Challenges of Reaching an ISO

Richard Linares

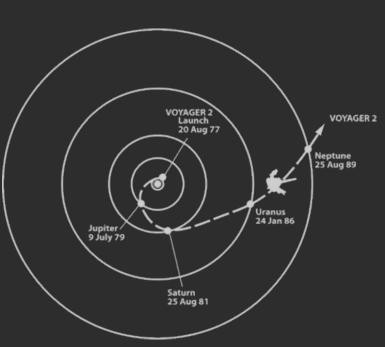
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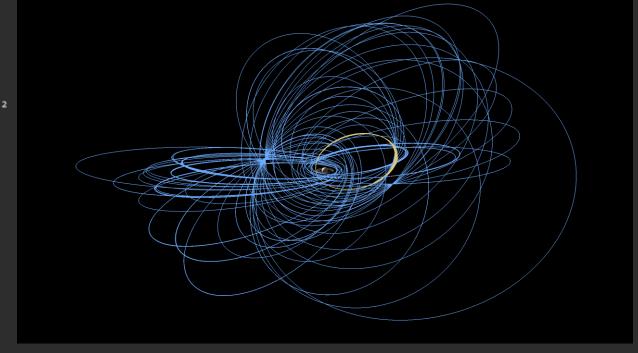


Traditional Missions, Voyager and Cassini, use Flybys



Voyager

Flybys require planning and alignment



Cassini

Challenges

- Engelhardt Suggested that there may be one ISO present within 1 AU at any given time
 - o Large Synoptic Survey Telescope will find at least 1 new ISO per year.
- ISO have high Heliocentric velocity, JPL KISS Study (Castillo-Rogez 2019)
 - o 'Oumuamua's relative velocity to the Sun was 36-50 km/s at the times of observation CFHT 2013/05/10-13-v
 - o Very short observation window (< 2 weeks)</pre>
- Orbital properties are not known in advance and can span a broad range of inclinations
 - o Due to their expected small sizes, dark albedo (from space weathering), they are hard to detect with sufficient lead time to launch a mission from Earth
- Flyby of Jupiter hard to plan for
- More intensive volatile and dust activity at long

Halley Comet 1986

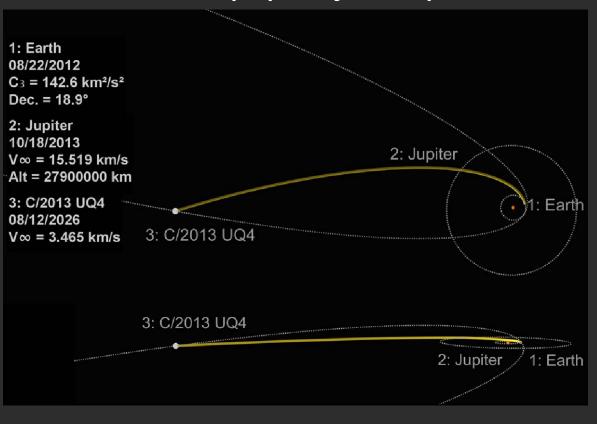


- Long-Period Comet (LPC)
 pose similar
 difficulties as ISO
- Halley comet is an example of LPC encounter 1986
 - o Encounters occurred at a relative velocity of up to 79 km/s.
 - (Julie C. Castillo-Rogez, 2019)
- Six spacecraft were sent by different space agencies:
 - NASA, ESA, Roskosmos (USSR), and JAXA (its first space mission)

