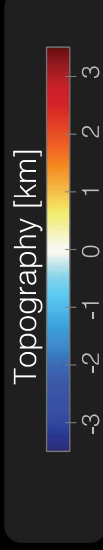
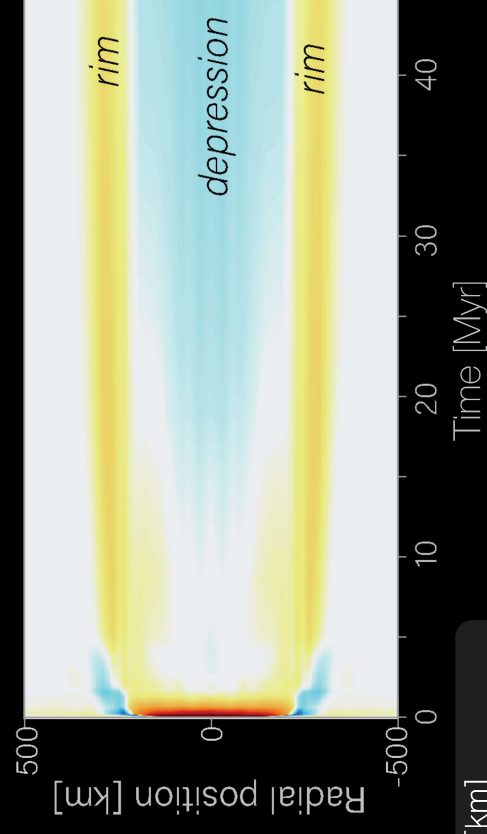
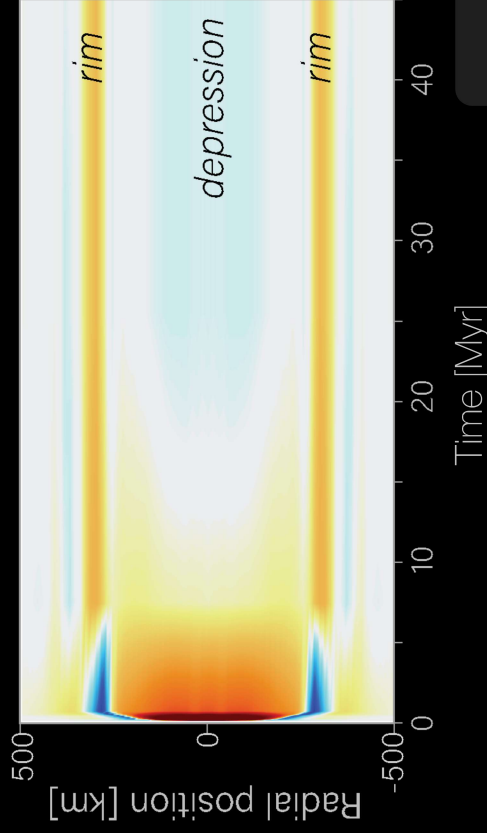
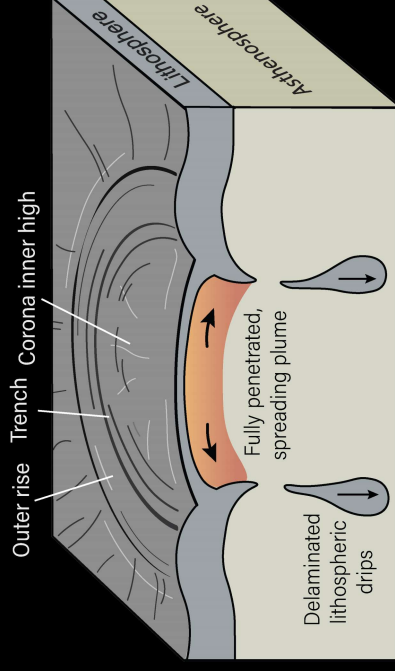
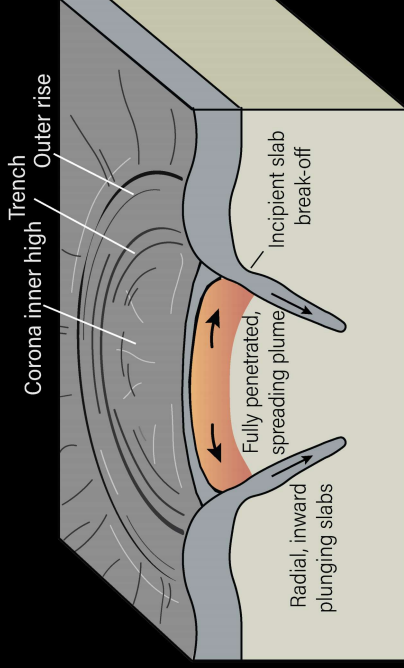
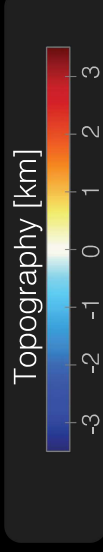
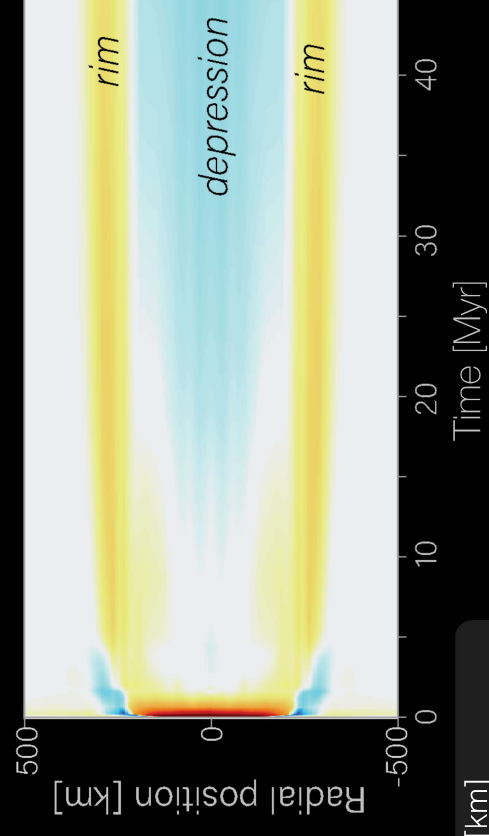
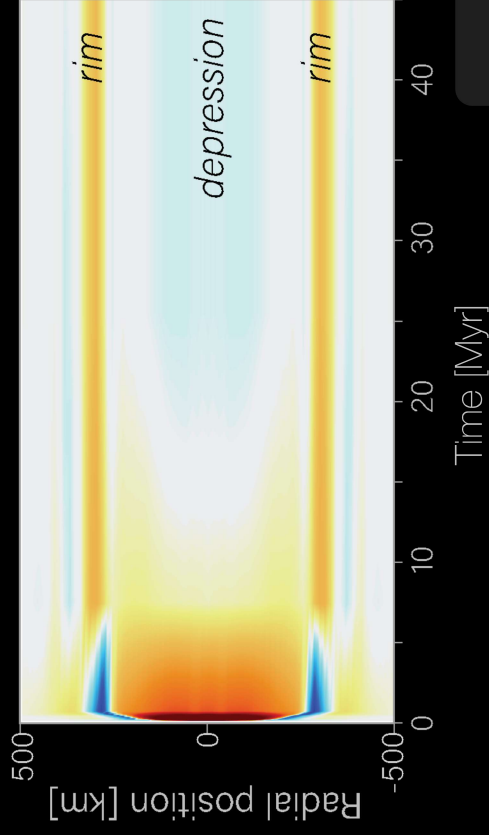
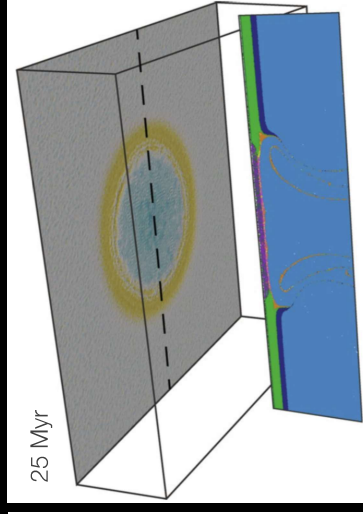
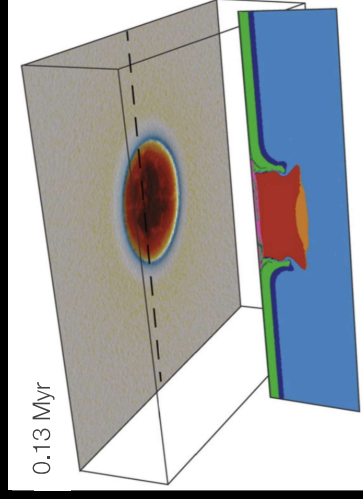
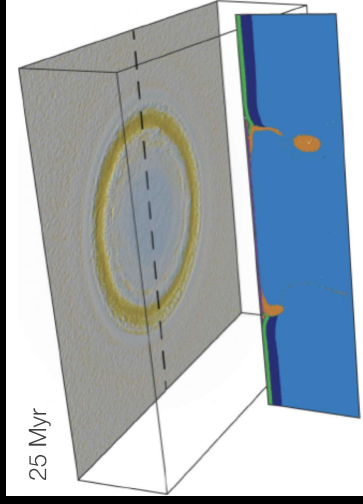
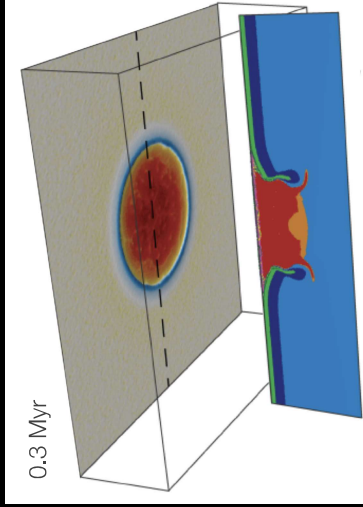


