

Predicting IMF B_z at 1 AU

-The past, present, and future

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BLUF

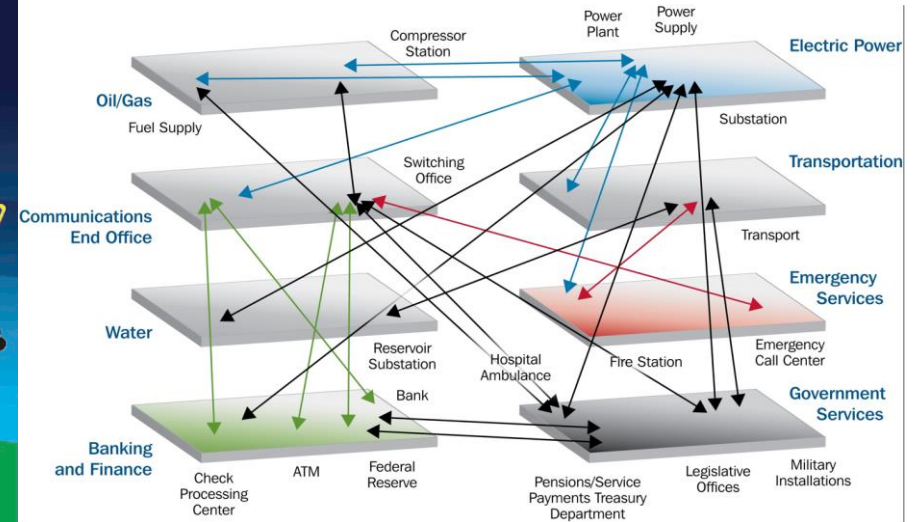
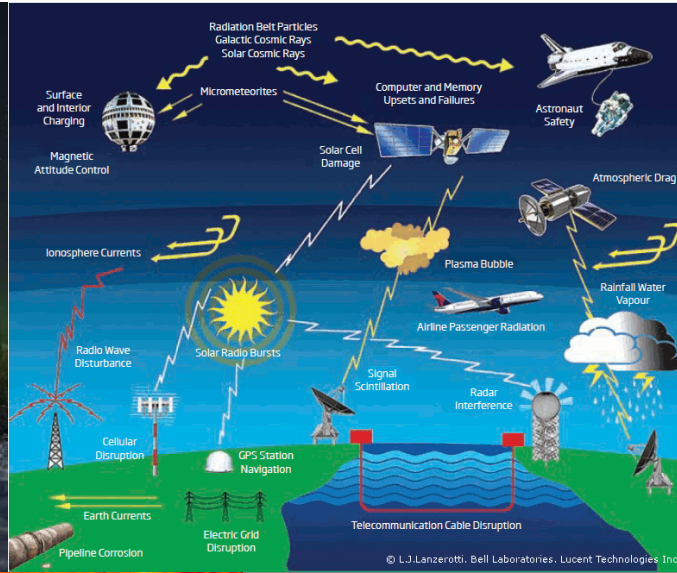
- We can't predict IMF B_z (well)
- Anyone who tries to convince you otherwise is misinformed, deluded, or lying
- Reproducing 1 AU observations is not the same as forecasting/predicting
- Being able to make useful forecasts of B_z is going to be more difficult than we (or at least I) thought
- If forecasting is the goal, we have to be goal-orientated, not theory-orientated
- If B_z prediction is really a priority, it will require significant investment

Overview

- Why do we care?
- What physical processes produce B_z ?
- How can we predict non-zero B_z ?
- Past and current status of B_z predictions
- Basic uncertainties that need to be resolved
- Conclusions
- Future Opportunities

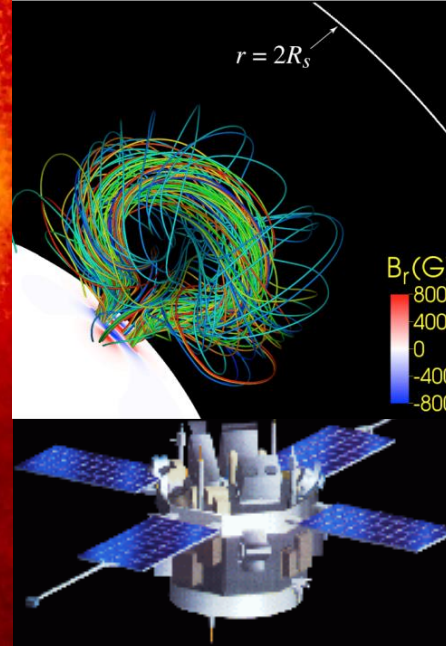
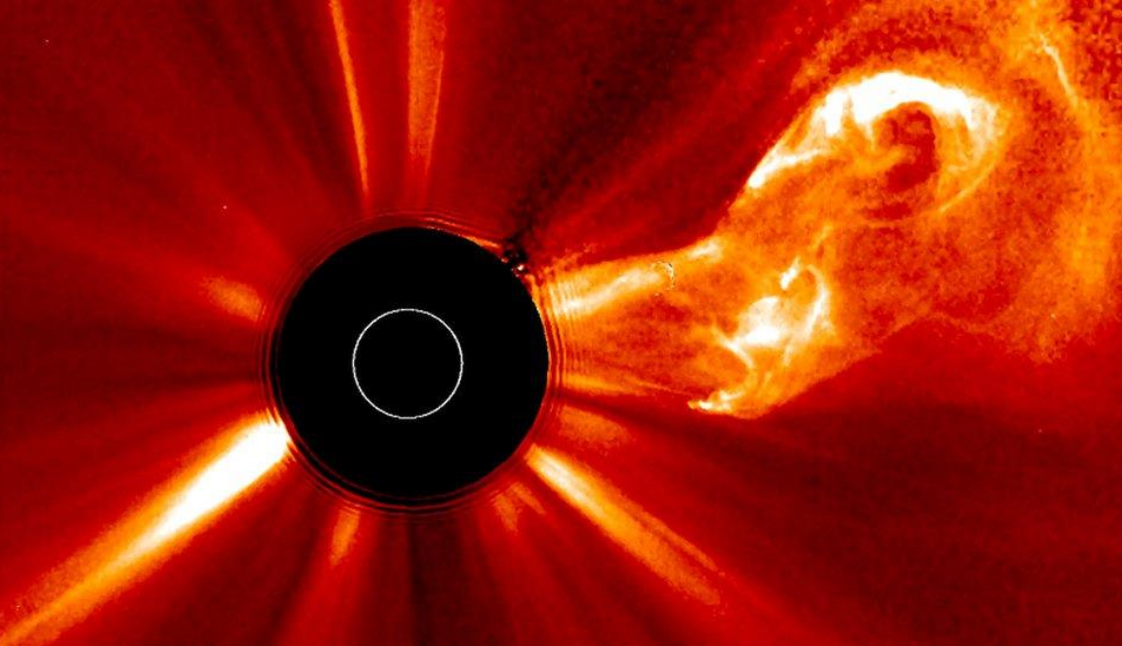
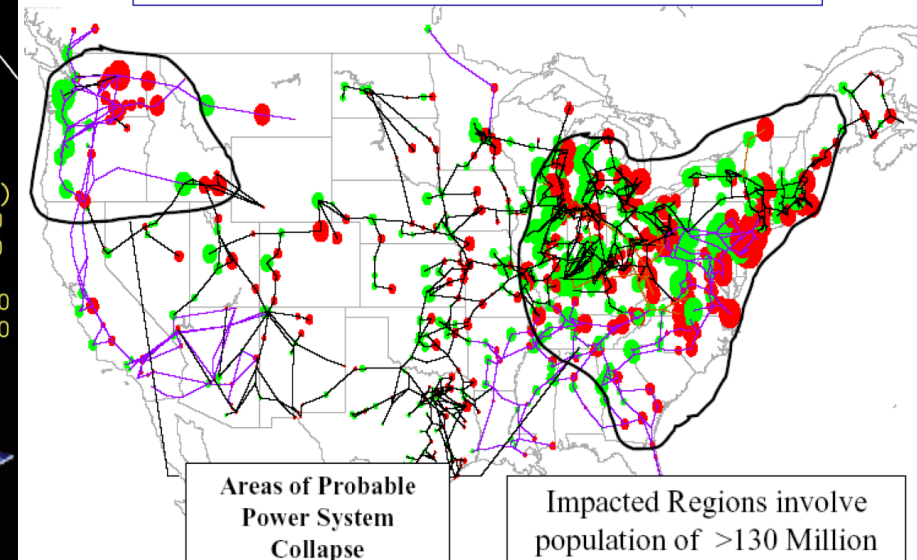
Why do we care about B_z ?