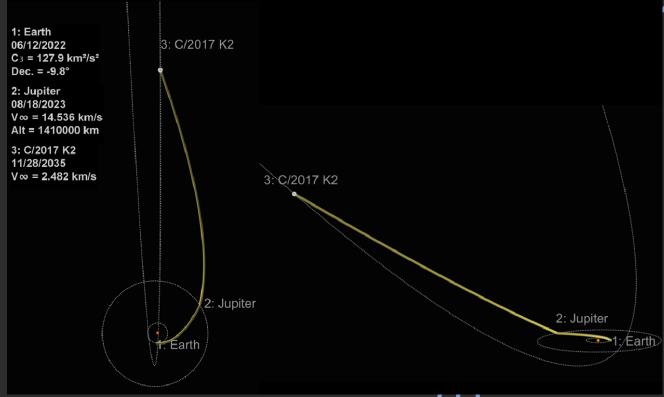


Traditional Trajectories, KISS <u>Study</u>

LPO Flyby Trajectory



LPO Fly-by Trajectory, Jupiter Flyby

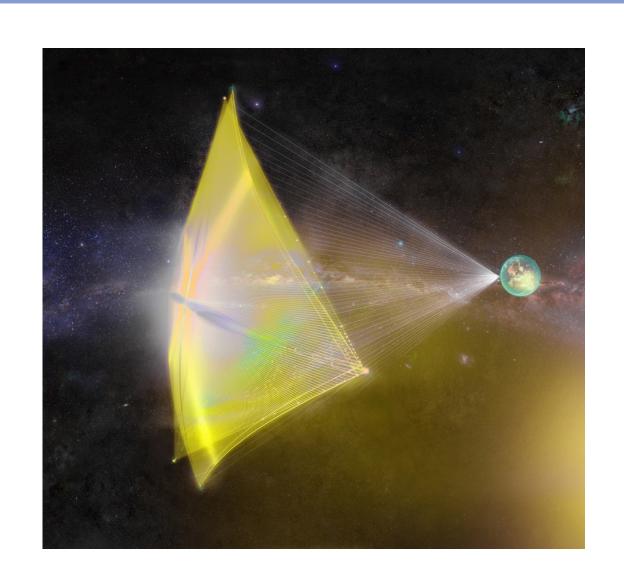


Images from Julie C. Castillo-Rogez 2019

Starshot Breakthrough Initiative

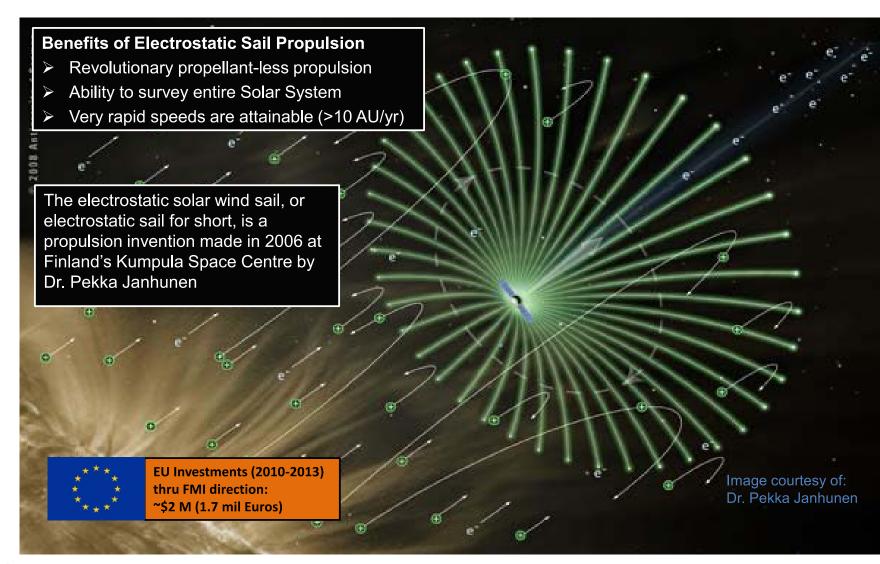
- •Using combination of Lasers and aggressive material design
 - o Ground to space laser illumination of a solar sail for DeltaV
 - Can achi eve hi gh heli ocentri c velociti es
- Material are very exotic but theoretically possible
 - oAtwater, Harry A., et al. "Materials