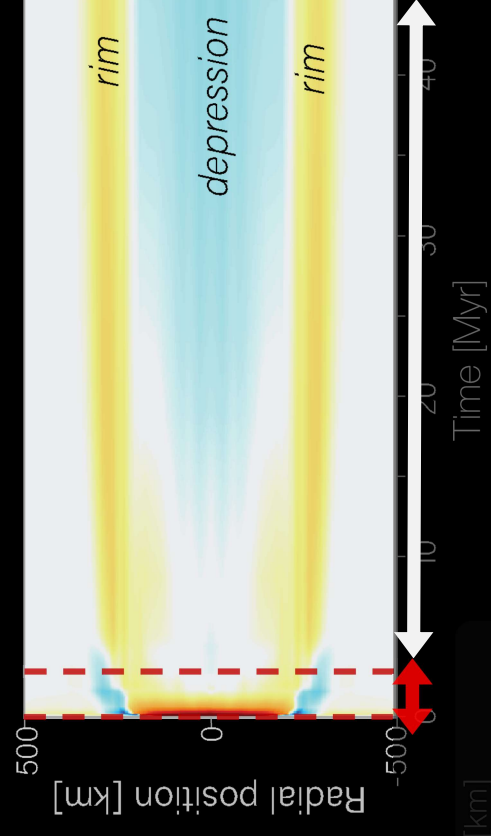
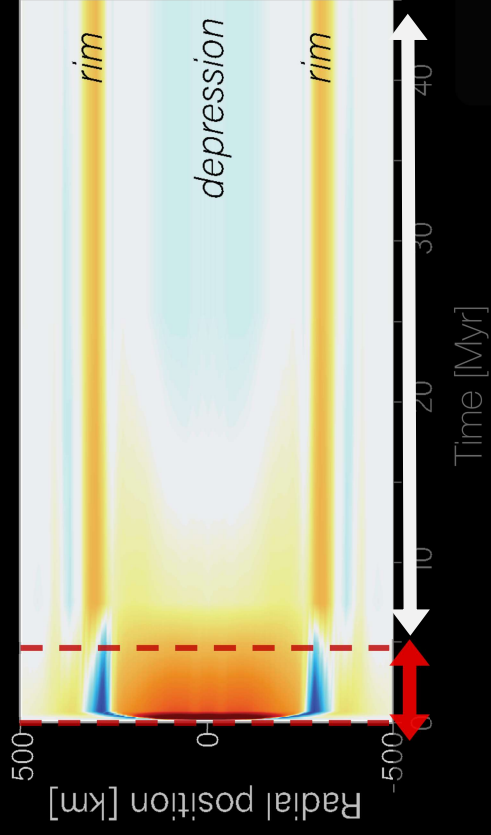
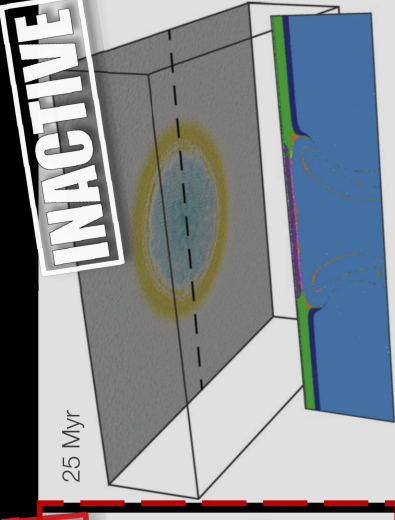
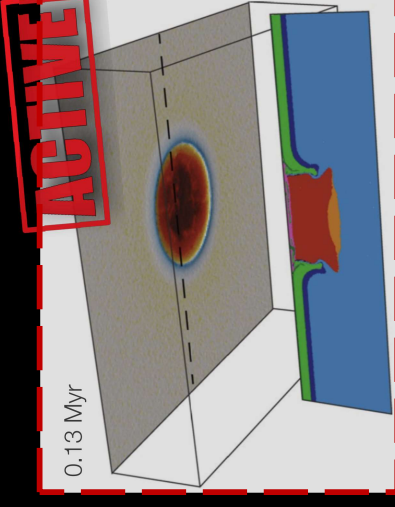
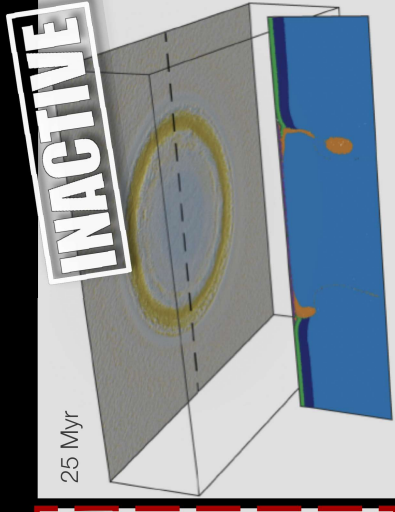
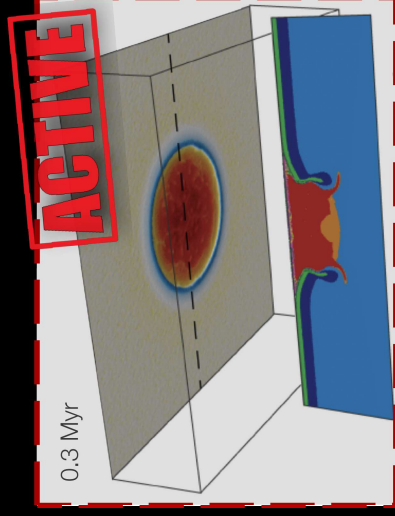
Temporal evolution of corona topography