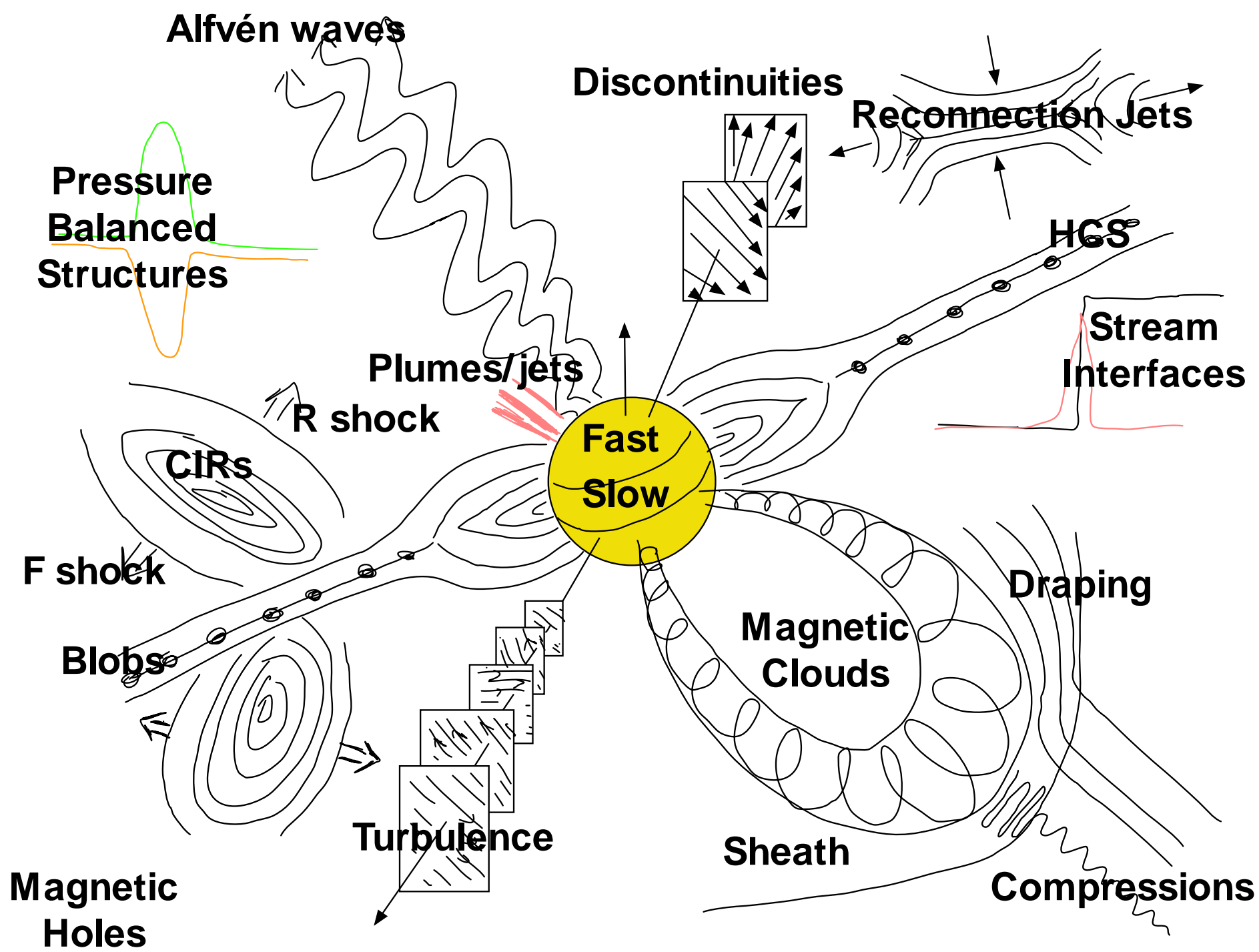
-Science and Society



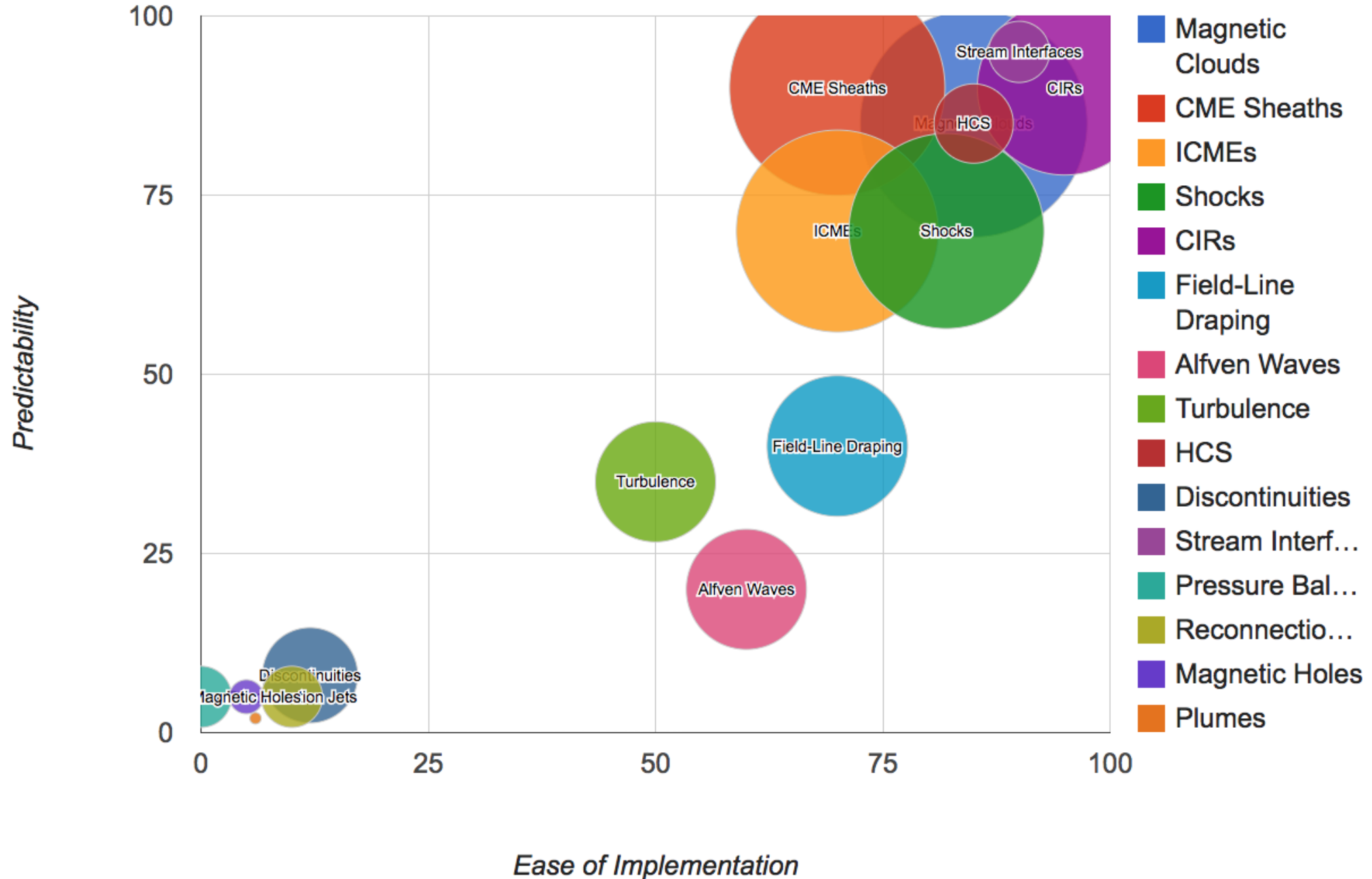
Severe Electrojet Disturbance Scenario

Power System Disturbance and Outage Scenario of Unprecedented Scale

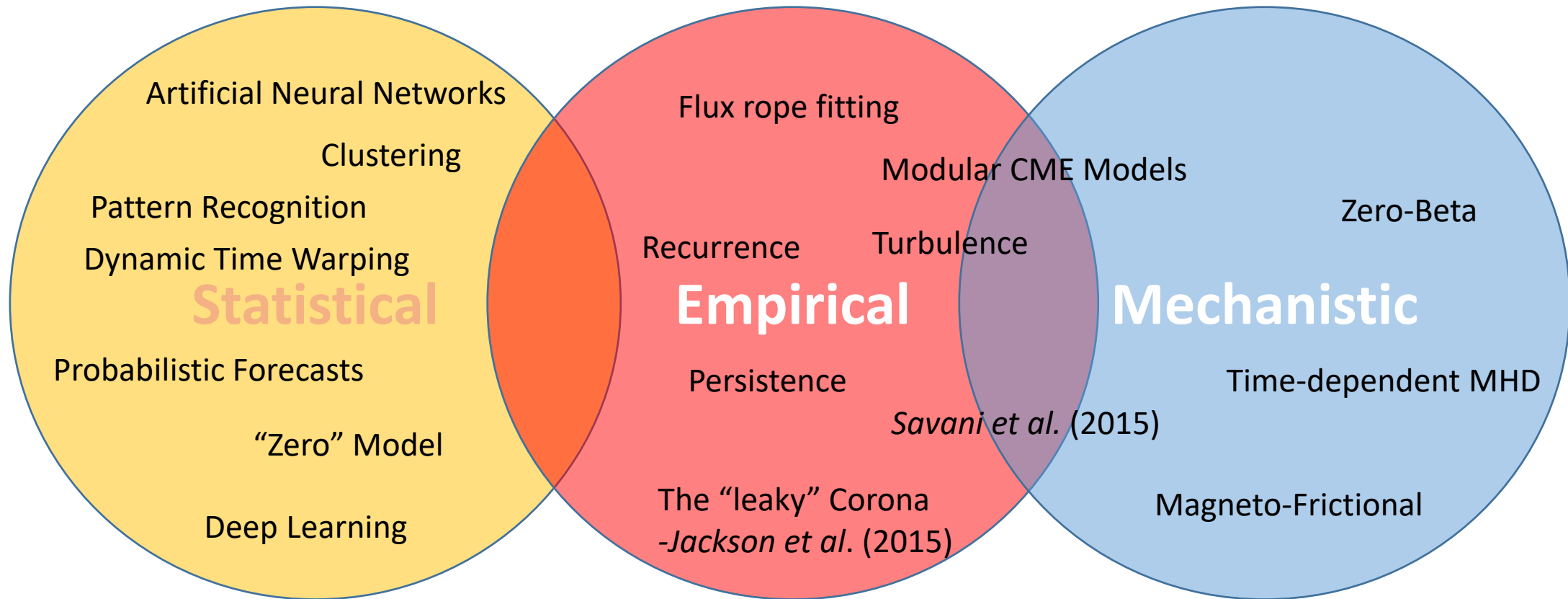




How do we Prioritize these Processes?