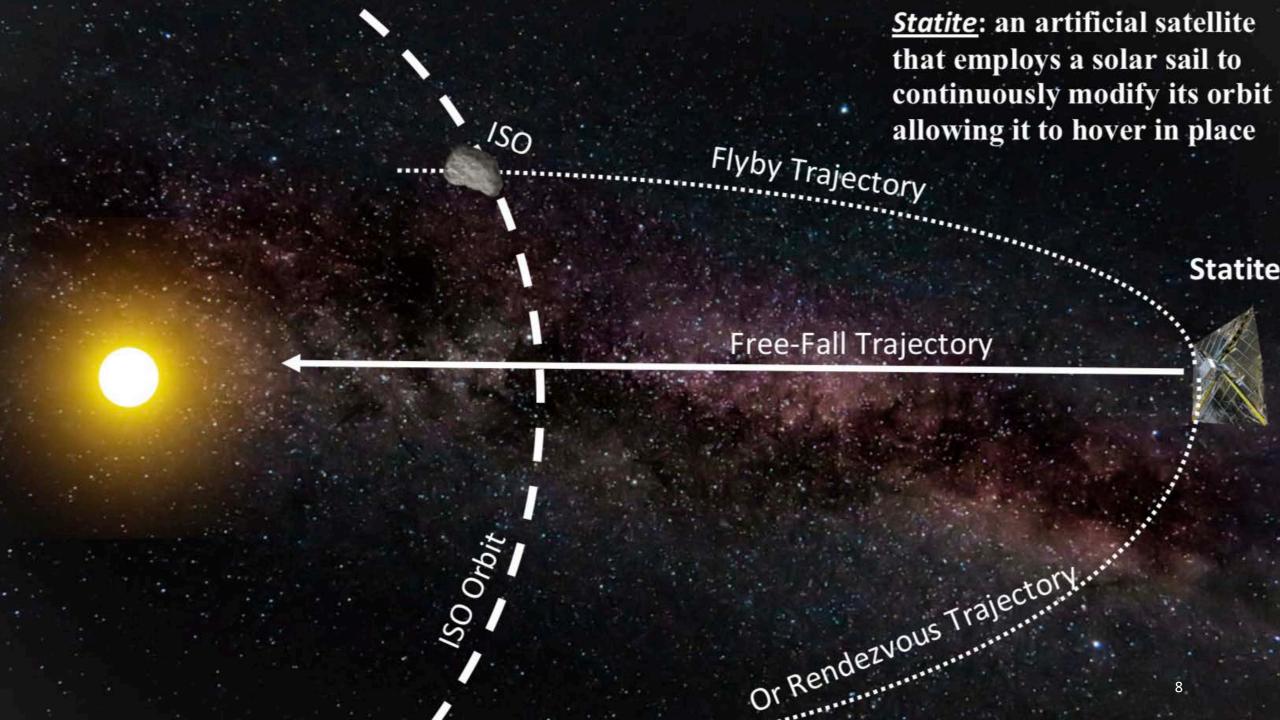




Electrostatic Sail Propulsion

- Unproven but promising
- •Heliocentric velocities of (>10 AU/yr)





Statites Concept

•Statites use solar sail technology to "cancel" out the gravitational acceleration from the Sun

$$F = \frac{GM_{Sun}m}{R^2} = 2\frac{A}{C}\frac{L}{4\pi R^2} \rightarrow \frac{m}{A} = \frac{L}{2\pi cGM_{Sun}}$$