Temporal evolution of corona topography



Topographic inversion



3D models – *key points*

- The morphology of coronae is related to what processes are going on underneath it

3D models – *key points*

- The morphology of coronae is related to what processes are going on underneath it
- Crustal thickness variations eventually lead to an isostasy-driven **topographic inversion**:
circular trench surrounding interior high → outer rim surrounding interior depression

3D models – key points

- The morphology of coronae is related to what processes are going on underneath it
- Crustal thickness variations eventually lead to an isostasy-driven **topographic inversion**:
circular trench surrounding interior high → outer rim surrounding interior depression
- The topographic pattern of coronae is therefore **indicative** for **plume (in)activity**

Address the **current state of plume activity** at large coronae on Venus

Coronae classification

- Analyse topography of 133 largest coronae on Venus

(Magellan topography dataset and USGS coronae database)

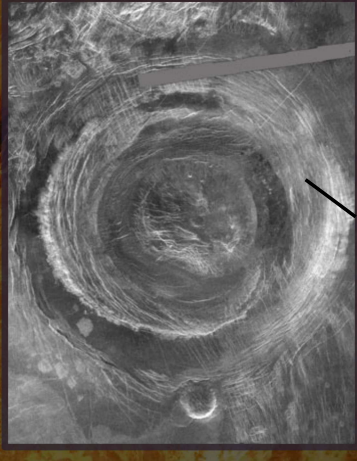
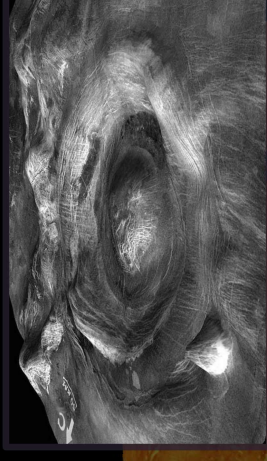
- Classify plume activity:

ACTIVE

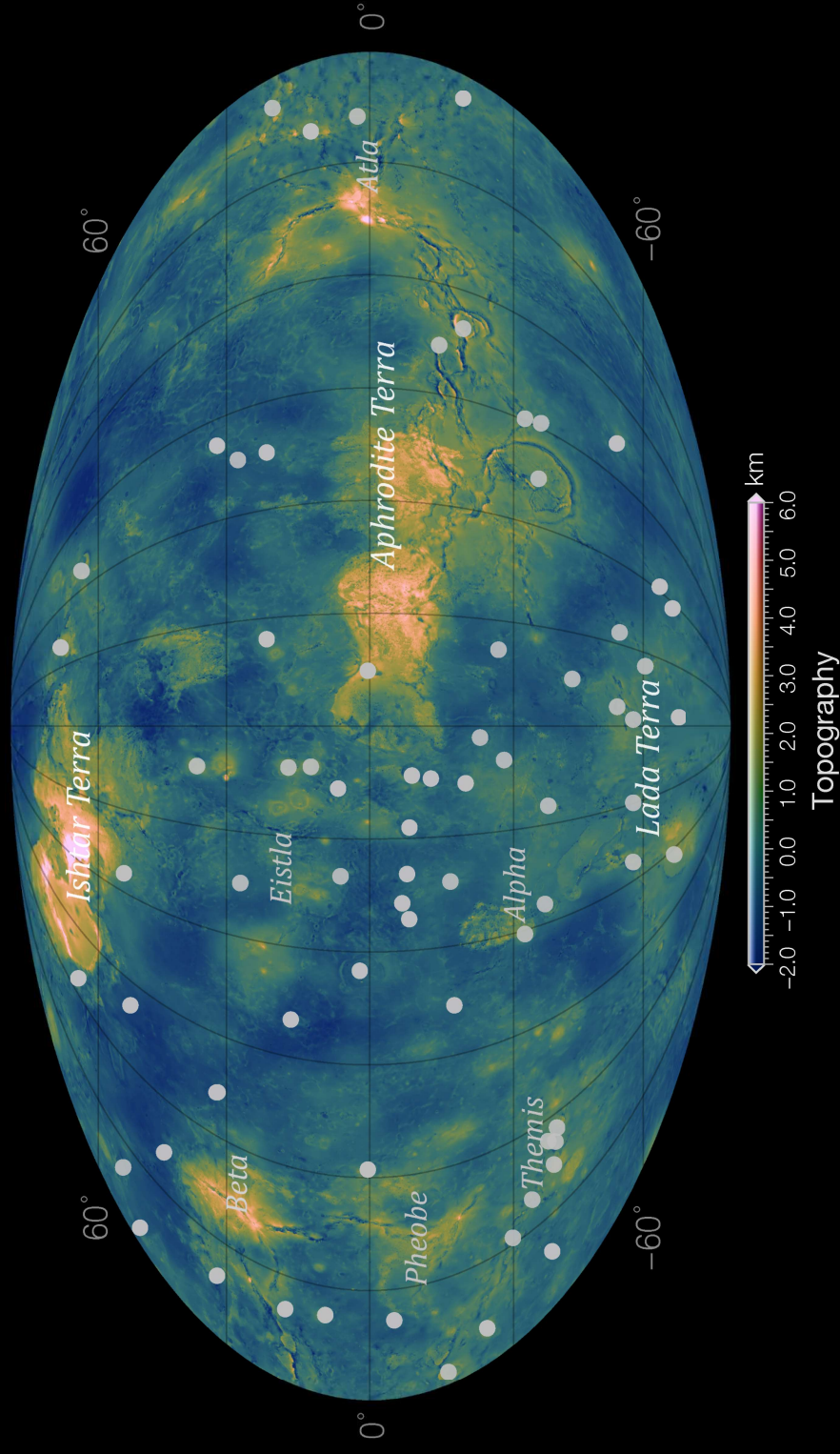
- ✓ Outer trench
- ✓ Outer rise

INACTIVE

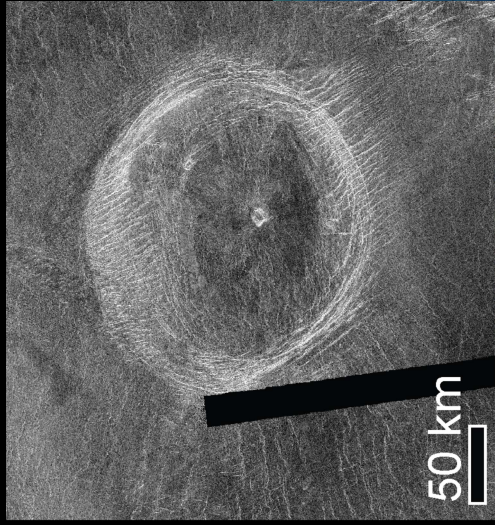
- ✓ Outer rim
- ✓ Inner depression



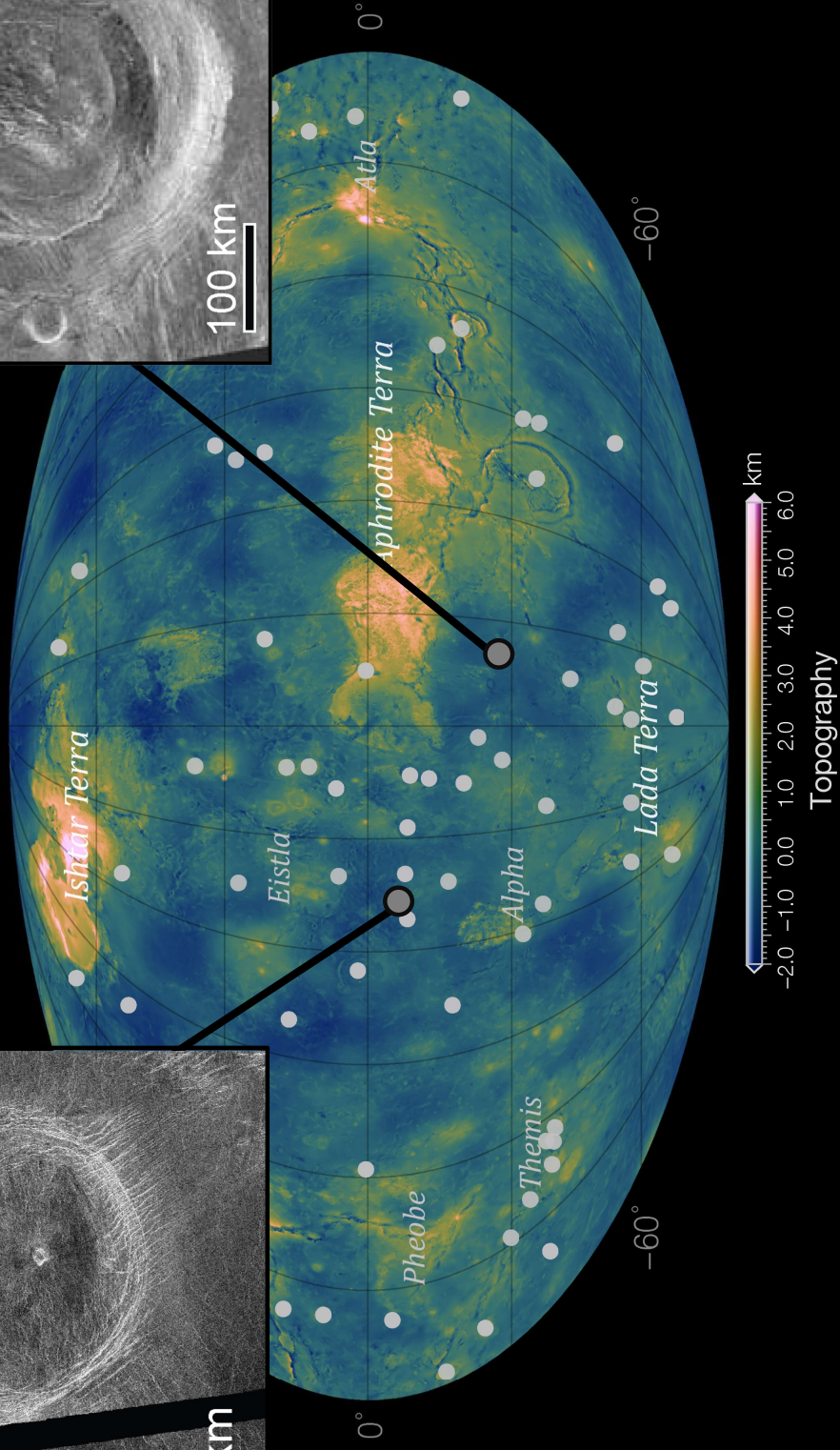
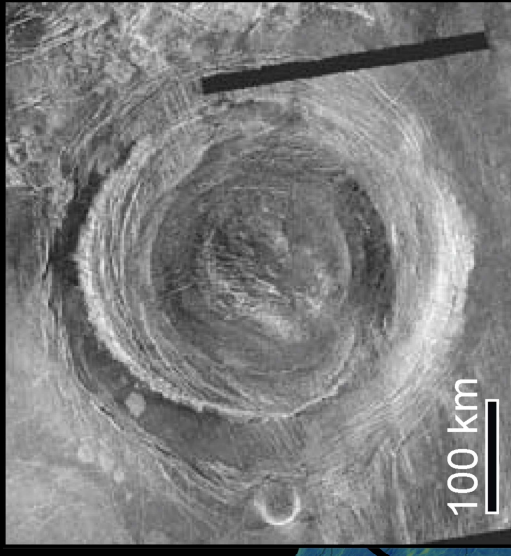
Coronae classification



Thouris corona (6.5°S, 12.9°E)