Techniques for predicting non-zero B_z



Bz 'prediction' studies...

- Search based on:
 - (interplanetary | "solar wind") AND (southward | Bz | "z component" | "z-component" | "magnetic vectors" | "magnetic field vectors") AND (prediction | predicting | forecast | forecasting)
- Resulted in ~20 relevant papers
 - Chen, James, Peter J. Cargill, and Peter J. Palmadesso. "Real-time identification and prediction of geoeffective solar wind structures." *Geophysical research letters* 23, no. 6 (1996): 625-628.
 - Riley, Pete, Michal Ben-Nun, Jon A. Linker, Mathew J. Owens, and T. S. Horbury. "Forecasting the properties of the solar wind using simple pattern recognition." *Space Weather* 15, no. 3 (2017): 526-540.
 - Jackson, B. V., H-S. Yu, A. Buffington, P. P. Hick, M. Tokumaru, K. Fujiki, J. Kim, and J. Yun. "A daily determination of BZ using the Russell-McPherron effect to forecast geomagnetic activity." *Space Weather* 17, no. 4 (2019): 639-652.
 - Sarkar, Ranadeep, Nat Gopalswamy, and Nandita Srivastava. "An Observationally Constrained Analytical Model for Predicting the Magnetic Field Vectors of Interplanetary Coronal Mass Ejections at 1 au." *The Astrophysical Journal* 888, no. 2 (2020): 121.
 - Savani, N. P., A. Vourlidas, A. Szabo, M. L. Mays, I. G. Richardson, B. J. Thompson, A. Pulkkinen, R. Evans, and T. Nieves-Chinchilla. "Predicting the magnetic vectors within coronal mass ejections arriving at Earth: 1. Initial architecture." *Space Weather* 13, no. 6 (2015): 374-385.
 - Möstl, Christian, Tanja Amerstorfer, Erika Palmerio, Alexey Isavnin, Charles J. Farrugia, Chris Lowder, Reka M. Winslow, Julia M. Donnerer, Emilia KJ Kilpua, and Peter D. Boakes. "Forward modeling of coronal mass ejection flux ropes in the inner heliosphere with 3DCORE." *Space Weather* 16, no. 3 (2018): 216-229.
 - Kay, C., and N. Gopalswamy. "Using the coronal evolution to successfully forward model CMEs' in situ magnetic profiles." *Journal of Geophysical Research: Space Physics* 122, no. 12 (2017): 11-810.

“Historical” attempts to predict B_z ...

