lo science

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Jet Propulsion Laboratory California Institute of Technology

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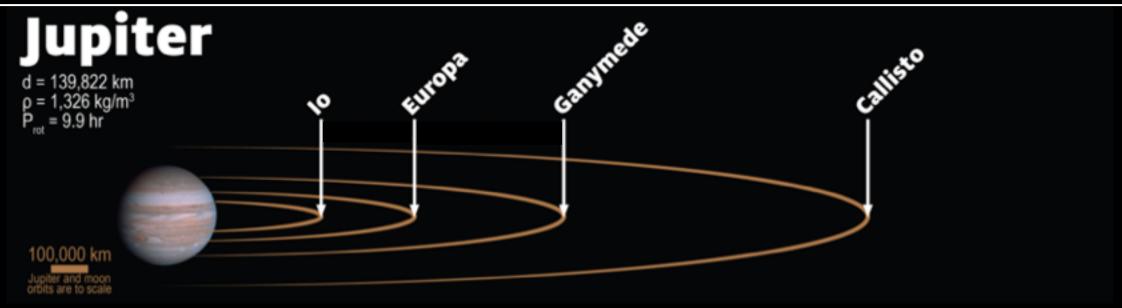
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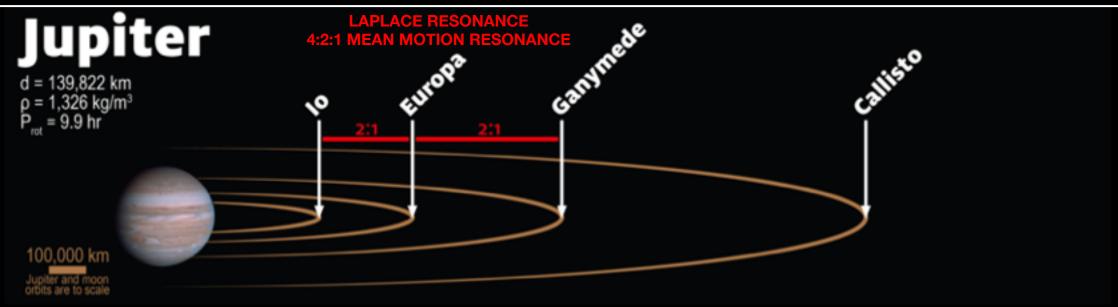
SPRING MEETING OF THE COMMITTEE ON ASTROBIOLOGY AND PLANETARY SCIENCE (CAPS)

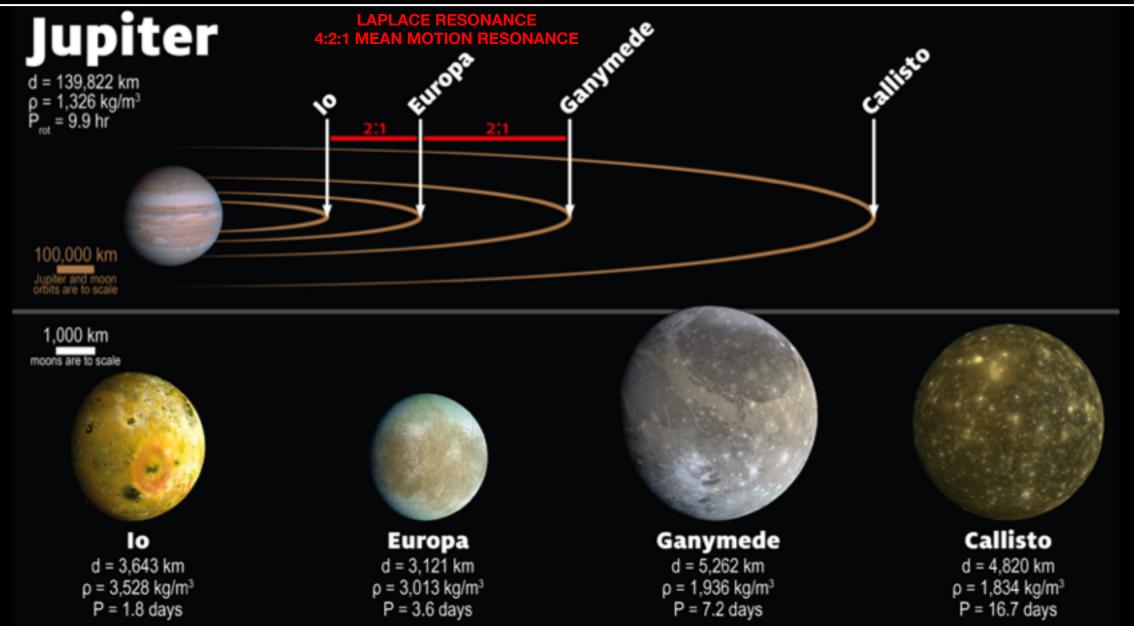
NATIONAL ACADEMIES OF SCIENCES, ENGINEERING, AND MEDICINE SPACE STUDIES BOARD

OUTLINE

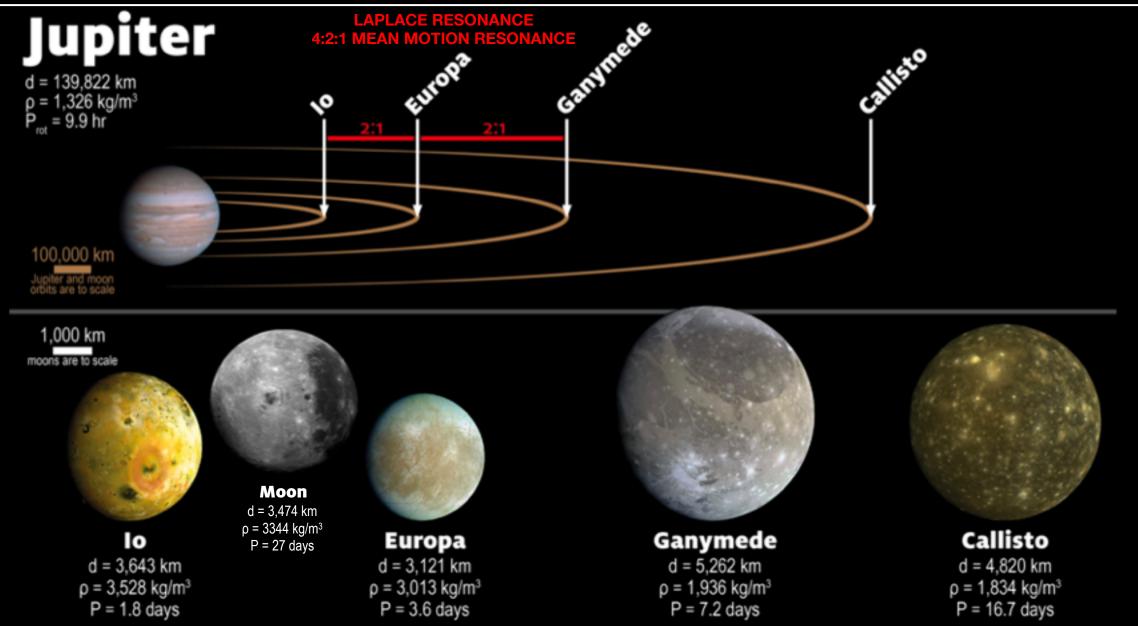
- A primer on Io.
- The science rationale for exploring lo.
- What did Vision and Voyages say?
- What's changed since Vision and Voyages?
- Io Volcano Observer.
- Discovery vs. New Frontiers.
- Conclusion.



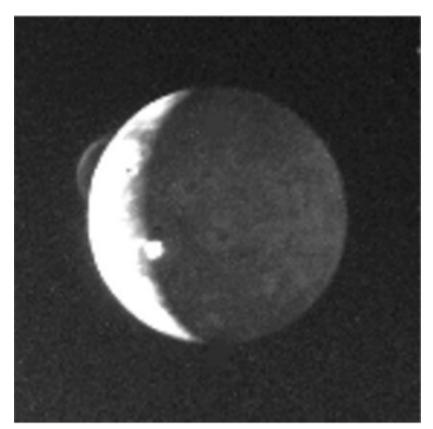




(NASA / JPL-Caltech / Voyager-ISS / Justin Cowart / Kevin M. Gill / Emily Lakdawalla / Jason Perry / Ted Stryk / Gordan Ugarkovic / Keck Institute for Space Studies / James T. Keane)



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↑ Voyager 1 discovery image if active volcanism and plumes on Io (NASA/JPL)

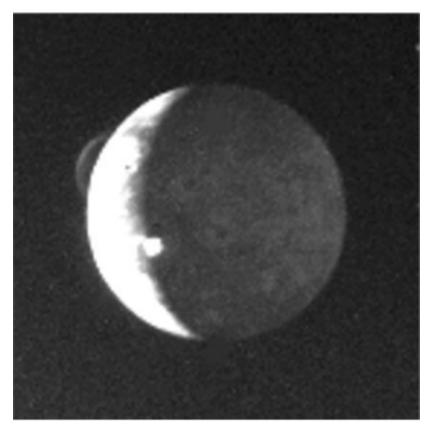


Peale et al.

"Melting of Io by tidal dissipation"
Science 203, 892-894 (2 Mar. 1979)

Morabito et al.

"Discovery of currently active extraterrestrial volcanism" Science 204, 972 (1 Jun. 1979)



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Witteborn et al.

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Although it is not possible to definitively rule out this explanation, we consider it to be unlikely by analogy with our experience with objects in the inner solar system. No such internally generated hot spot having the approximate size and temperature of the one postulated for Io has ever been observed on global infrared surveys of Earth, the moon, Mercury, or Mars (11). ↑ Voyager 1 discovery image if active

Witteborn et al.

