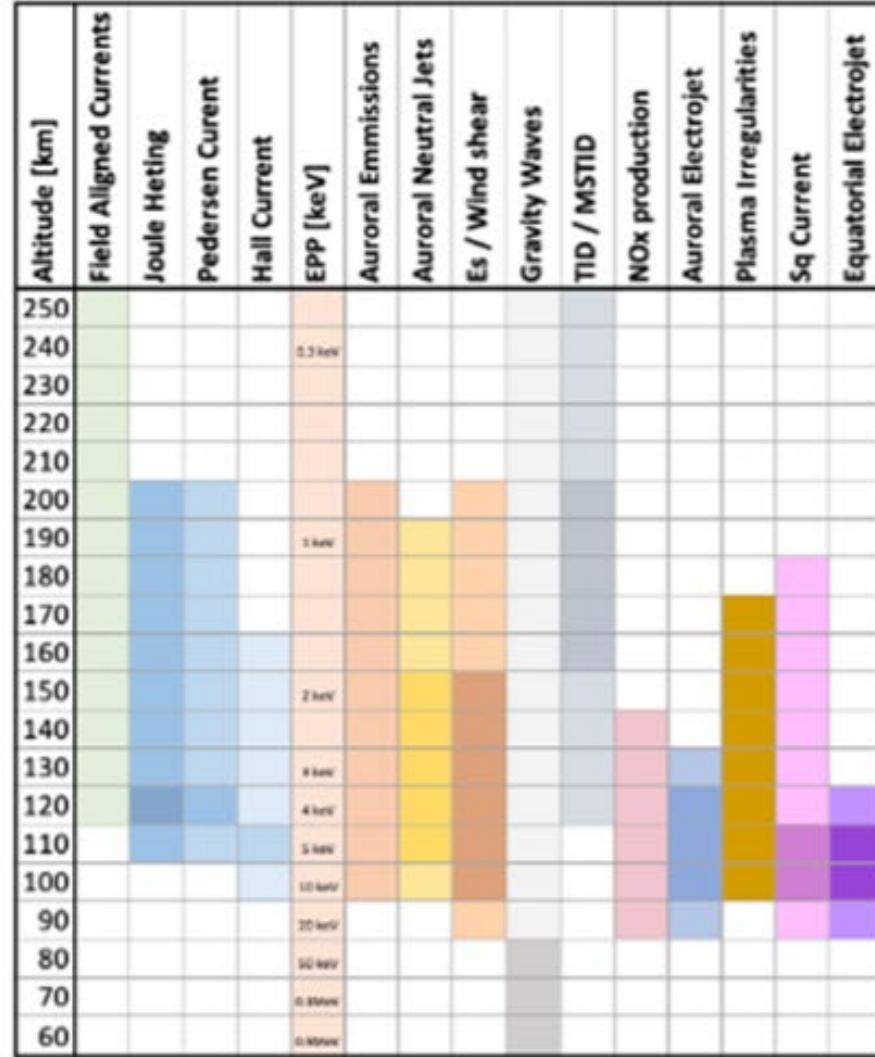


Science Return as a Function of Altitude

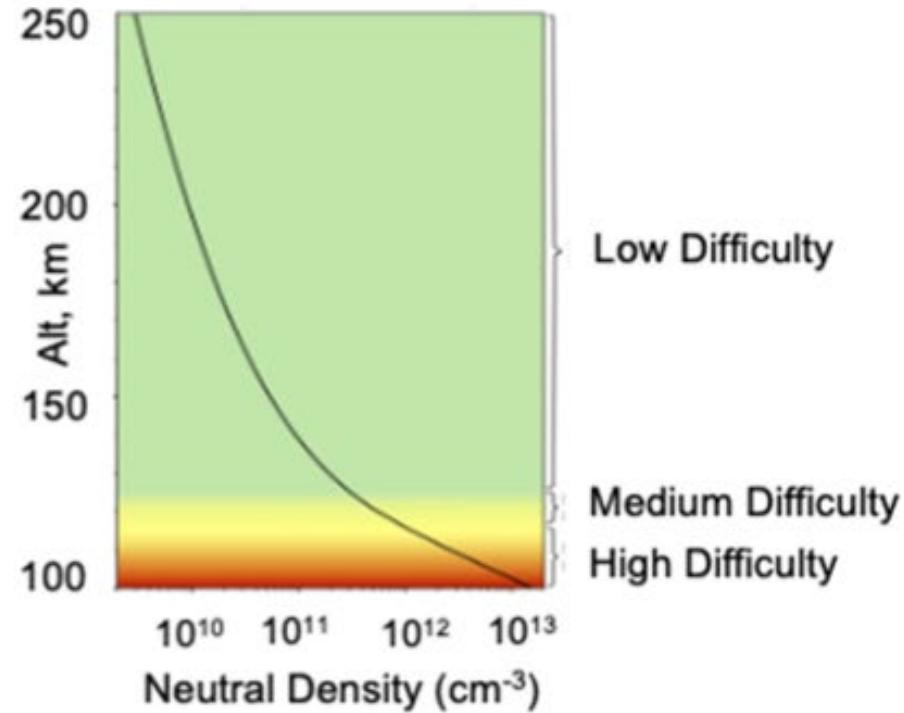


Key LTI Phenomena

Altitude, km



Atmosphere Neutral Density



Likely Instrument Performance on Low Perigee Satellite

(Orbital Velocity ~ 8 km/s)

Geophysical Observable

Altitude	N_i	V_i	T_i	N_{ix}	N_e	T_e	N_n	U_n	T_n	N_{nx}	ρ_n	E	B
>200 km	green	green	green	green	green	green	green	green	green	green	green	green	green
150 km	green	green	green	green	green	green	green	green	green	green	green	green	green
130 km	green	green	green	green	green	green	green	green	green	green	green	green	green
125 km	green	green	green	green	green	green	green	green	green	green	green	green	green
120 km	green	green	green	green	green	green	green	green	green	green	green	green	green
115 km	yellow	yellow	yellow	yellow	yellow	yellow	yellow	yellow	yellow	yellow	yellow	yellow	green
110 km	yellow	yellow	yellow	yellow	yellow	yellow	yellow	yellow	yellow	yellow	yellow	yellow	green
105 km	yellow	yellow	yellow	yellow	yellow	yellow	yellow	yellow	yellow	yellow	yellow	yellow	green
100 km	red	red	red	red	red	red	red	red	red	red	red	red	green