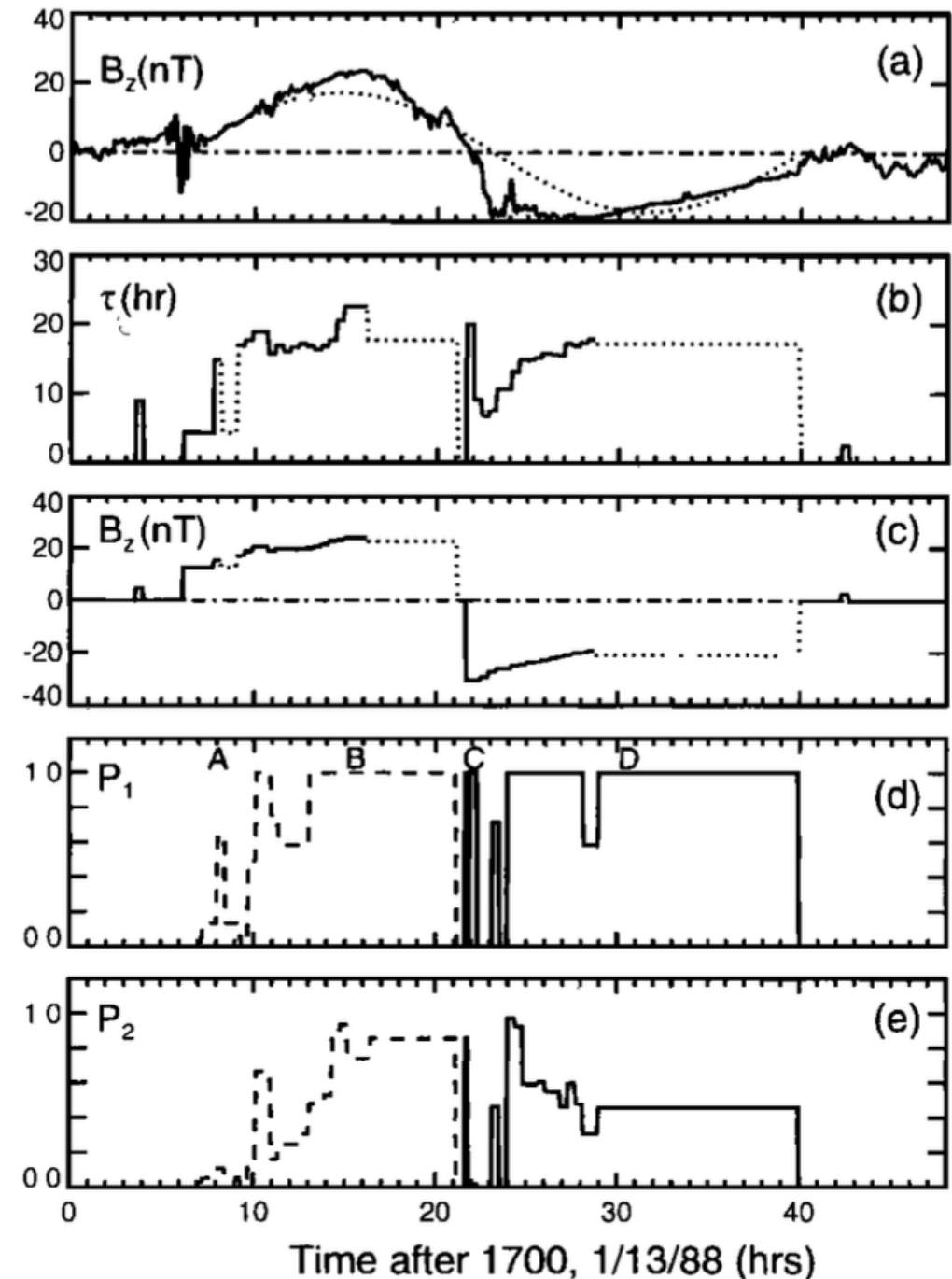
GEOPHYSICAL RESEARCH LETTERS, VOL. 23, NO. 6, PAGES 625-628, MARCH 15, 1996

Real-time identification and prediction of geoeffective solar wind structures

James Chen, Peter J. Cargill, Peter J. Palmadesso

Plasma Physics Division, Naval Research Laboratory

- Feature-based classification technique
- Extrapolation of coherent signal
- Precursor to flux-rope fitting and/or pattern recognition
- Focuses on sinusoidal variation of B_z
 - If $B_z < 1$ during first-half => limited prediction
 - If $B_z > 0$ during first-half => longer prediction
- Danger of false positives
- Only demonstrated for a few case studies