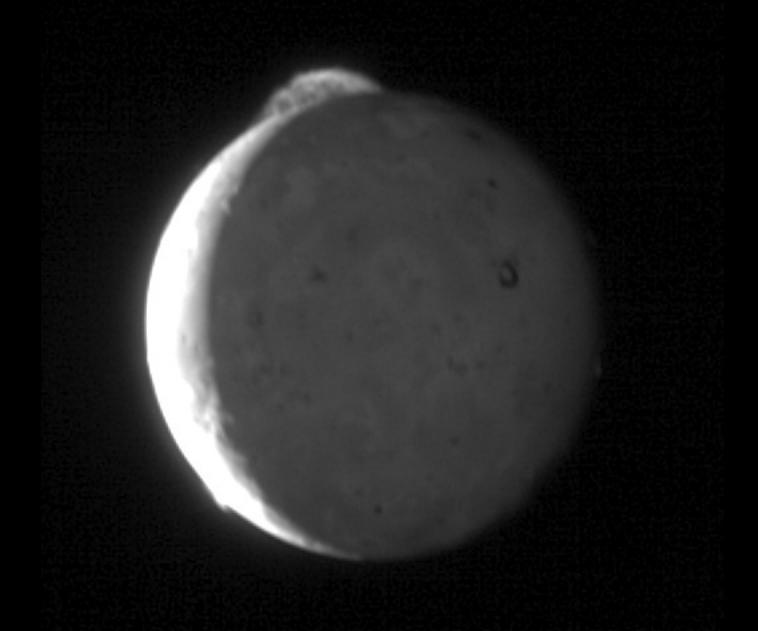
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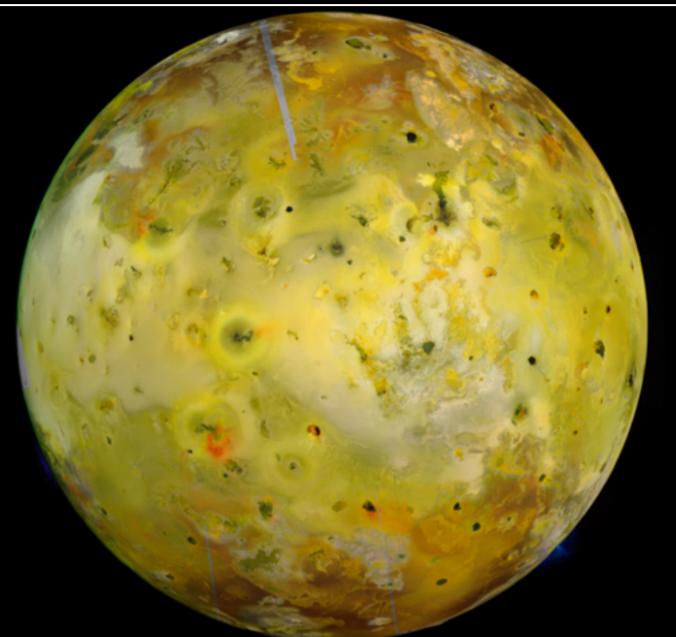
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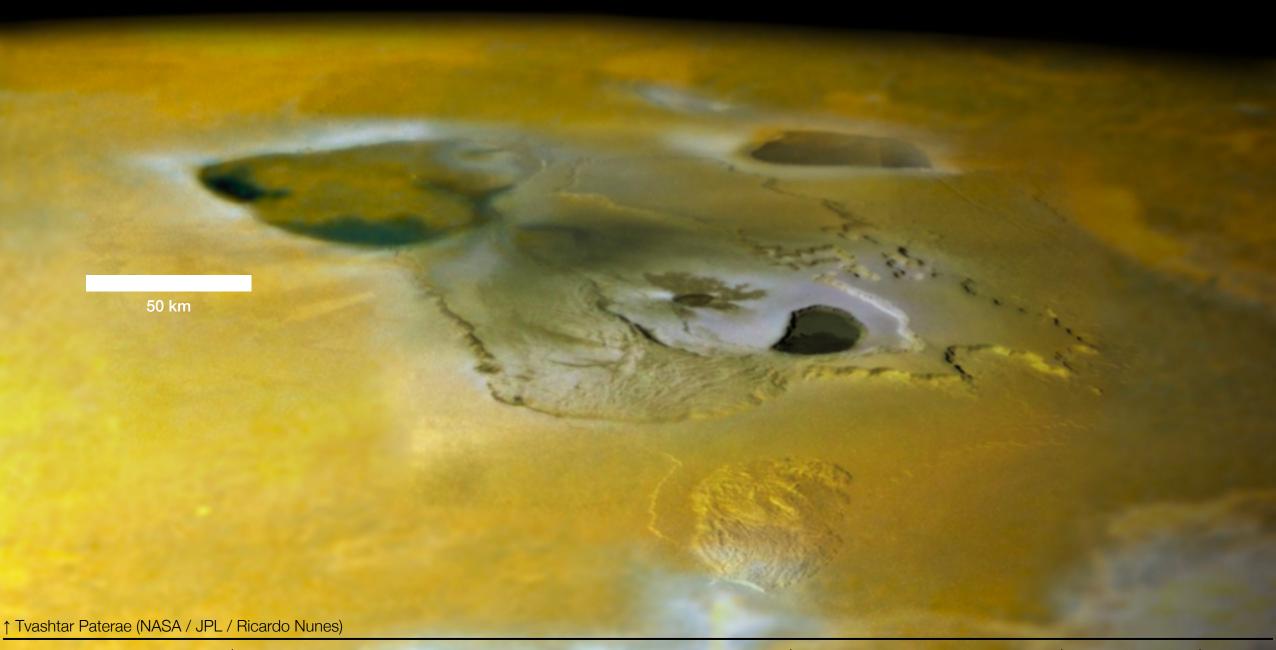
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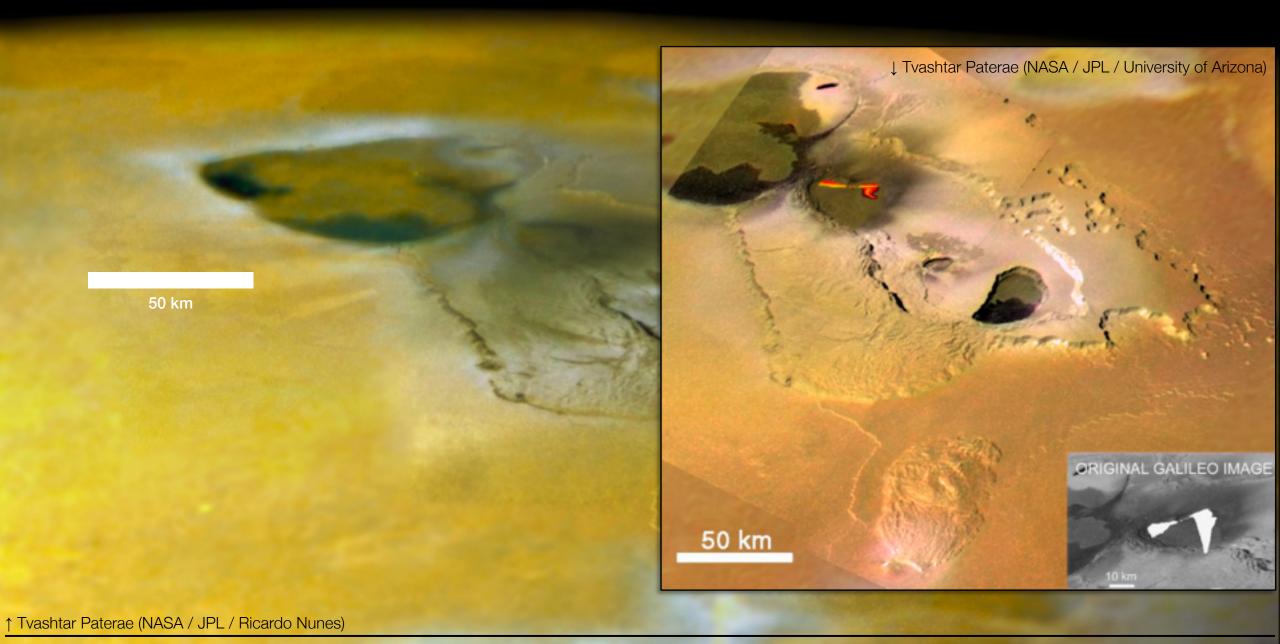


← New Horizons movie of the Tvashtar plume, during its 2007Jupiter flyby (NASA / JHUAPL / SwRI)

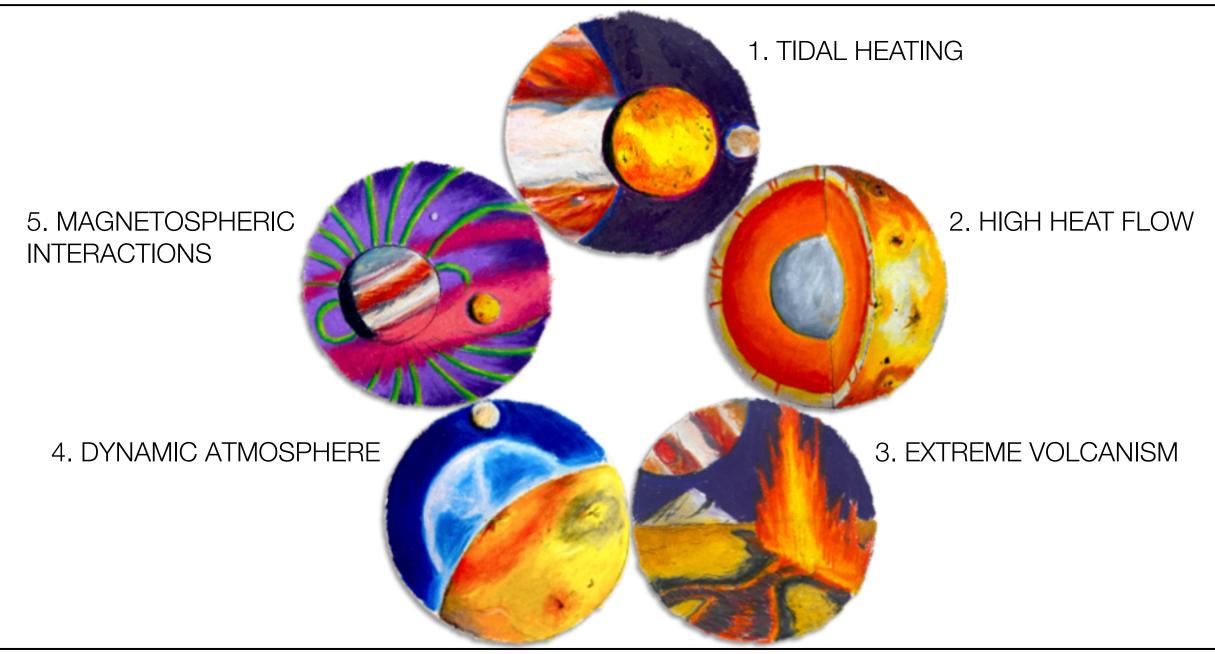


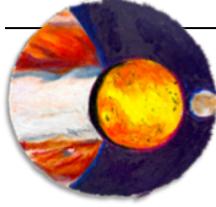
← *Galileo* image mosaic of lo's anti-Jovian hemisphere (NASA / JPL / Jason Perry)