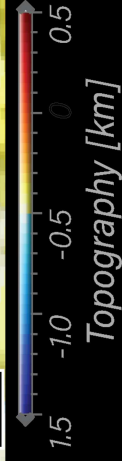
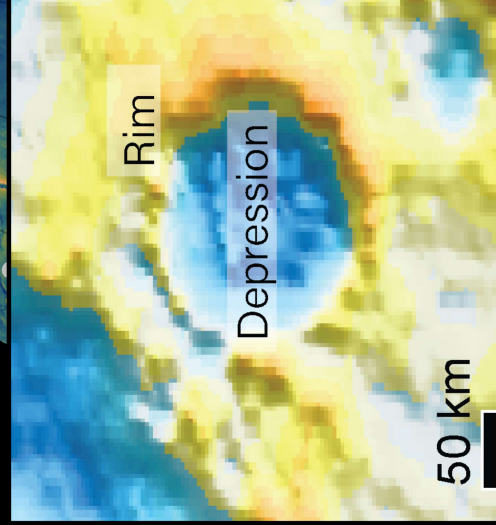
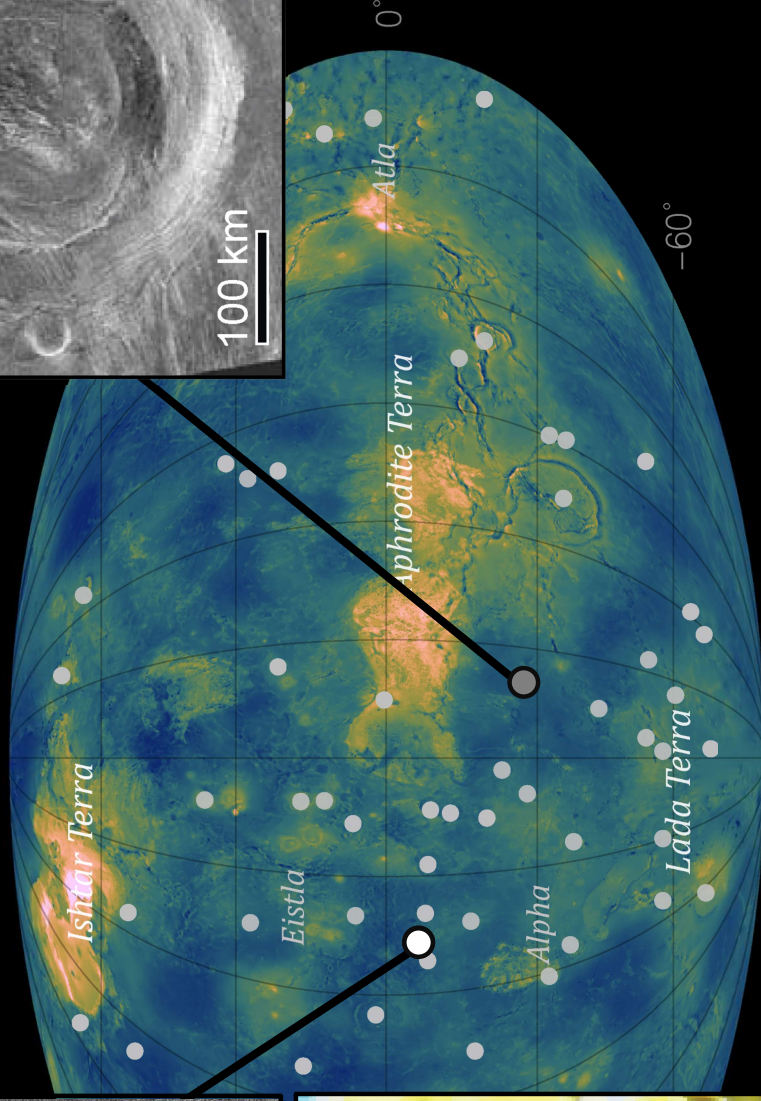
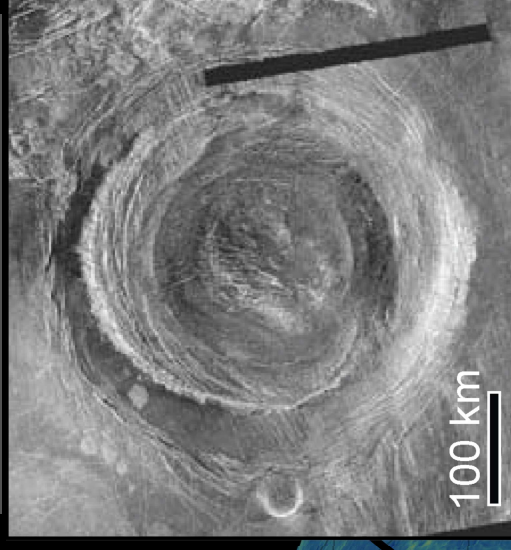
Aramaiti corona (25.5°S, 82.0°E)



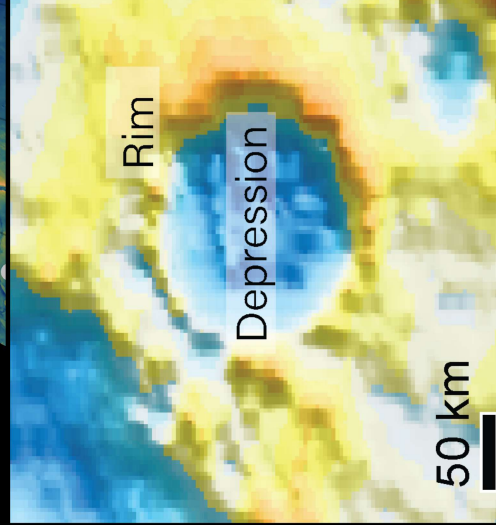
Thouris corona (6.5°S, 12.9°E)