Predictive Science - Project Zed

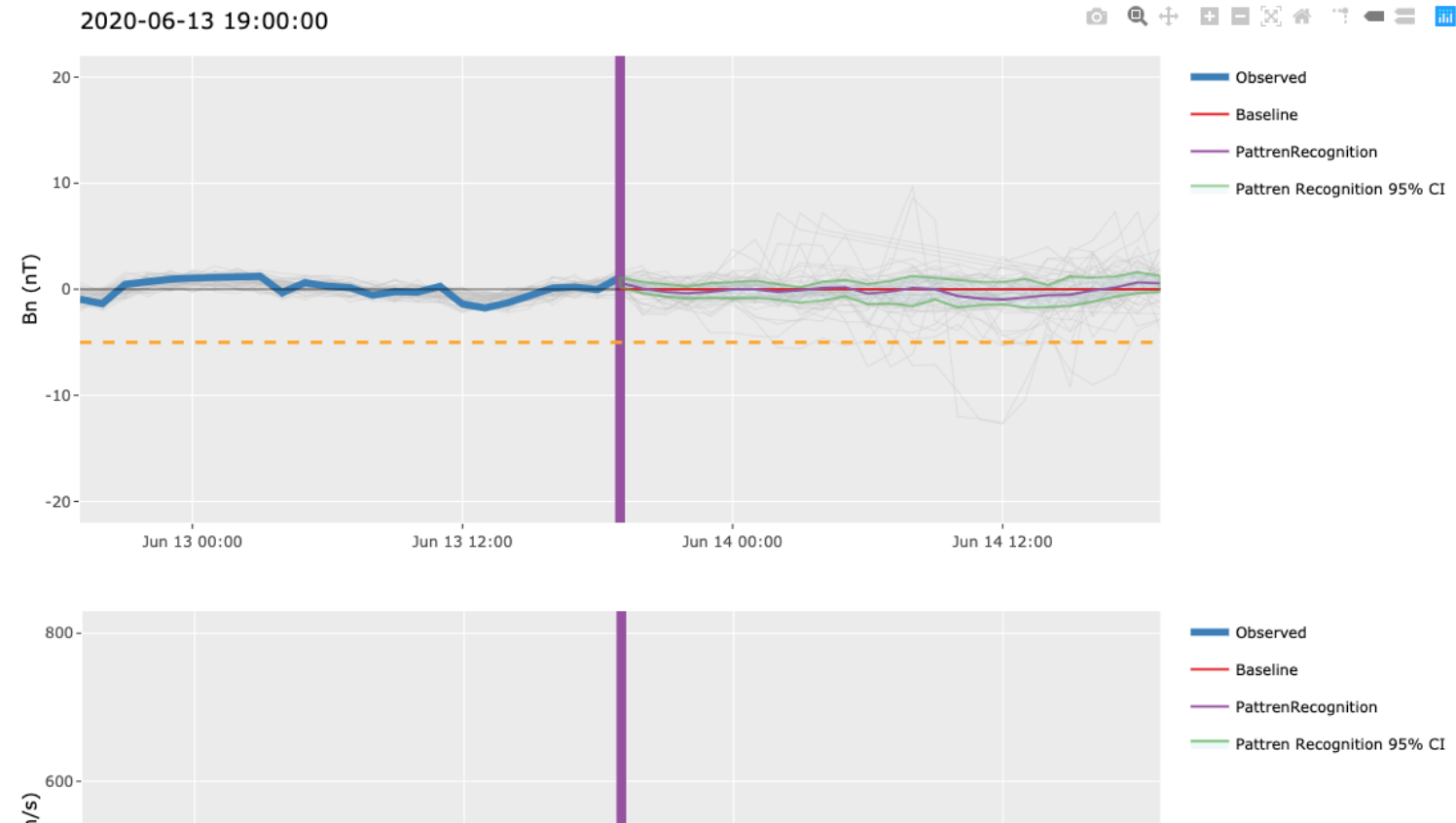
Data & Forecast

Statistics

About