CROSS-CUTTING THEMES OF IO EXPLORATION

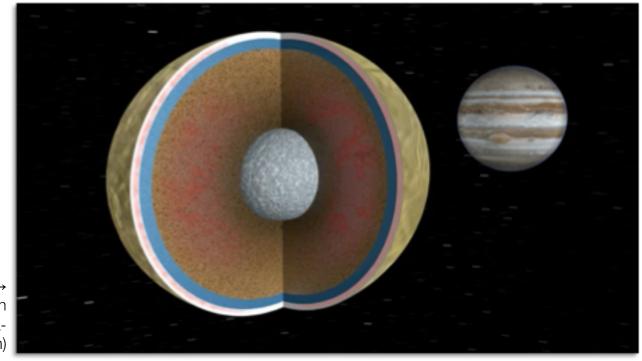




CROSS-CUTTING THEME #1:

TIDAL HEATING

Io is the best place in the solar system to study tidal heating—a critical process for shaping planetary bodies and creating habitable environments across the cosmos.



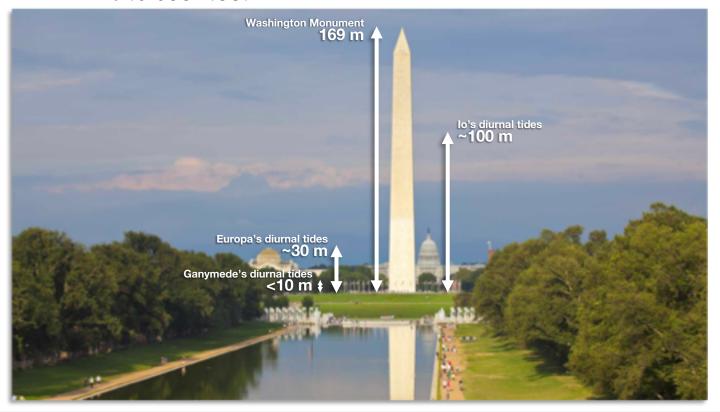
Tidal flexing animation of Europa (NASA / JPL-Caltech)

- What is the magnitude and spatial distribution of Io's heat flow?
- What is Io's interior structure and does it have a magma ocean?
- Where and how is tidal heat dissipated in the interior?
- How is Io being tidally deformed?
- What are the present-day orbital migration rates of the Galilean satellites, and is the system in equilibrium?
- How does tidal heating shape crustal structure?

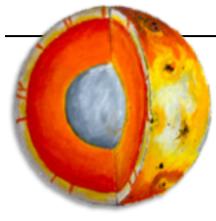


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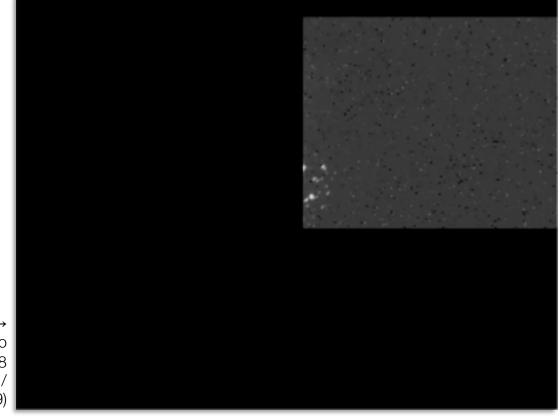
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CROSS-CUTTING THEME #2:

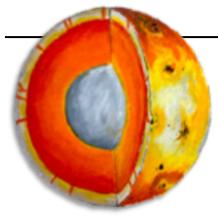
HIGH HEAT FLOW

Io is a living example of how rocky planets respond to extreme heating, akin to the Earth when life first emerged.



Juno JIRAM observations of lo entering eclipse, December 2018 (NASA / JPL-Caltech / SwRI / ASI / INAF / JIRAM / J. Roger / @landru79)

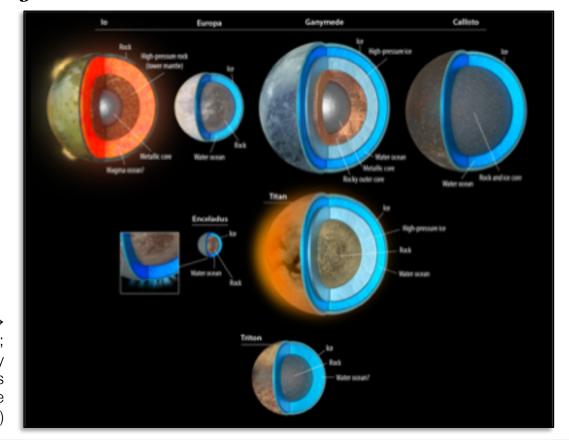
- How does Io lose its internal heat, and what is the balance between heat loss mechanisms (conduction, heat-pipe volcanism, etc.)?
- How do volcanic eruptions and their characteristics relate to the deep interior?
- What is the thickness, composition, and stress-state of Io's lithosphere?
- Does Io have an internally-driven magnetic field?
- Can stable isotopes inform longterm evolution of Io?



CROSS-CUTTING THEME #2:

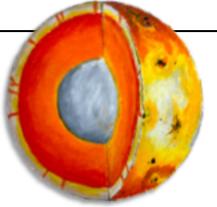
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The solar system's ocean worlds; scale of interior layers are only approximate (Chuck Carter / James T. Keane / Keck Institute for Space Studies)

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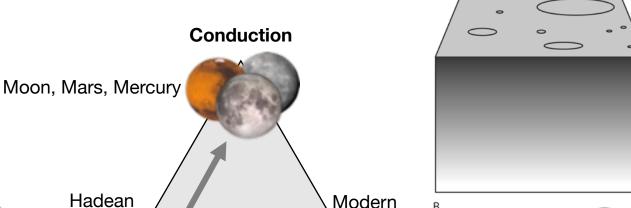


CROSS-CUTTING THEME #2:

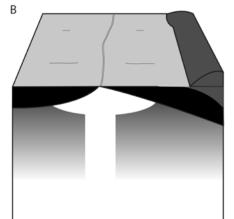
Earth

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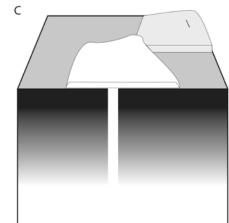


Earth



PRIORITY SCIENCE QUESTIONS:

- How does Io lose its internal heat, and what is the balance between heat loss mechanisms (conduction, heat-pipe volcanism, tectonics)?
- How do volcanic eruptions and their characteristics relate to the deep interior?
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Heat Pipe Plate Volcanism Tectonics

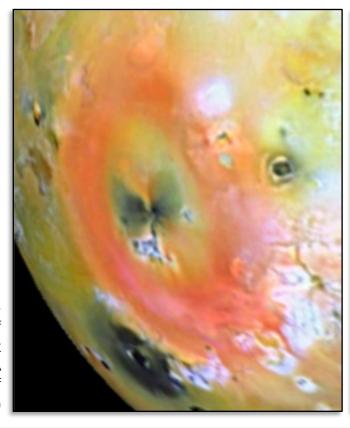
End-members for planetary heat loss (Lauren Jozwiak, James T. Keane)

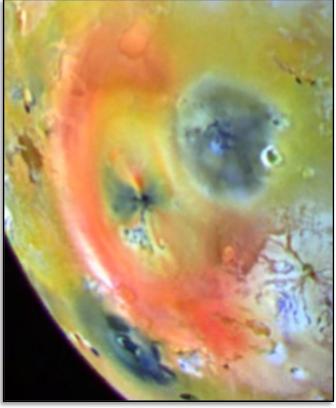


CROSS-CUTTING THEME #3:

EXTREME VOLCANISM

Io is the most volcanically active world in the solar system, and it is a window into volcanic processes both familiar and exotic.





PRIORITY SCIENCE QUESTIONS:

- What is the composition and chemistry of Io's magmas?
- How does volcanism operate in extreme environments?
- How do Io's large volcanic edifices (paterae) form?
- What processes create and destroy Io's mountains?
- How do tectonism and magmatism interact on both global and local scales?
- How do Io's plumes operate, and what can they tell us about Io's interior processes?

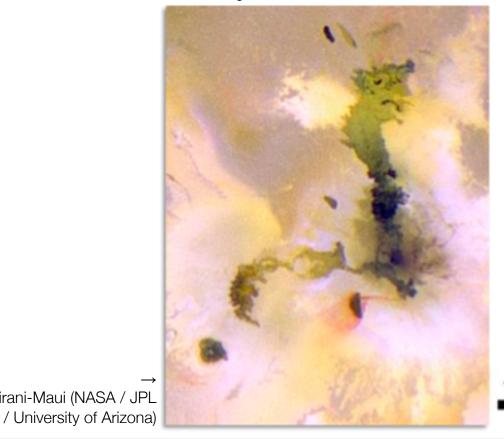
Galileo images of Pele and Pillan, six months apart (NASA / JPL / University of Arizona)



CROSS-CUTTING THEME #3:

TREME VOLCANISM

Io is the most volcanically active world in the solar system, and it is a window into volcanic processes both familiar and exotic.