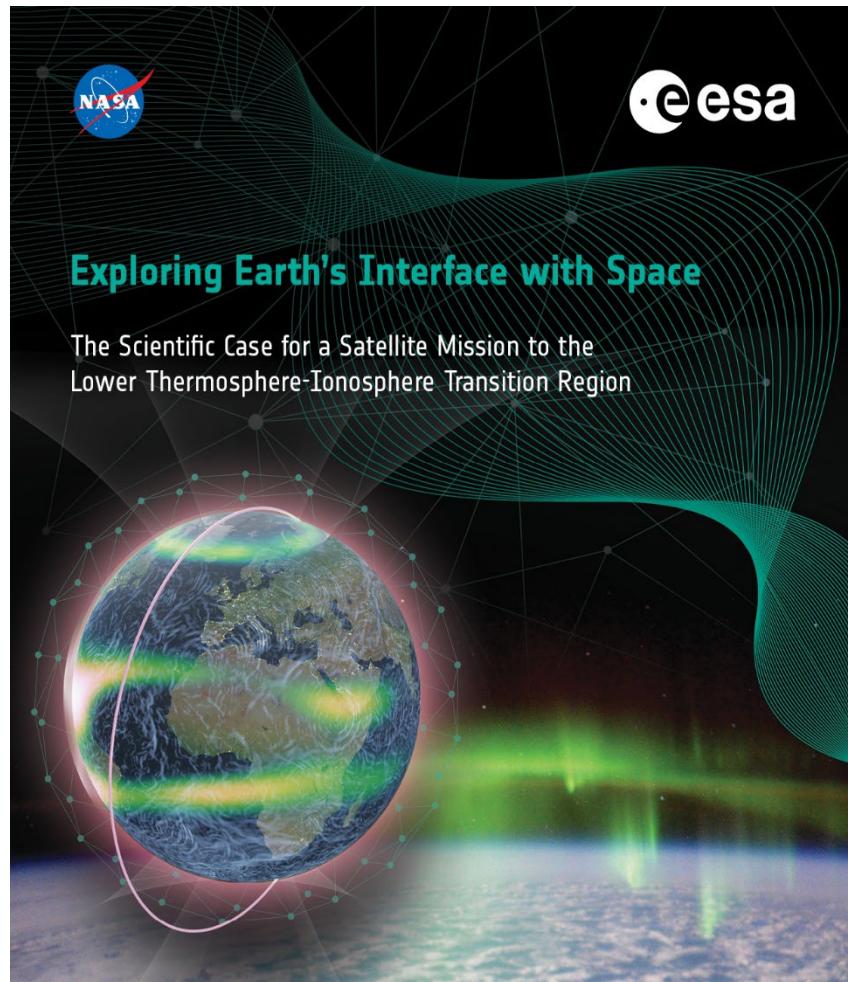
ENLoTIS Report: Summary and Next Steps



- ❑ In situ measurements in the 100-200 km altitude range with a comprehensive instrument suite will provide, for the first time, all geophysical observables needed for the unambiguous quantification of key processes in the LTI
- ❑ These processes need to be quantified in-situ, with statistically representative sampling over the mission lifetime.

Next steps:

1. Flow-down of the scientific needs expressed in this report into firm mission requirements supported by analyses, tools and capabilities to justify and verify these requirements;
2. Assessment of instrumental capabilities needed to meet measurement requirements;
3. Study of orbits, propulsion, as well as mission and science operational concepts capable of meeting sampling requirements;
4. Further study on mitigation approaches to the special environmental challenges listed above that affect the measurement techniques, data processing, spacecraft design choices and concepts of operations.



<http://doi.org/10.5270/ESA-NASA.LTI-SC.2024-07-v1.0>

Questions?

NASA POC: john.p.mccormack@nasa.gov

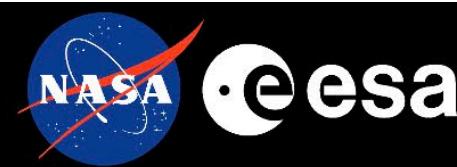
ESA POC: Alex.Hoffmann@esa.int

The WG enables ESA-NASA cooperation on future LTI satellite mission concepts by:

- a) Reviewing and consolidating **consensus science questions or goals, mission objectives, and high-level mission requirements** that would inform the eventual definition and design of (a) future mission concept(s)
 - Not starting from “blank slate” – leverage knowledge from past and current mission studies
 - Input/feedback from research community throughout initial phase will be key
 - From Heliophysics perspective, initial phase of ENLoTIS WG would resemble an “SDT” or Science Definition Team.
- b) Identifying **scientific and technical challenges and constraints** associated with these high-level requirements from (a) in view of facilitating trade-offs and identifying candidate measurements.
 - Balancing science and feasibility – how low should we go vs. how low can we go?
- c) Coordinating with **on-going and planned activities** between NASA & ESA supporting (a) and (b)

Programmatic context (ESA)

The Daedalus concept, an ESA Earth Observation Programme Earth Explorer 10 mission candidate (Phase 0)



- Targets a better understanding of the **atmosphere-space** (thermosphere-ionosphere) **coupling**, to shed light on key ion-neutral interaction processes affecting structure, energetics, composition and dynamics of the upper atmosphere, by
- Exploring the **transition region** (~120 to 200 km altitude) **in situ**, using a deep diving spacecraft.