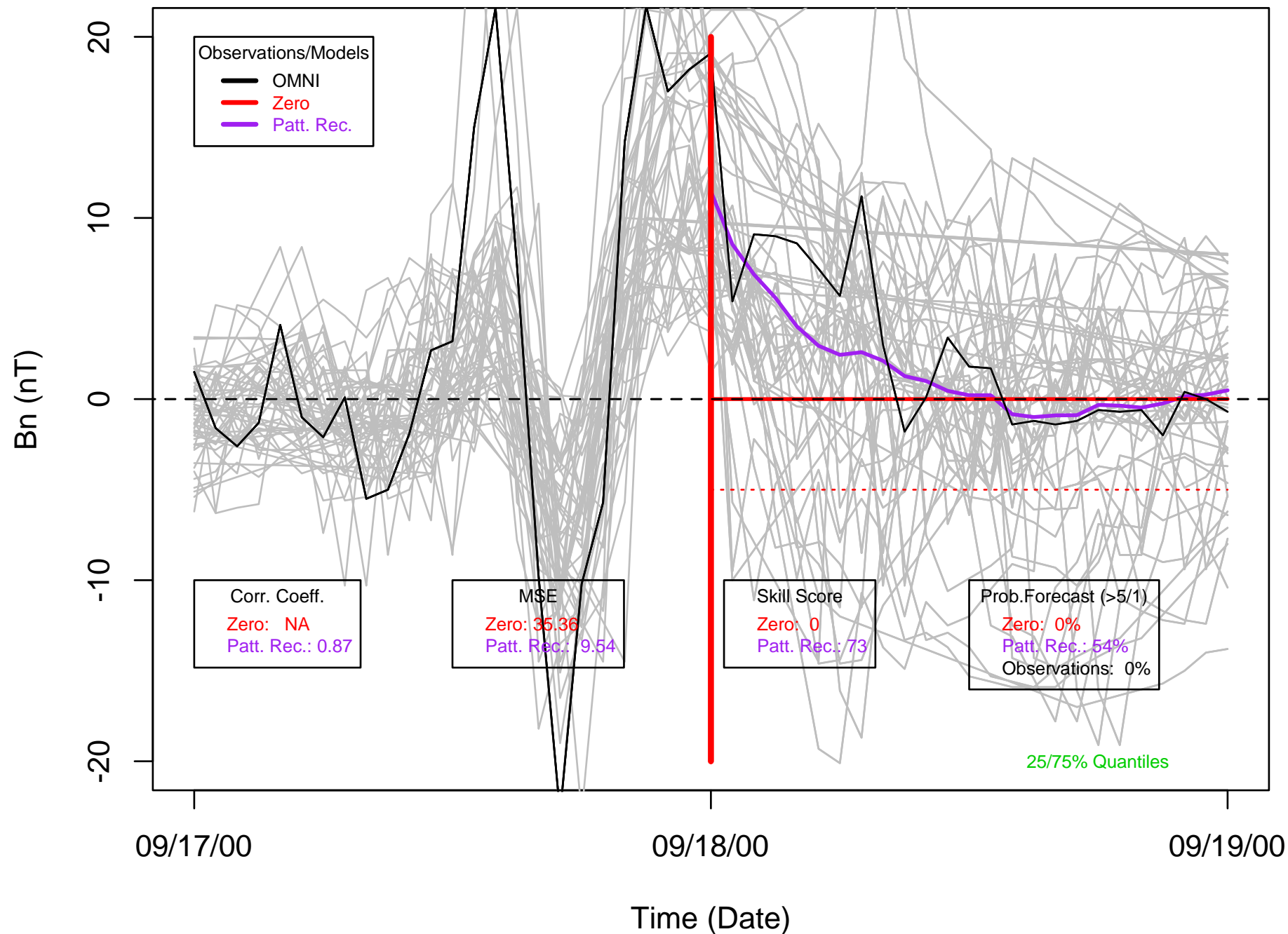
Contact

Solar Wind Data and Forecast

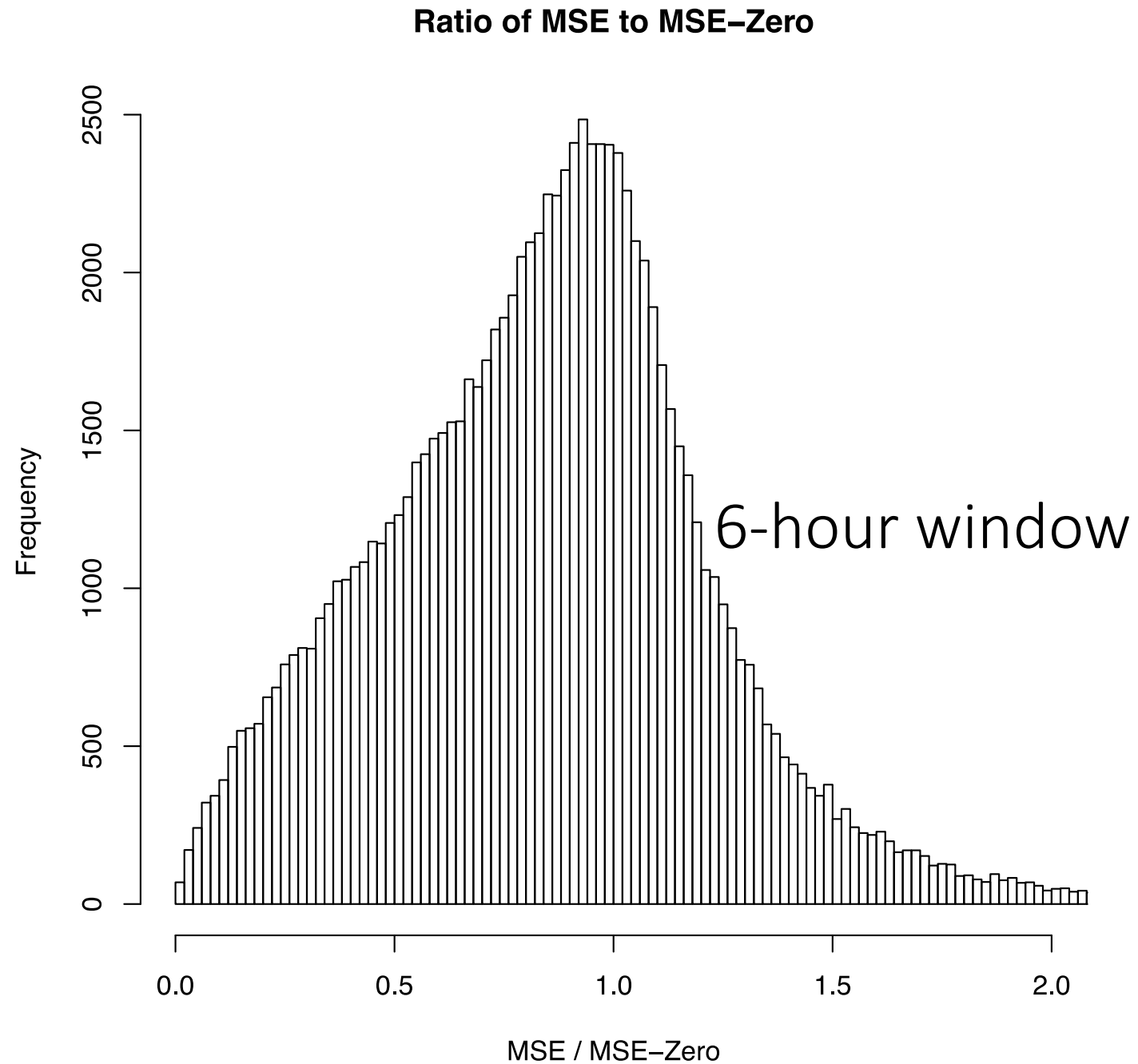


Riley et al. (2017)