Amirani-Maui (NASA / JPL

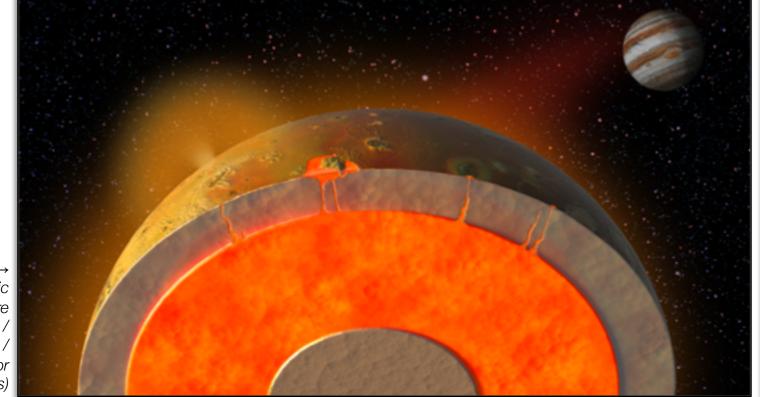
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CROSS-CUTTING THEME #4:

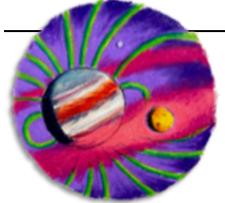
DYNAMIC ATMOSPHERE

Io's atmosphere is unlike any other, with a rich array of interactions that test our understanding of fundamental planetary science.



lo's dynamic atmosphere (Chuck Carter / James T. Keane / Keck Institute for Space Studies)

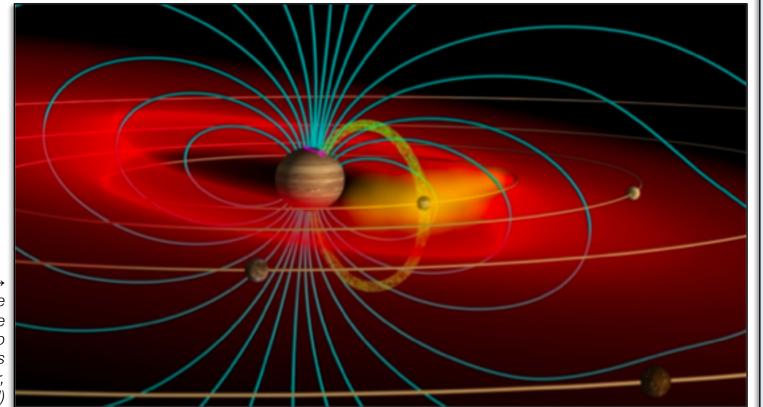
- What are the dominant sources and sinks of Io's atmosphere?
- What are the dynamics of Io's atmosphere and plumes?
- What physical and chemical processes are responsible for Io's complex surface color patterns?
- What was Io's original volatile inventory, and how were those volatiles lost?
- What is the sulfur speciation in *Io's volcanic plumes and* atmosphere, and what does it tells about the chemical state of Io's deep interior?



CROSS-CUTTING THEME #5:

MAGNETOSPHERES

Io drives the activity in the Jovian magnetosphere, yielding a wonderland for understanding fundamental planetary physics.



→ The magnetosphere of Jupiter and lo

ot Jupiter and io plasma torus (John Spencer, SwRI)

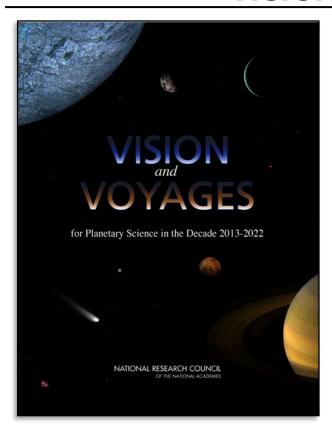
- Is Io's ionosphere global, does it vary, and how is it maintained?
- How do Io's volcanoes affect the supply of plasma, gas, and dust in the Jovian system?
- How much Iogenic material reaches the surface of Europa?
- How do plasma-atmosphere interactions vary with time and location around the moon?
- How does Io interact with the inward motion of plasma that can eventually rain onto Jupiter's ionosphere, generating aurorae?

IO IS A WORLD OF SYNERGIES



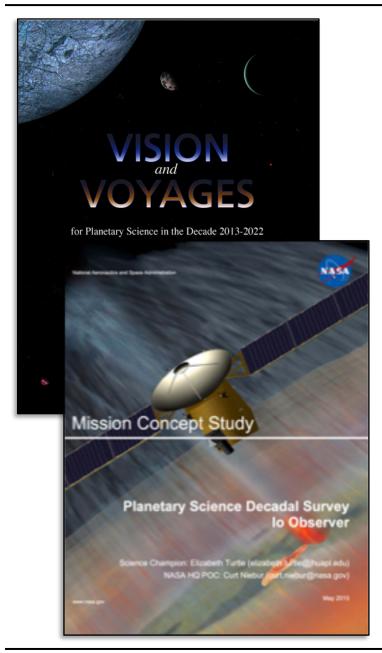
IO IS A WORLD OF SYNERGIES





The science case for lo:

"[Io is] the ideal target to study tidal dissipation and the resulting variety of volcanic and tectonic processes in action, with fundamental implications for the thermal co-evolution of the Io-Europa-Ganymede system as well as for the habitable zone around other stars."



The science case for lo:

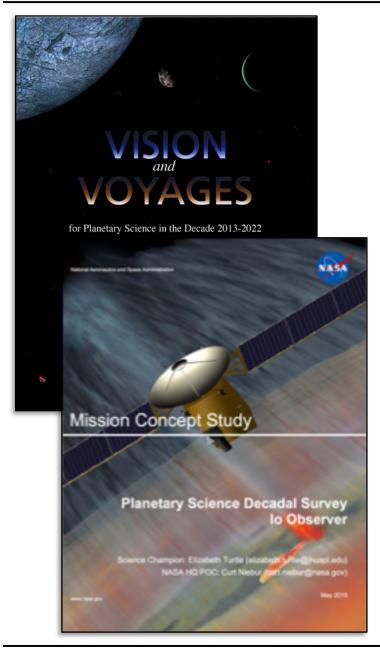
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New Frontiers Io Observer mission concept:

"Io Observer—The focus of this mission is to determine the internal structure of Io and to investigate the mechanisms that contribute to the satellite's intense volcanic activity from a highly elliptical orbit around Jupiter, making multiple flybys of Io."

New Frontiers to Observer science objectives:

- 1. Study Io's active volcanic processes;
- 2. Determine the melt fraction of Io's mantle;
- 3. Constrain tidal heating mechanisms;
- 4. Study tectonic processes;
- 5. Investigate interrelated volcanic, atmospheric, plasma-torus, and magnetospheric mass- and energy-exchange processes;
- 6. Constrain the state of Io's core via improved constraints on whether Io generates a magnetic field; and
- 7. Investigate endogenic and exogenic processes controlling surface composition.



New Frontiers 4 options:

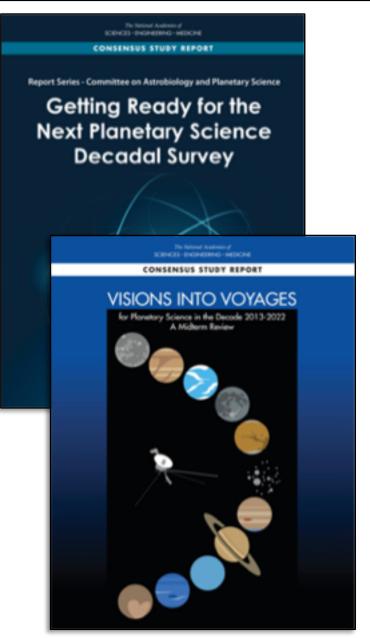
- Comet Surface Sample Return
- Lunar South Pole–Aitken Basin Sample Return
- Saturn Probe
- Trojan Tour and Rendezvous
- Venus In Situ Explorer
- Ocean Worlds (Enceladus/Titan)

Additional New Frontiers 5 options:

- Io Observer
- Lunar Geophysical Network







Getting ready for the next decadal:

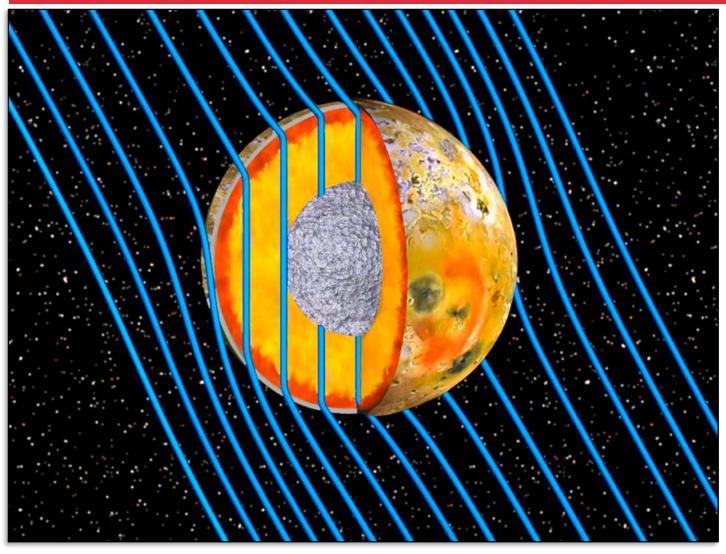
"Given that the announcement of opportunity for New Frontiers 5 will likely come in the early 2020s, the revalidation of the Io Explorer as having appropriate scientific merit for inclusion in the New Frontiers program will fall to the next planetary science decadal survey [...]. As such, the technical feasibility of the Io Explorer concept [...] could be fruitfully reexamined."

Midterm review:

"...the committee believes [Io Observer and Lunar Geophysical Network] still remain valid missions for New Frontiers 5."