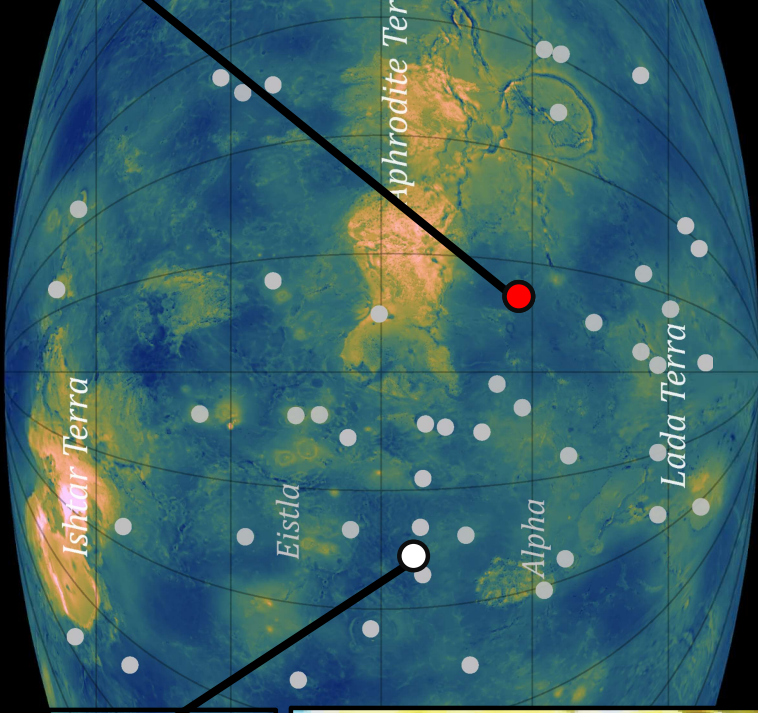
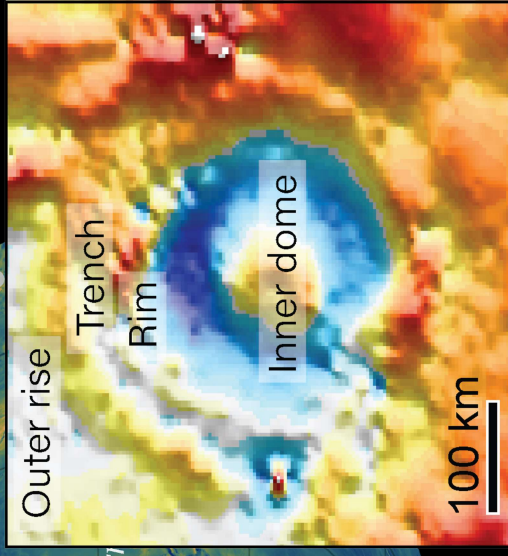
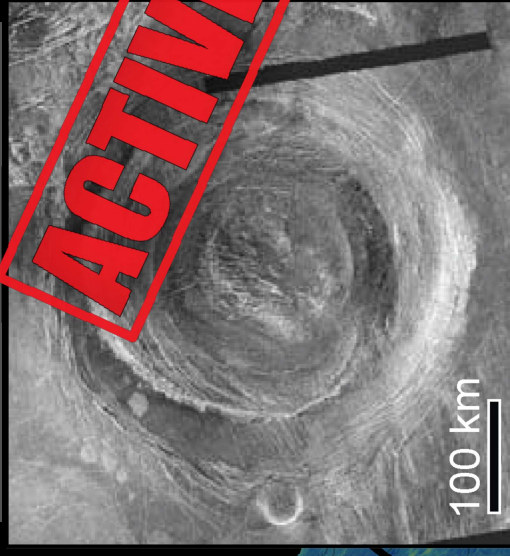
Aramaiti corona (25.5°S, 82.0°E)



Thouris corona (6.5°S, 12.9°E)



Aramaiti corona (25.5°S, 82.0°E)

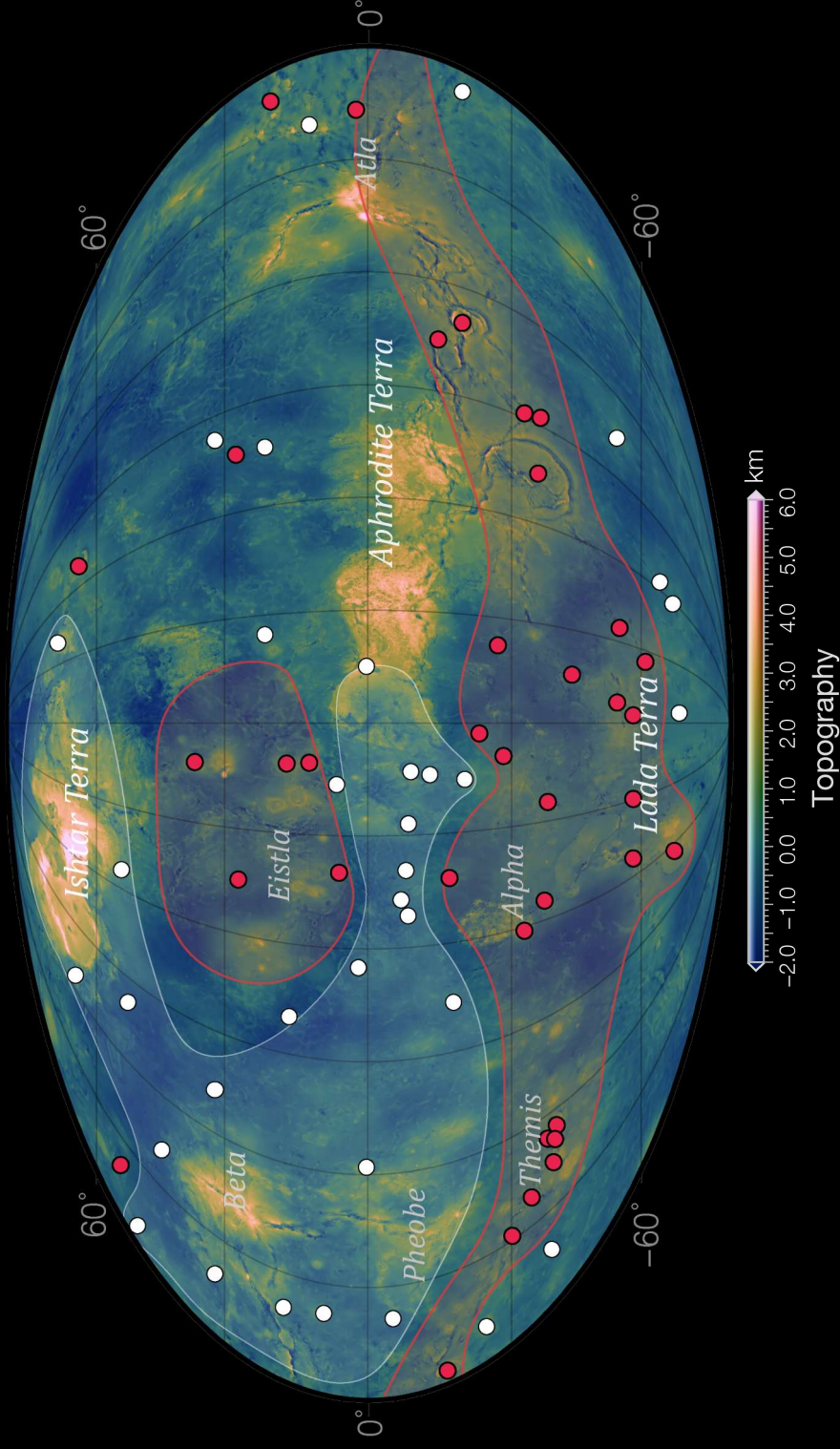


Coronae classification

Database **freely available**:

Gülcher et al., 'Venus coronae classification',
Zenodo, DOI:10.5281/zenodo.3608692

● **Ongoing activity** ● **Inactive**

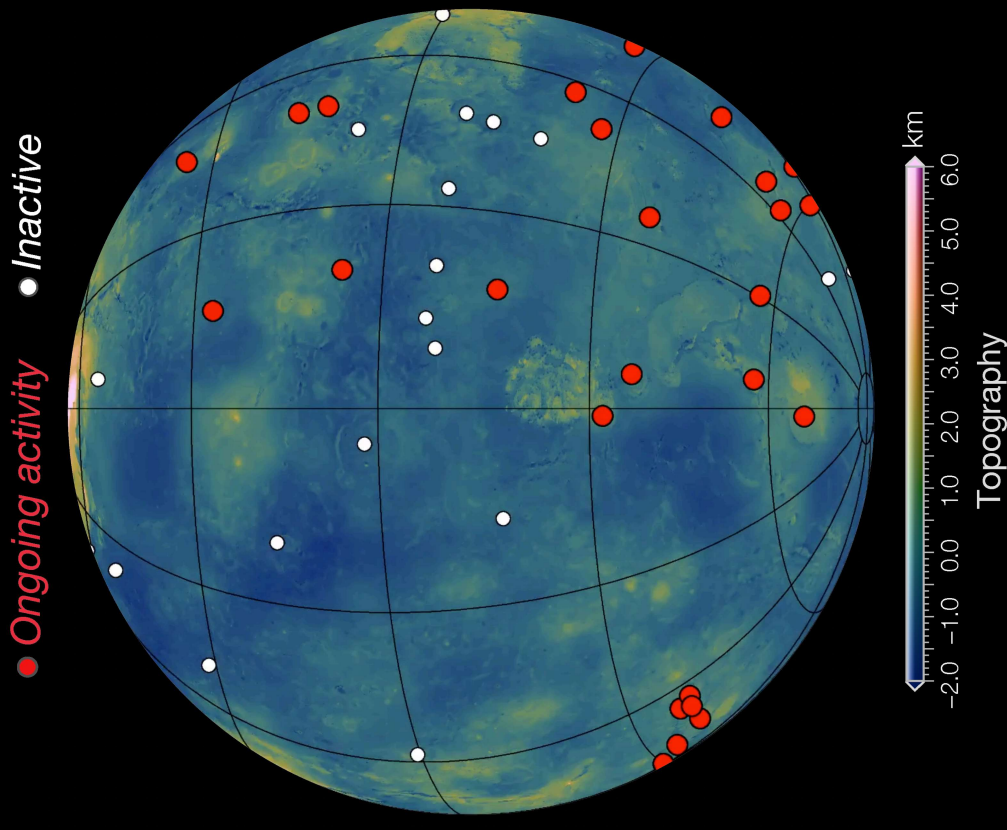


Summary of our work

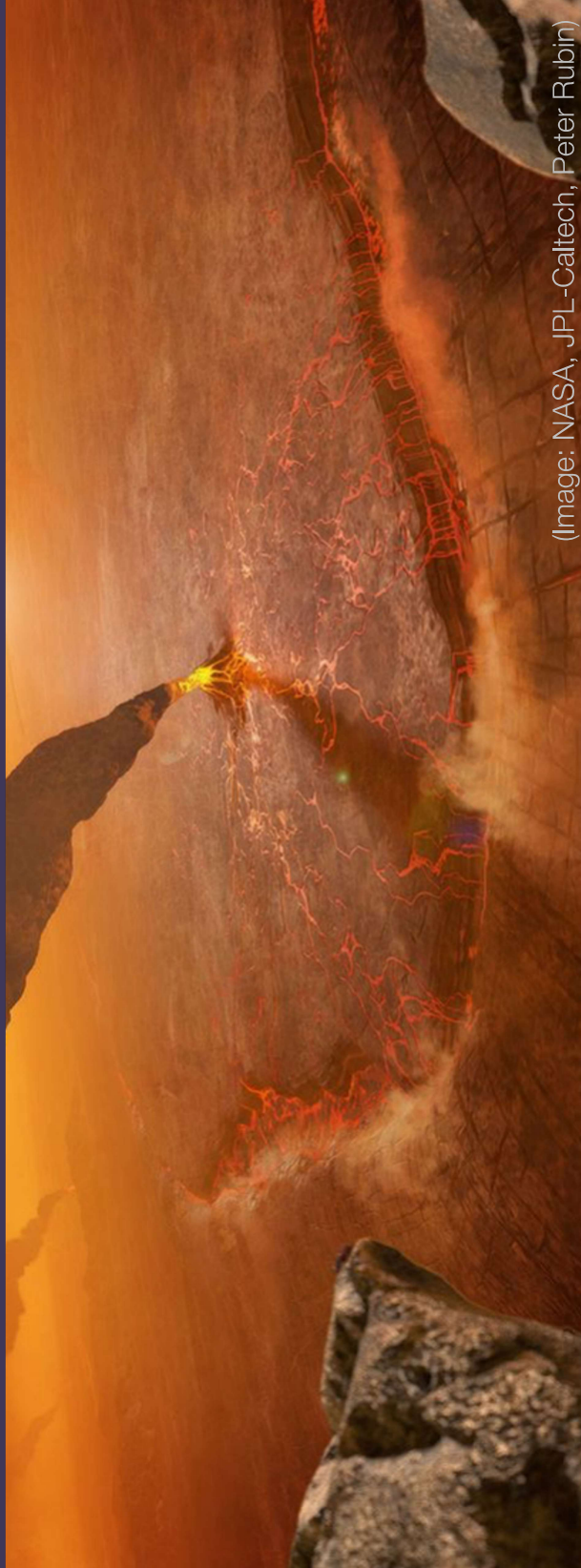
Outlook:

- Verify results with **other datasets** (gravity; geology/SAR imagery; heat flow; ..) *but limited in available datasets and resolution*
- Integrate distribution with other proposed “active” magmatic/volcanic structures
- How does **distribution** of surface activity reflect **circulation patterns** in the deep interior?