Pattern Recognition *Riley et al.* (2017)



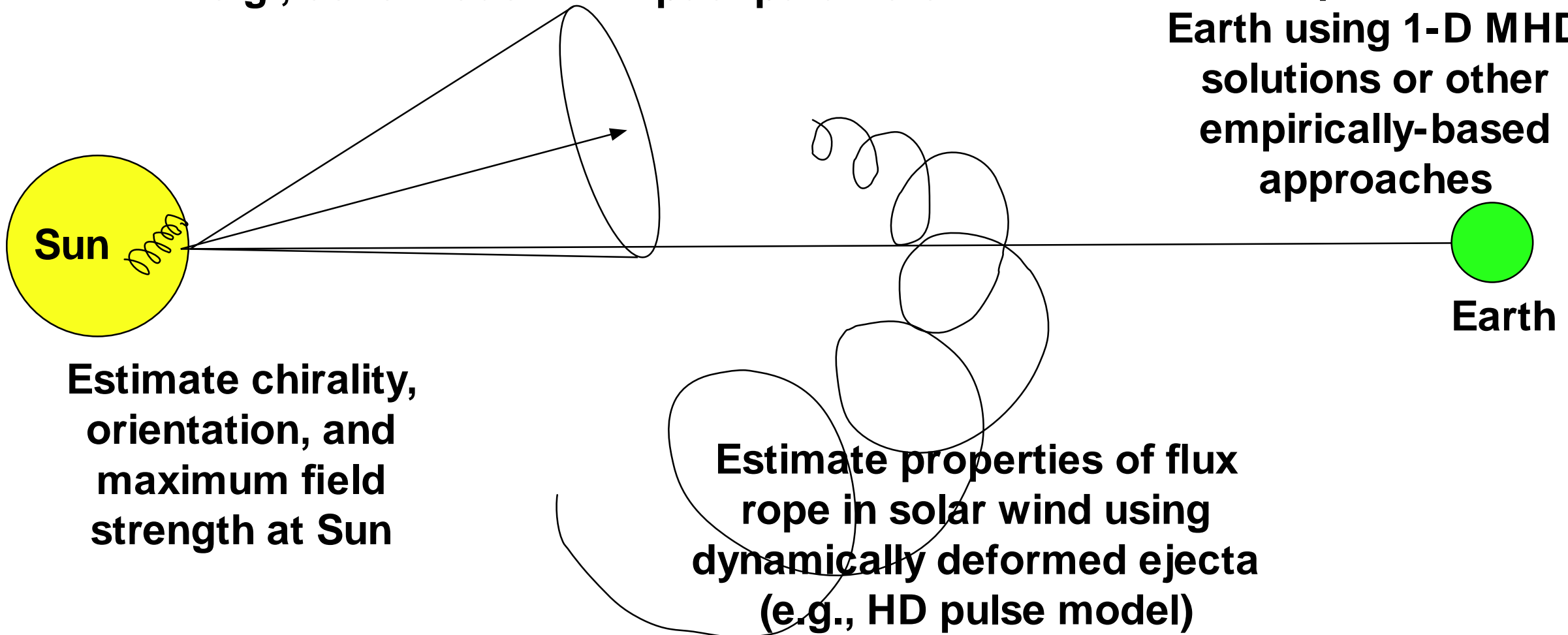
Pattern Recognition



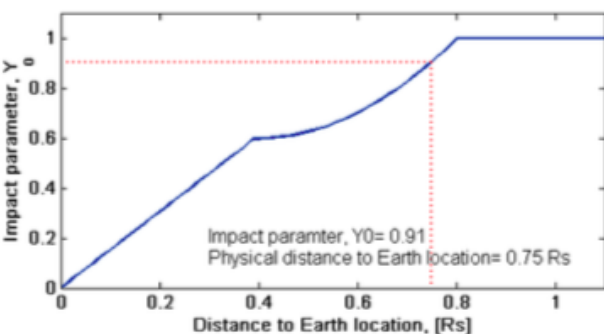
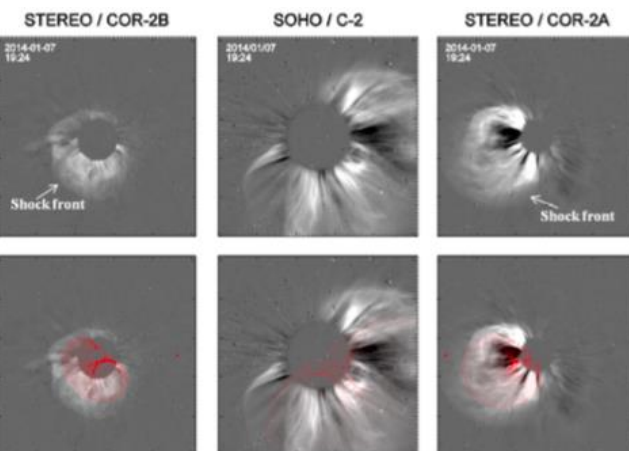
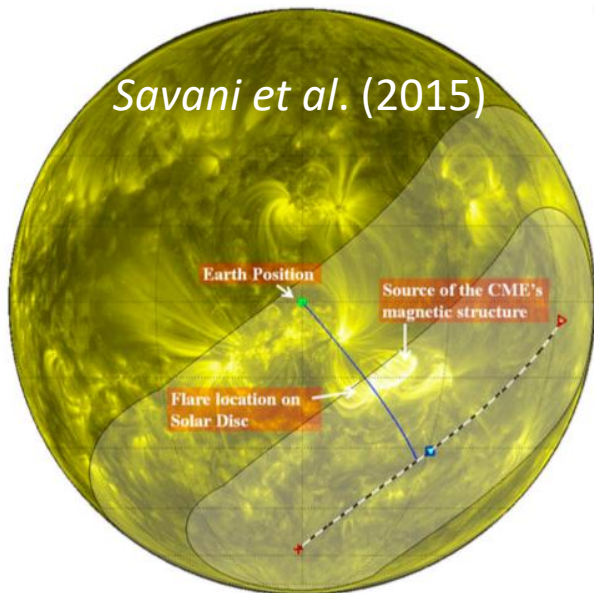
Empirical Modular CME (EMC) Models