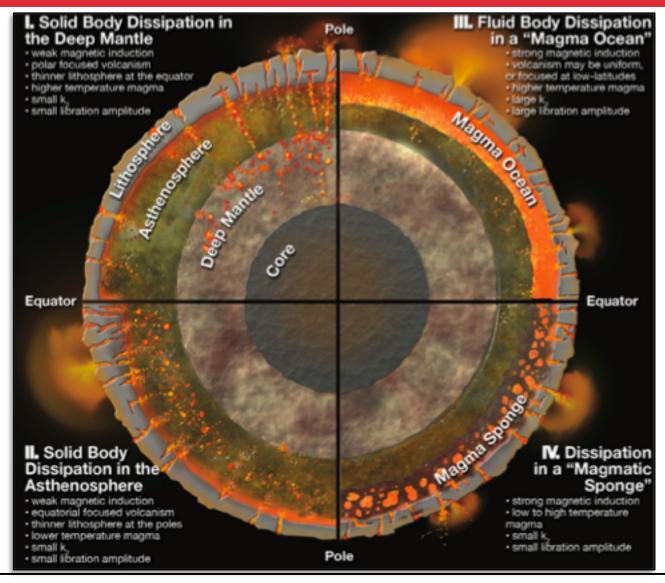
Example 1: EVIDENCE FOR A MAGMA OCEAN



Reanalysis of Galileo magnetometer data suggests Io may have a global magma ocean (Khurana et al. 2011 Šebek et al. 2019), although this interpretation is debated (Roth et al. 2017, Blöcker et al. 2018).

← lo's induced magnetic field (Xianzhe Jia / Krishan Khurana)

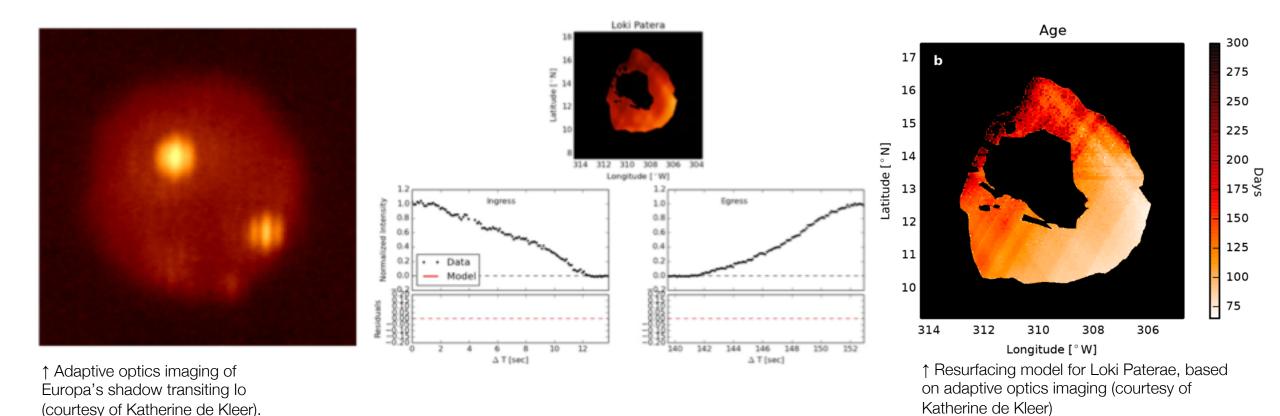
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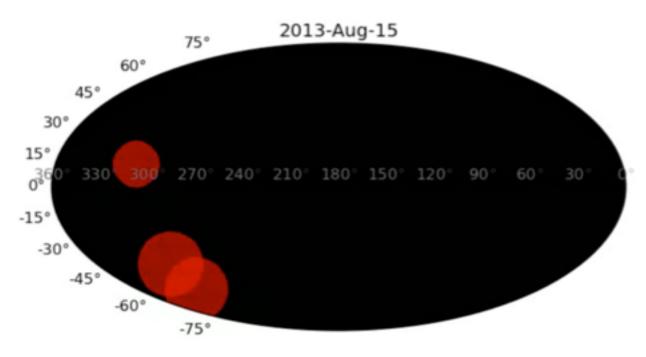
[←] Four competing models for lo's interior structure (Chuck Carter / James T. Keane / Keck Institute for Space Studies)

Example 2: ADAPTIVE OPTICS IMAGING



Adaptive optics imaging has provided extremely high-resolution ground-based imaging of individual volcanoes, enabling unprecedented investigation of certain aspects of Io's volcanism (e.g., de Kleer et al. 2017).

Example 3: EVIDENCE FOR A MAGMA OCEAN



↑ Five years of lo observations (courtesy of Katherine de Kleer).

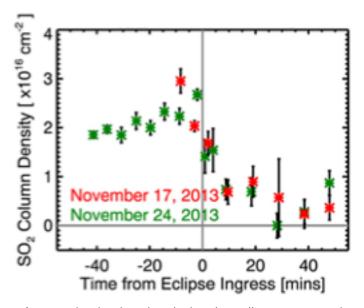
High-cadence monitoring of Io's volcanic activity from ground-based telescopes have revealed new trends in Io's volcanism, identified new volcanic outbursts, and hint at the geophysical processes modulating Io's activity (de Pater et al. 2016, Cantrall et al. 2018, de Kleer et al. 2019a-b).

Example 4: ATMOSPHERIC COLLAPSE



↑ Keck AO observations of lo coming out of eclipse (courtesy of Katherine de Kleer)

New observations from ground-based telescopes have been able to characterize some of the dynamics of Io's atmosphere—like its partial collapse during eclipse (Tsang et al. 2016).



† Atmospheric density during lo eclipse, as monitored by TEXES on *Gemini-north* (Tsang et al. 2016)

Example 5: ISOTOPES



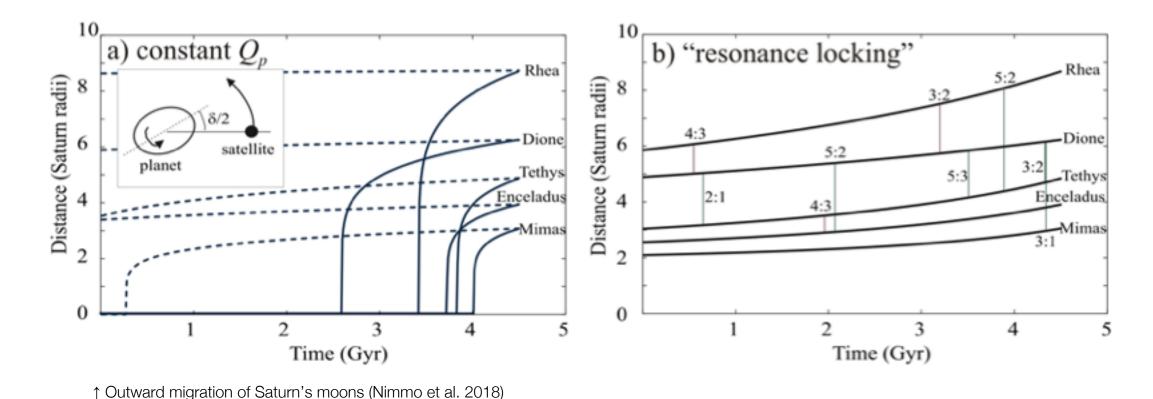
Isotopic ratios have been transformative for understanding the evolutionary histories of many bodies, from Titan to Pluto (e.g., Liang et a. 2007, Mandt et al. 2009, Mandt et al. 2012a-b, Nixon et al. 2012), Mandt et al. 2014, Mandt et al. 2017).

New millimeter and radio observatories have enabled the first measurements of isotopic ratios in Io's atmosphere (34S/32S, Moullet et al. 2013). Stable isotopes may provide a window into the deep history of Io, a world where geologic record is otherwise constantly obliterated.

← Atacama Pathfinder Experiment (APEX) telescope

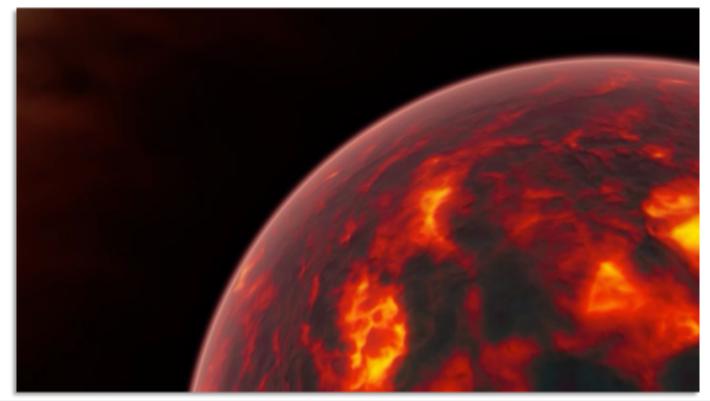
Example 6: ASTROMETRY

Astrometric observations of Saturn's moons by Cassini (Lainey et al. 2012, 2017) have revealed that Saturn's moons are migrating much faster than expected. This has motivated a new generation of orbital evolution models (Fuller et al. 2016), which are still in their infancy.



Example 8: THE EXOPLANET RENNAISSANCE

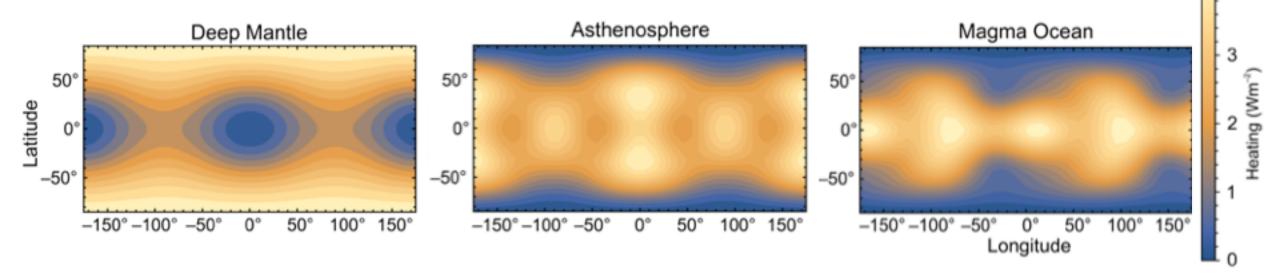
Io is an important analog for extremely heated exoplanets—including both tidally heated worlds (e.g., the resonant super-Ios in TRAPPIST-1, Barr et al. 2018), and worlds heated by extreme solar insolation (e.g., the close-in "lava worlds" like 55 Cnc e, Demory et al. 2016). Future telescopes may be able to characterize these worlds.