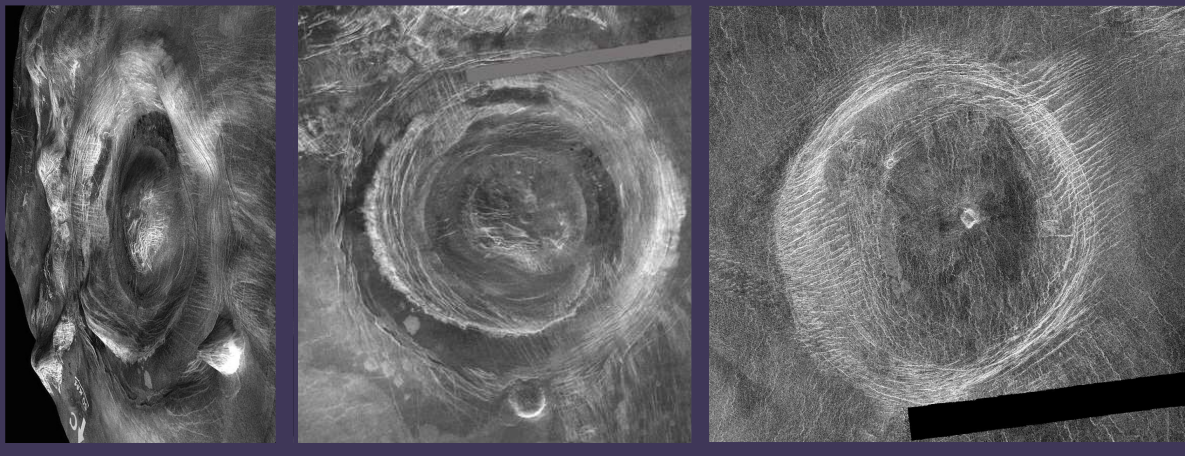
Need for new and higher-resolution geological/geophysical data!



- The formation of **coronae** may be key for unraveling Venus' interior **geodynamic processes**, and their present-day activity
- Global analysis revealed **37** coronae that bear testimony of **ongoing plume activity** underneath
- Supporting the hypothesis of a **“Hotspot” Venus** 



(Image: NASA, JPL-Caltech, Peter Rubin)



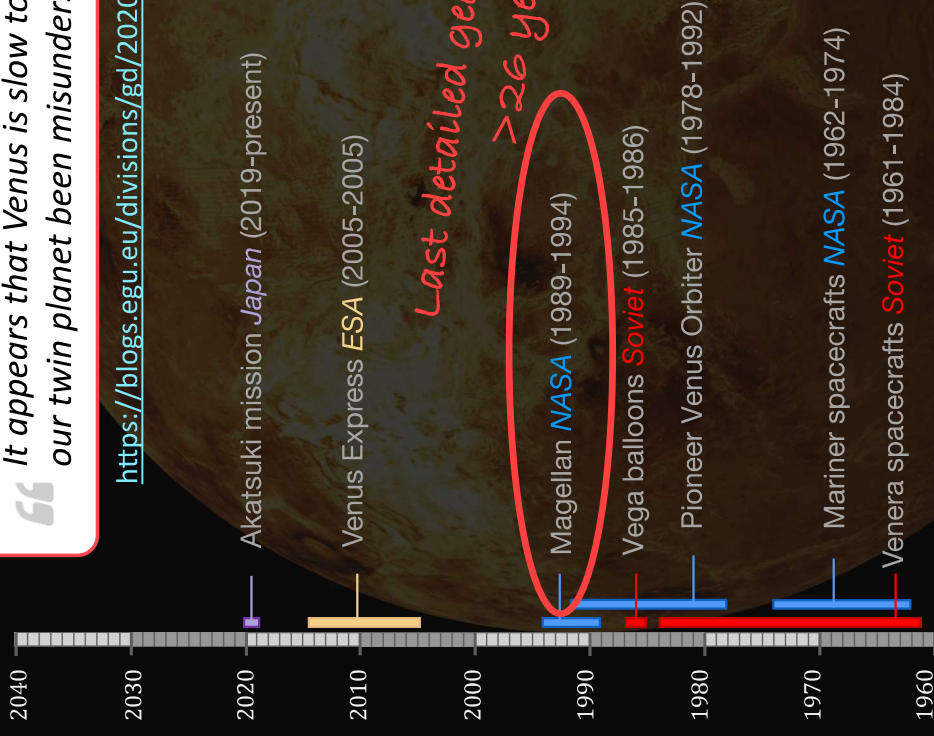
Key questions

- Did Venus have a past epoch of habitability?
- What are the physical mechanisms responsible for Venus' geologic activity?
- How have Earth and Venus diverged down such different geological paths?

“ It is clear that there are tremendous gaps in our knowledge that can only be fulfilled with new investigations of our sister world

“ It appears that Venus is slow to reveal its secrets. Has our twin planet been misunderstood all this time?

<https://blogs.egu.eu/divisions/gd/2020/09/23/venus-science-today/>



References

- Turcotte, D. L. An episodic hypothesis for Venusian tectonics. *J. Geophys. Res.* 98, 17061–17068 (1993)
- Strom, R. G., Schaber, G. G. & Dawson, D. D. The global resurfacing of Venus. *J. Geophys. Res.* 99, 10899–10926 (1994)
- Herrick, R. R. & Rumpf, M. E. Postimpact modification by volcanic or tectonic processes as the rule, not the exception, for Venusian craters. *J. Geophys. Res.* 116, E02004 (2011)
- Bjornes, E. E et al. Equilibrium resurfacing of Venus: results from new Monte Carlo modeling and implications for Venus surface histories. *Icarus* 217, 451–461 (2012)
- Smrekar, S. E. et al. Recent hotspot volcanism on Venus from VIRTIS emissivity data. *Science* 328, 605–608 (2010)
- O'Rourke, J. G. & Smrekar, S. Signatures of lithospheric flexure and elevated heat flow in stereo topography at coronae on Venus. *J. Geophys. Res.* 123, 369–389 (2018)
- Campbell, B. A. et al., Pyroclastic flow deposits on Venus as indicators of renewed magmatic activity. *J. Geoph. Res.*, 122, 1580–1596 (2017)
- Cordier, D. et al. The Physical Origin of the Venus Low Atmosphere Chemical Gradient, *The Astrophysical Journal*, 880 (2019)
- Filiberto, J., et al. Present-day volcanism on Venus as evidenced from weathering rates of olivine. *Sci. Adv.* 6, eaax7445 (2020)
- Sandwell, D.T. & Schubert, G. Evidence for retrograde lithospheric subduction on Venus. *Science* 257, 766–770 (1992)
- Stofan, E. R. et al., Corona structures on Venus: models of origin. *J. Geophys. Res.* 96, 20933–20946 (1991)
- Davaille, A. et al., Experimental and observational evidence for plume-induced subduction on Venus. *Nat. Geosci.* 10, 349–355 (2017)
- Smrekar, S. E. & Stofan, E. R. Corona formation and heat loss on venus by coupled upwelling and delamination. *Science* 277, 1289–1294 (1997)
- Gerya, T. V. Plume-induced crustal convection: 3D thermomechanical model and implications for the origin of novae and coronae on Venus. *Earth Planet. Sci. Lett.* 391, 183–192 (2014)
- McGovern, P. J. et al. The influence of lithospheric flexure on magma ascent at large volcanoes on Venus. *J. Geophys. Res.* 118, 2423–2437 (2013)
- Gerya, T. V. *Introduction to Numerical Geodynamic Modelling* (Cambridge Univ. Press, 2010)