**Estimate kinematic/geometric
properties of ejecta near Sun
e.g., cone model => Impact parameter**

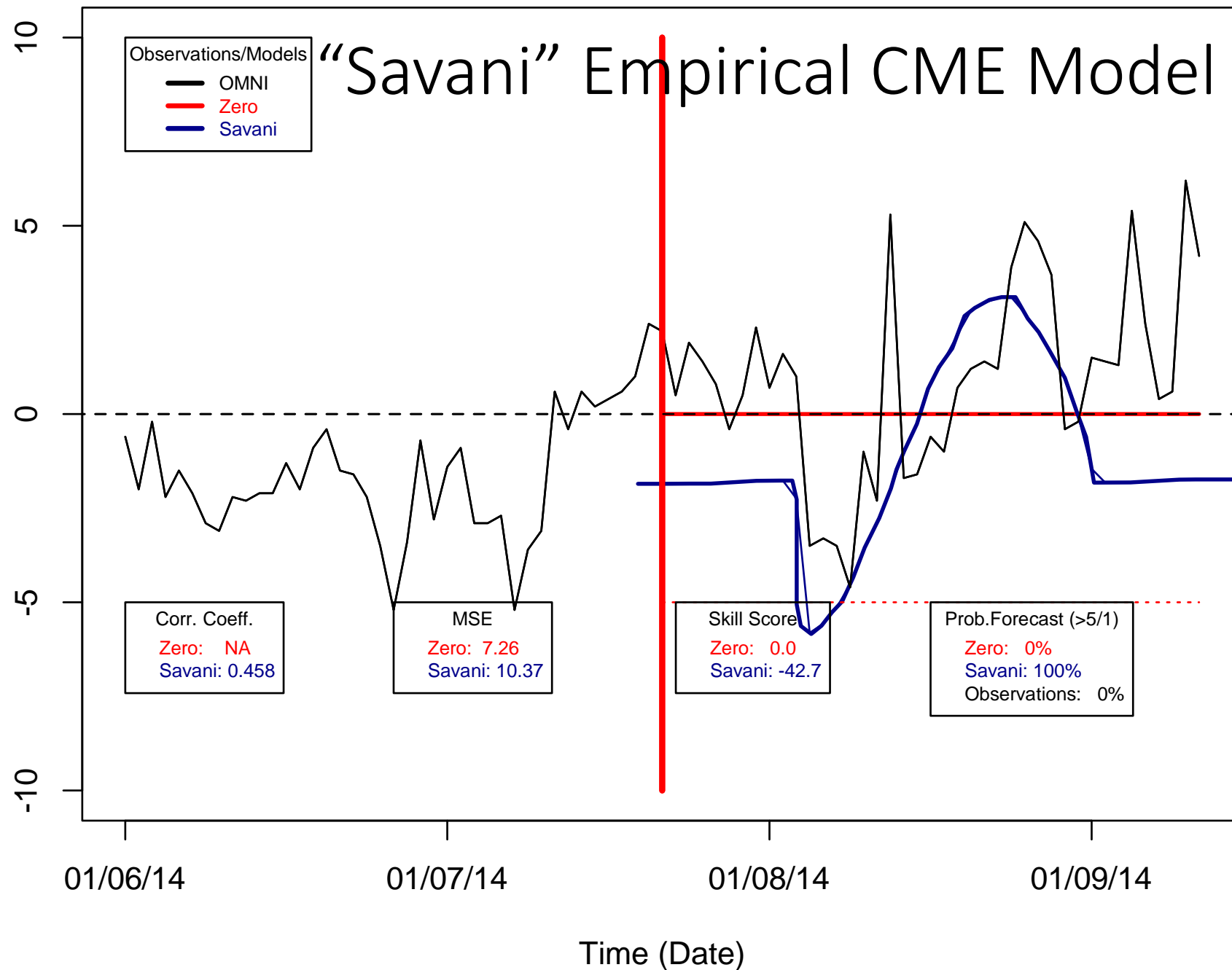
**Estimate transit time (τ)
of flux rope from Sun to
Earth using 1-D MHD
solutions or other
empirically-based
approaches**



Savani et al. (2015)



Bn (nT)



B_z Prediction: A Comedy of Errors

- Initial helicity (accuracy ~ 0.75)
- Photospheric FR axial orientation ($\theta \pm 5^\circ$, $\phi \pm 5^\circ$)
- Flux rope rotation ($\sim \theta \pm 20^\circ$, $\phi \pm 20^\circ$)
- Ejecta speed ($\sim \pm 100$ km/s)
- Ejecta mass ($\sim \pm 1 \times 10^{12}$ kg)
- Ejecta propagation direction ($\theta \pm 15^\circ$, $\phi \pm 15^\circ$).

B_z Prediction: CME/ICME relationships

- 46% of ICMEs have no obvious solar event associated with them (*Richardson et al. (2010)*)
 - But, many slow, weak events that are convected out with the solar wind.
- 10% of the ICMEs driving “Dst <-100 nT” storms had no clear associated solar event (*Zhang et al., 2007*)
- Only ~56% of halo CMEs produce MC at 1 AU

Conclusions

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Suggestions for Future Paths for Forecasting B_z ?

- Invest time/money specifically into forecasting, not coopting research studies
- Work on specific components of the Sun-Earth chain to reduce uncertainties
- Investigate statistical approaches more carefully (c.f. weather forecasting)
- Optimize NASA/ESA/NOAA mission designs to measure critical inputs
- Start with simpler objectives: Predict speed (time of arrival), B , B_r , and B_t
- Identify and focus on “lowest hanging fruit”: CME sheaths, ultra-fast events, etc.