← Artist impression of 55 Cnc e (ESA / *Hubble* / M. Kornmesser)

Example 9: THEORETICAL MODELS

Theoretical models for tidal heating have substantially improved, enabling exploration of more realistic and expansive parameter spaces, including dissipation in both solid and fluid layers (Beuthe 2013, Tyler et al. 2015, Matsuyama et al. 2018, Hay & Matsuyama 2019), modeling melt migration (Schools et al. 2019), accounting for mantle convection (Kiefer 2019), and modeling feedbacks between tidal heating and rheology (Steinke et al. 2019).



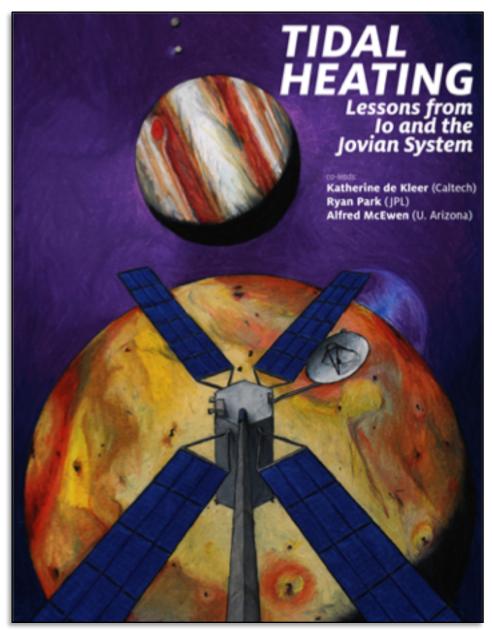
† Expected patterns of surface heat flux on lo for three end-member cases (Courtesy of Isamu Matsuyama, in de Kleer et al. 2019)



† Participants of the October 2018 KISS workshop (Keck Institute for Space Studies)

Tidal Heating—Lessons from Io and the Jovian System

Caltech Keck Institute for Space Studies, 15–19 October 2018, 3 December 2018



Tidal Heating—Lessons from lo and the Jovian System Coltach Kook Institute for Space Studies

Caltech Keck Institute for Space Studies, 15–19 October 2018, 3 December 2018

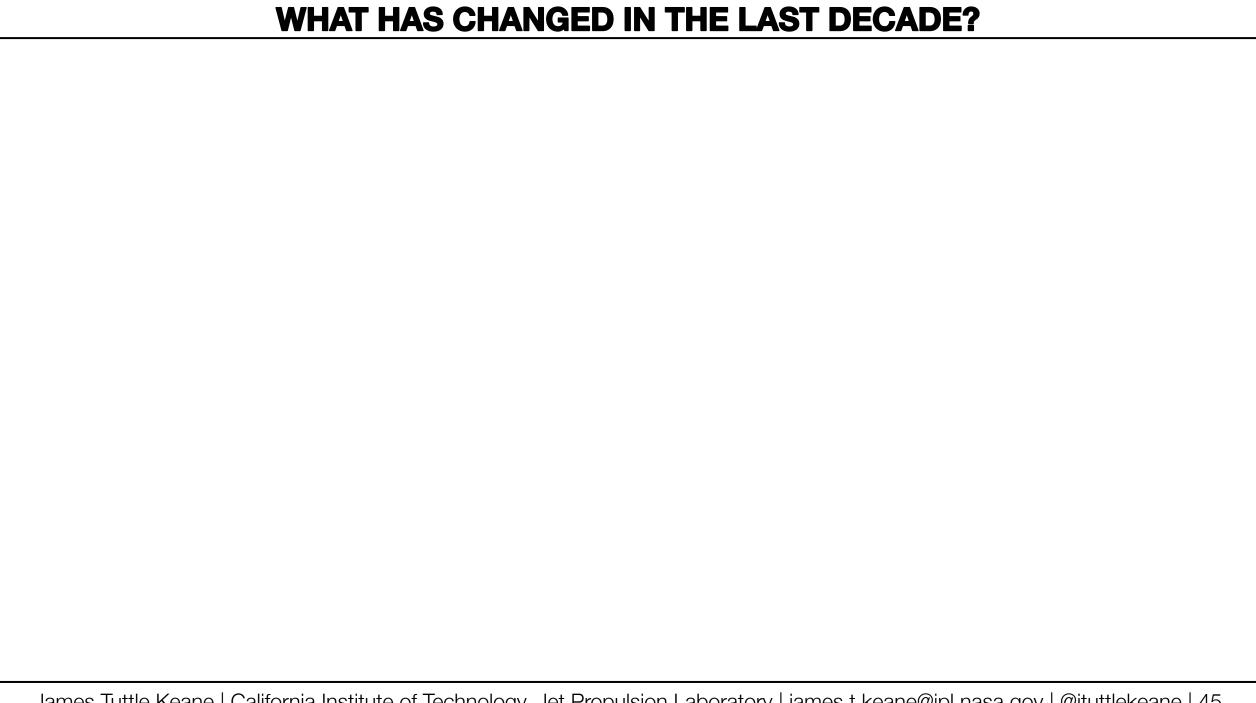
Full, 99-page report:

https://www.kiss.caltech.edu/programs.html#tidal_heating

Five key questions:

- 1. What do volcanic eruptions tell us about the interiors of tidally heated bodies?
- 2. How is tidal dissipation partitioned between solid and liquid materials?
- 3. Does Io have a magma ocean?
- 4. Is the Jupiter/Laplace system in equilibrium?
- 5. Can stable isotope measurements inform long-term evolution of tidally heated bodies?

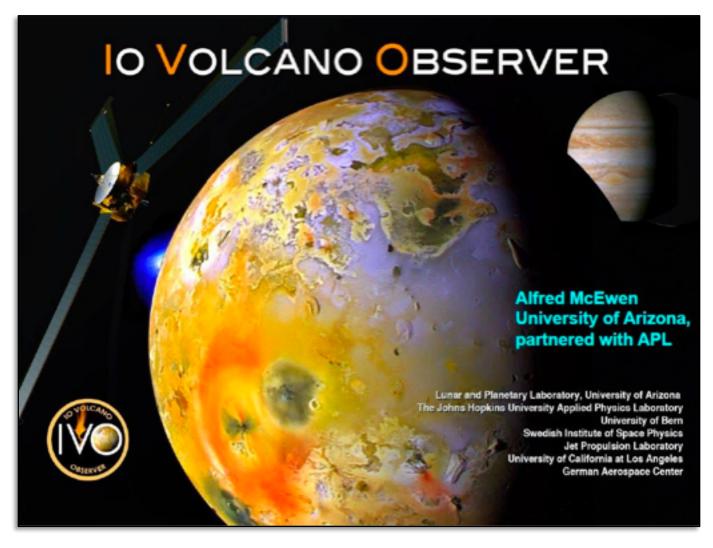
[←] Report cover (James T. Keane / Keck Institute for Space Studies)



KEY POINTS:

- The limited Voyager, Galileo, New Horizons, and Juno datasets have revealed new mysteries.
- 2. Earth-based observations have dramatically improved, enabling new scientific discoveries.
- 3. Nonetheless, there are major knowledge gaps—
 particularly with regards to geophysical and
 geochemical measurements (e.g., gravity/magnetic
 fields, atmospheric chemistry, etc.). These can only
 be addressed by a new mission to Io.

IVO is one of four finalists for the current round of NASA Discovery Program.



← Artist's conception of lo Volcano Observer (JHUAPL / U. Arizona)

More information about: https://ivo.lpl.arizona.edu/

PRE-DECISIONAL INFORMATION— FOR PLANNING AND DISCUSSION PURPOSES ONLY



IVO would be the first mission dedicated to Io, the most volcanically active world in the solar system. The IVO mission design and payload would transform our understanding of this unique world and how tidal heating shapes planetary worlds across the cosmos.

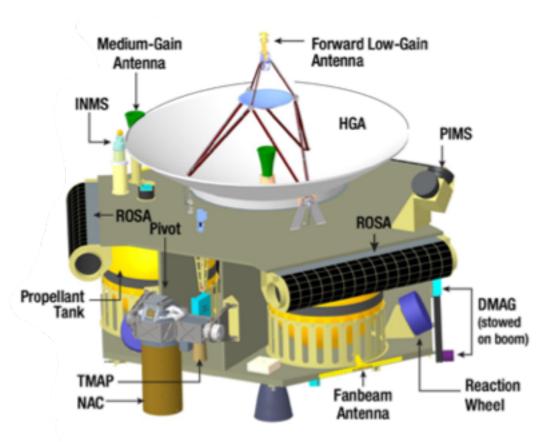
IVO's primary goal would be to 'Follow the Heat' and understand how tidal heat (A) is generated, (B) is lost, and (C) drives the evolution of Io.

IVO science objectives:

- A1. Determine the degree and distribution of melt in within Io's interior.
- B1. Determine Io's lithospheric structure.
- *B2. Determine where and how Io is losing heat.*
- C1. Determine Io's orbit evolution.
- C2. Determine Io's volatile loss processes and rates.

← Artist's interpretation of *IVO* interrogating the interior structure of Io (James T. Keane)

PRE-DECISIONAL INFORMATION — FOR PLANNING AND DISCUSSION PURPOSES ONLY



† Artist's conception of IVO design, with solar arrays and boom stowed (JHUAPL / U. Arizona)

IVO mission design

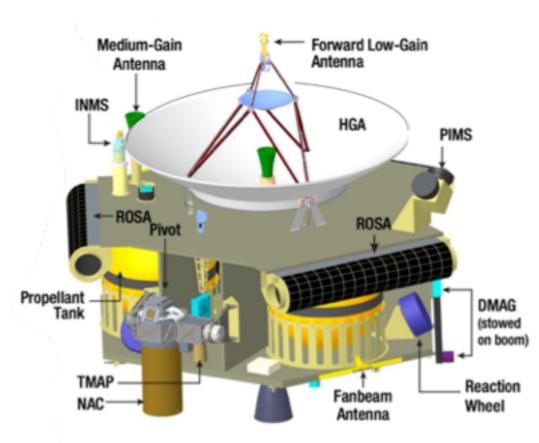
- Jupiter-orbit with 10 close (300-1,000 km altitude) flybys of Io over 3.9 years.
- Orbit is optimized for geophysics, near global mapping, and radiation mitigation (<10% total ionizing dose of Galileo, Juice, or Europa Clipper).

IVO spacecraft design:

Solar powered.

IVO baseline science experiments:

- Narrow Angle Camera (NAC)
- Thermal Mapper (TMAP)
- Dual Fluxgate Magnetometers (DMAG)
- Plasma Instrument for Magnetic Sounding (PIMS)
- Gravity Science (GS)
- Ion and Neutral Mass Spectrometer (INMS)



↑ Artist's conception of IVO design, with solar arrays and boom stowed (JHUAPL / U. Arizona)

IVO mission design

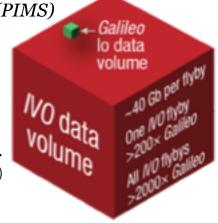
- Jupiter-orbit with 10 close (300-1,000 km altitude) flybys of Io over 3.9 years.
- Orbit is optimized for geophysics, near global mapping, and radiation mitigation (<10% total ionizing dose of Galileo, Juice, or Europa Clipper).

IVO spacecraft design:

Solar powered.

IVO baseline science experiments:

- Narrow Angle Camera (NAC)
- Thermal Mapper (TMAP)
- Dual Fluxgate Magnetometers (DMAG)
- Plasma Instrument for Magnetic Sounding (PIMS)
- Gravity Science (GS)
- Ion and Neutral Mass Spectrometer (INMS)



IVO data volume (JHUAPL / U. Arizona)

THE BILLION DOLLAR QUESTIONS:

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#1. "Is Io still relevant for New Frontiers?"

or

#2. "Since IVO has been selected for Phase-A study, is Io exploration still relevant for New Frontiers 5?"

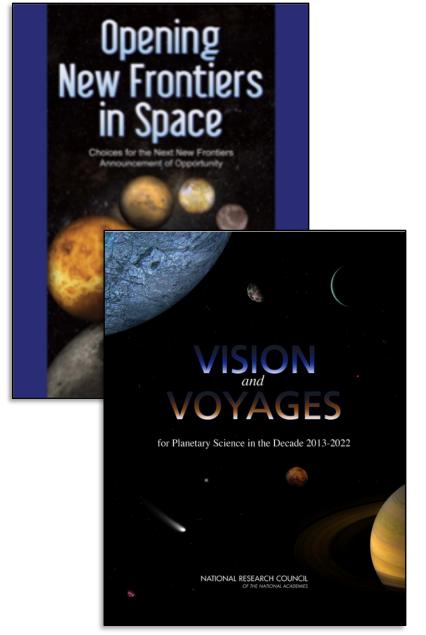
or

#3. "What would a New Frontiers mission do that a Discovery mission couldn't do?"

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NEW FRONTIERS v. DISCOVERY: Q: IS IO STILL RELEVANT?



Opening New Frontiers in Space:

"However, unlike Discovery, New Frontiers missions must be firmly grounded in scientific priorities established by the decadal survey without relying on new scientific or technology developments."

Vision and Voyages:

"In contrast to those for the Discovery program, solicitations for New Frontiers are more **strategic**, restricting proposals to a small number of specific mission goals. **New Frontiers missions** address focused science goals that cannot be implemented within the Discovery cost cap but that do not require the resources of a flagship mission."

NEW FRONTIERS v. DISCOVERY: Q: IS IO STILL RELEVANT?

Midterm review:

"...the committee
believes [Io
Observer and
Lunar Geophysical
Network] still
remain valid
missions for New
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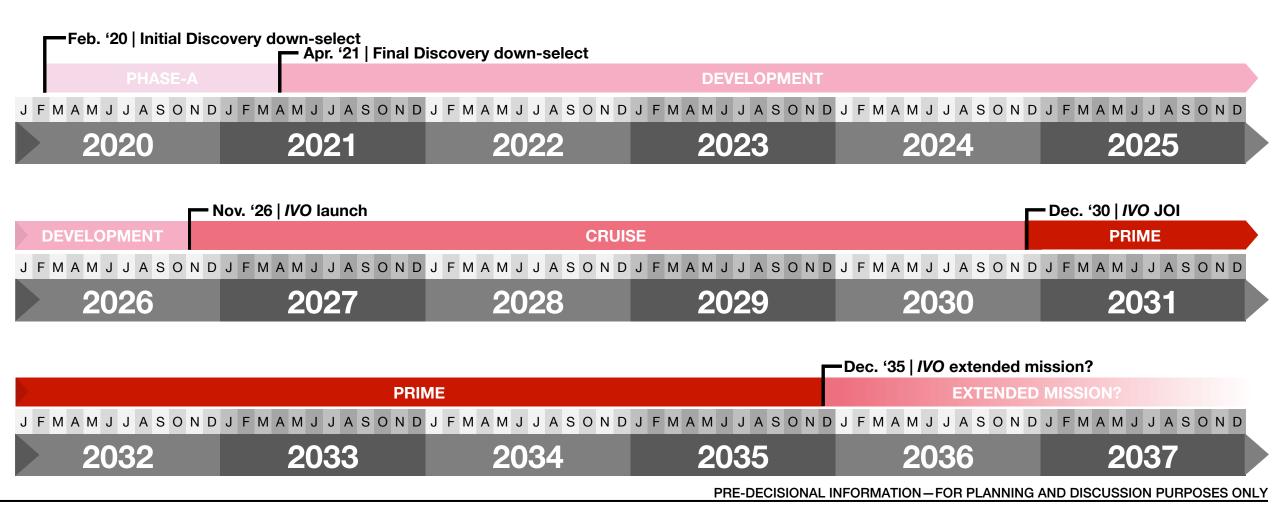
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KEY POINT:

IVO has not yet been selected for flight!

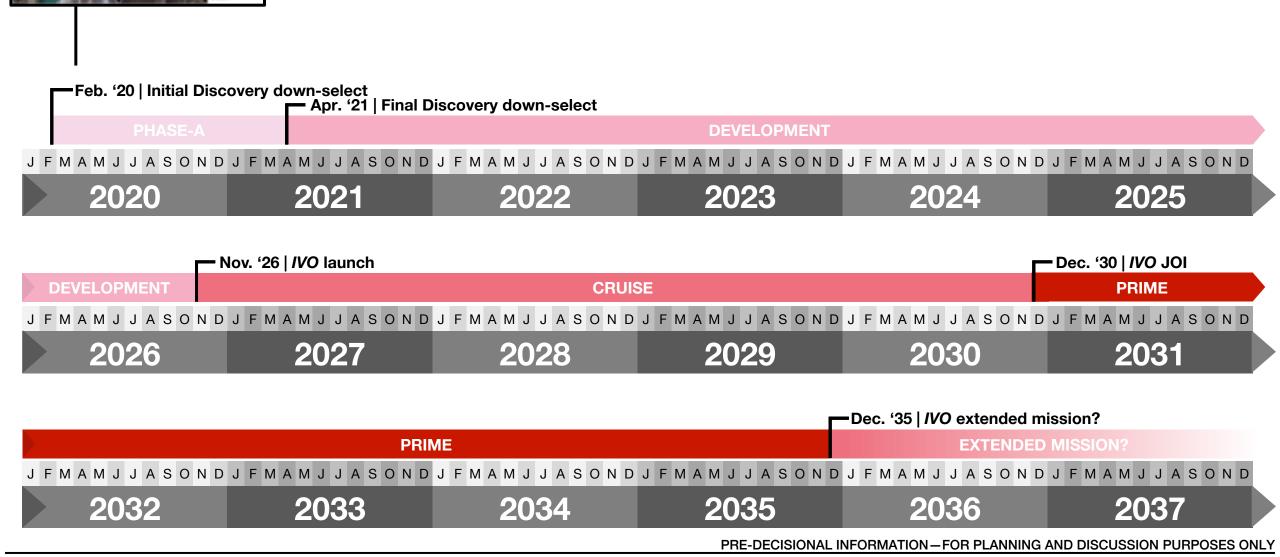
IVO TIMELINE

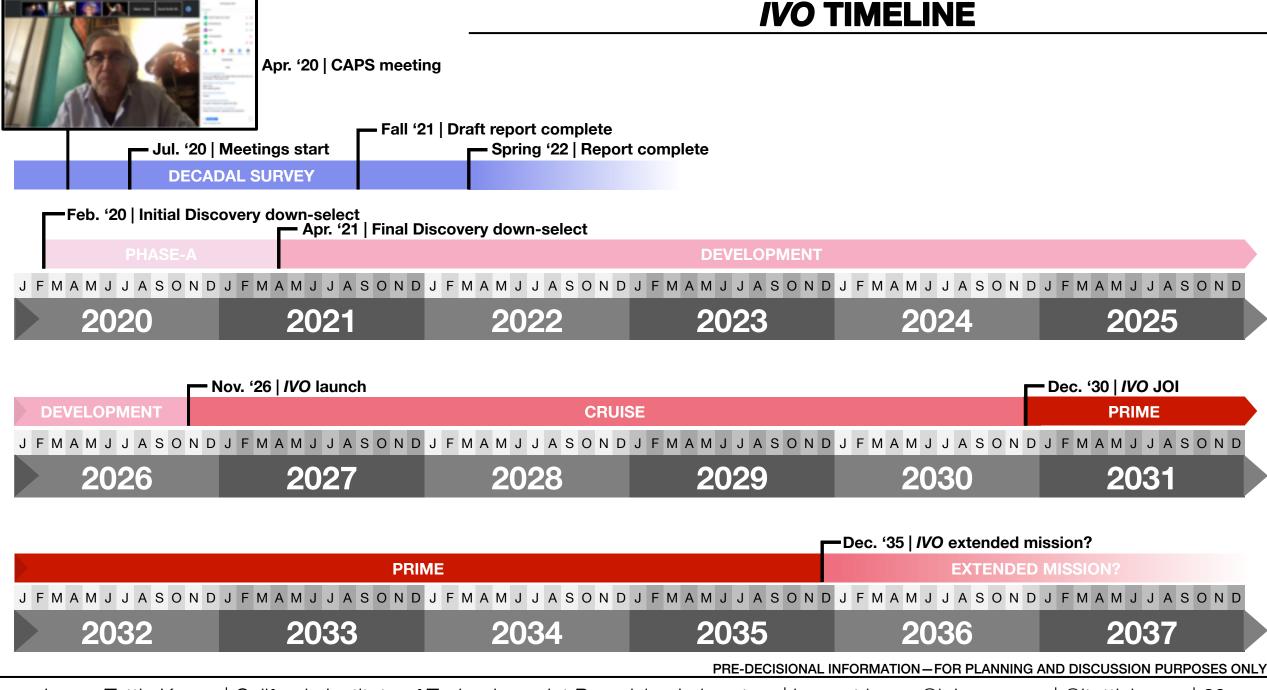


James Tuttle Keane | California Institute of Technology, Jet Propulsion Laboratory | james.t.keane@jpl.nasa.gov | @jtuttlekeane | 58

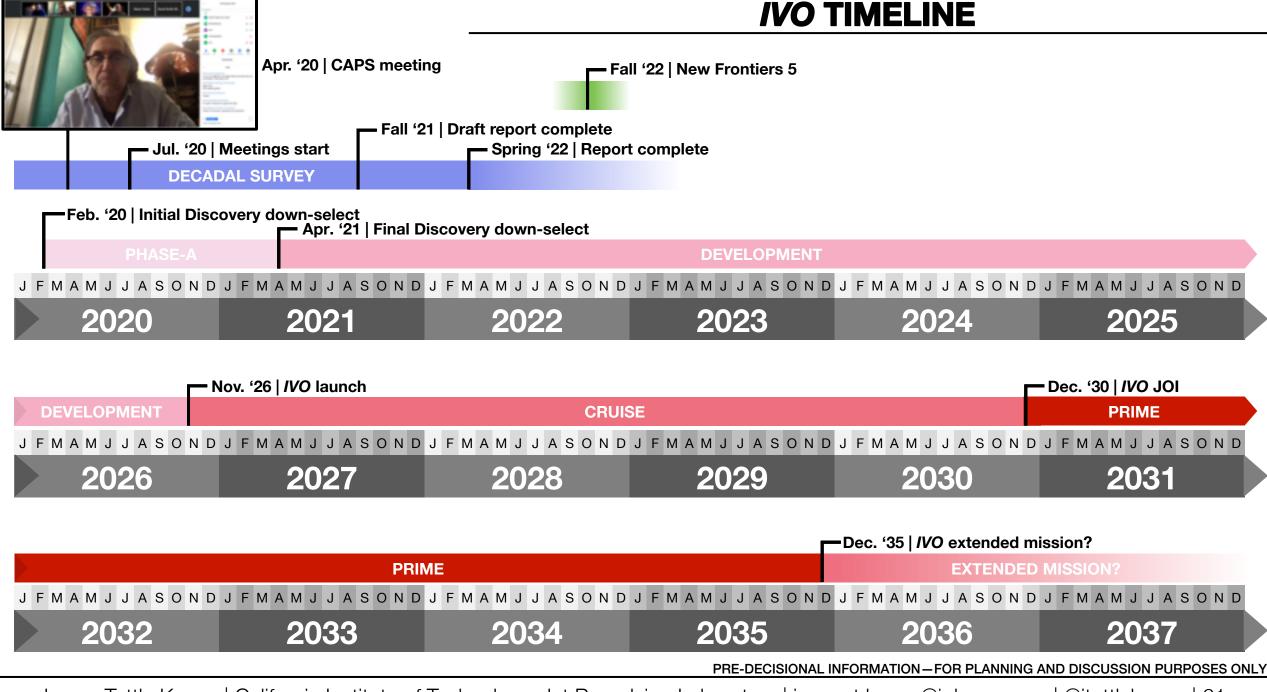
Apr. '20 | CAPS meeting

IVO TIMELINE





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NEW FRONTIERS 1
New Horizons





NEW FRONTIERS 3
OSIRIS-REx



NEW FRONTIERS 4

Dragonfly



NEW FRONTIERS 1 New Horizons



Juno

Proposed three times in Discovery (previously: INSIDE Jupiter)



NEW FRONTIERS 3 OSIRIS-REX

Proposed twice in Discovery (previously: OSIRIS)



NEW FRONTIERS 4 Dragonfly



NEW FRONTIERS 1 New Horizons

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(but many funded concept studies in/near the Discovery cost range: Pluto 350, Pluto Fast Flyby, etc.)



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NEW FRONTIERS 2
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NEW FRONTIERS 4

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Never proposed to Discovery

(but other innovative Titan exploration has been proposed: TIME)



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KEY POINT:

Discovery proposals are an important step in maturing concepts for New Frontiers proposals.

THE BILLION DOLLAR QUESTIONS:

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KEY POINT:

IVO has not yet been selected for flight!

We should let the Discovery proposal and selection process run its course. If IVO is selected, then the decadal survey will certainly need to reevaluate Io's inclusion on any New Frontiers list.

...Just like we are currently reevaluating the inclusion of Trojan Tour and Rendezvous and Ocean Worlds on the New Frontiers list after the selections of Lucy (Discovery) and Dragonfly (New Frontiers). I assume the decadal survey would also reevaluate the inclusion of Venus In-Situ Explorer on the New Frontiers list if VERITAS (Discovery Phase-A) and/or DAVINCI+ (Discovery Phase-A) were selected, or even revisit the choice between Uranus and Neptune flagship if Trident (Discovery Phase-A) was selected.

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A New Frontiers mission could do exactly what IVO is proposing to do.

...It is entirely possible IVO will not be selected for Discovery based on cost. As proposed to Discovery, IVO would address all of the science objectives outlined for a New Frontiers class Io Observer mission, and thus would be relevant to New Frontiers without any changes to the spacecraft, mission design, payload, etc.

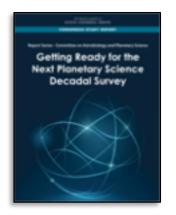
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KEY POINT:

The planetary science community would benefit from a new Io mission concept study.

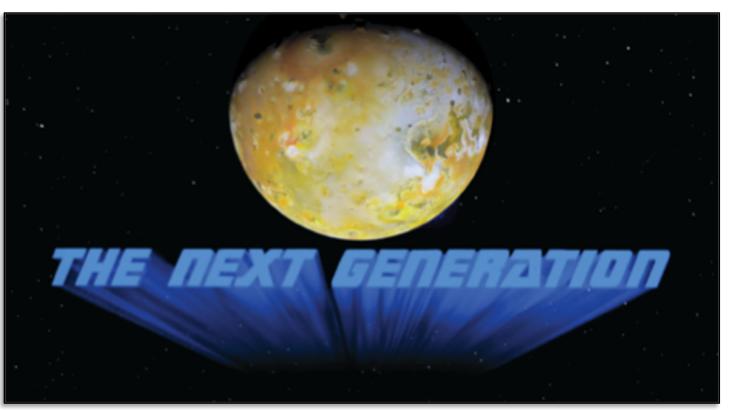
This was also recommended in "Getting Ready for the Next Decadal"

RECOMMENDATION: A NEW FRONTIERS IO MISSION CONCEPT STUDY



At present, the only available mission concept study is the Io Observer concept from the previous decadal. While compelling, this study does now outdated due to scientific and technological advances.

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Additional technologies include: high-dynamic range thermal imaging (Davies et al. 2016), geodetic cameras (Park et al. 2016), radar (Steinbrügge et al. 2018), microwave radiometers, penetrators, CubeSats, etc.

RESPONSE TO THE QUESTIONS IN THE CAPS STATEMENT OF TASK

1. Has scientific understanding or external factors such as programmatic developments or technological advances, significantly changed since the release of the planetary science decadal survey or its midterm review?

2. Has scientific understanding or external factors, such as programmatic developments or technological advances, been sufficiently substantial to warrant reconsideration of the four targets for inclusion in the New Frontiers 5 announcement of opportunity, scheduled for release in early 2022?

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No; if anything, the scientific, technological, and programmatic advances since the last decade only further demonstrate the need for a dedicated lo mission in the coming decade. The science has been refined and focused, the technology has been advanced (largely due to Europa Clipper and other NASA investments), and the programmatic push for habitability highlights our critical knowledge gaps in the underlying physical processes (e.g., tidal heating).

While Io exploration is currently being considered in the Discovery program (Io Volcano Observer), this should not negatively impact the consideration of Io for New Frontiers at this time.

SUMMARY

Io is the best destination for studying tidal heating and extreme volcanism-prime processes shaping planetary evolution and habitability.

There have been multiple major scientific advances in the past decade that further motivate and focus Io exploration. There is a priority need for in situ geophysical and geochemical measurements.

While challenging, Io exploration is more compelling than ever before, and still entirely relevant for New Frontiers.

Io Volcano Observer is a bold Discovery mission concept for exploring Io, however it is not yet selected and should not negate Io's inclusion on the New Frontiers list.

A new mission concept study, if deemed appropriate by the decadal survey, could leverage major scientific and technological developments in the past decade.