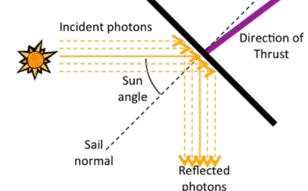
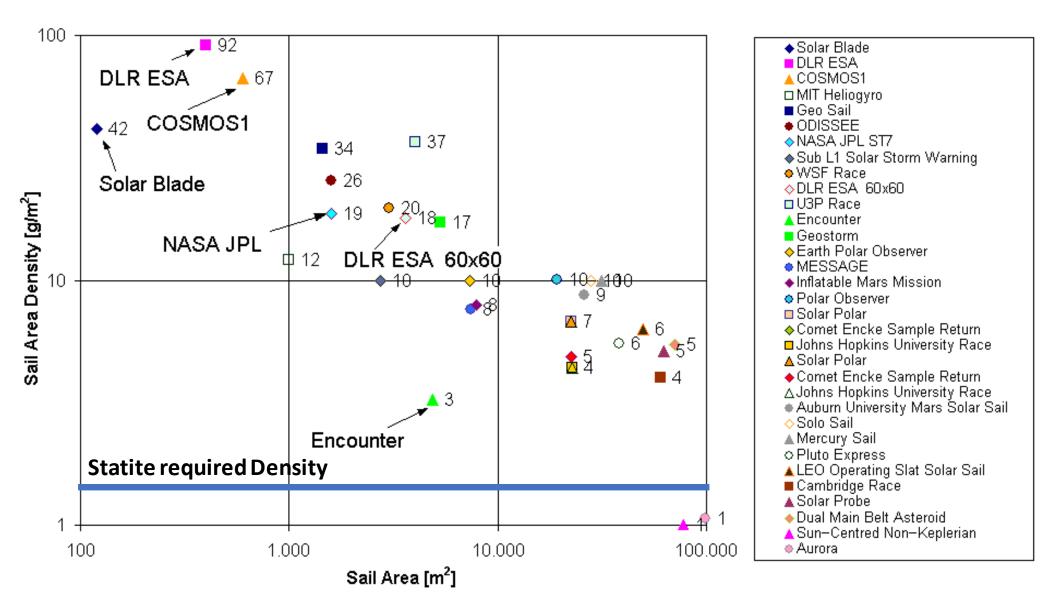


Figure 3: Solar Sail Reflection Geometry (credit www.deorbitsail.com).

- o Area-to-mass ratio is 0.625 m²/g
- oThese values are potentially achievable with exotic materials (SnS).



Herbeck, Lars, et al. "Solar sail hardware developments." European Conference on Spacecraft Structures, 2002.

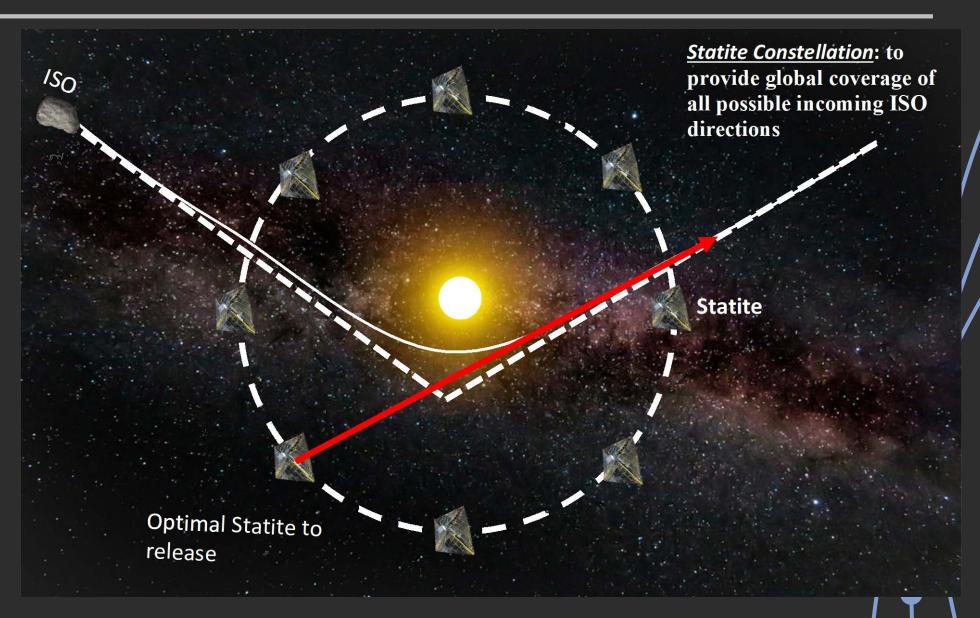
Release Trajectories

 Additionally, since the solar sail is levitating or hovering, its velocity is zero and therefore if it is Semi-major axis rel wi t

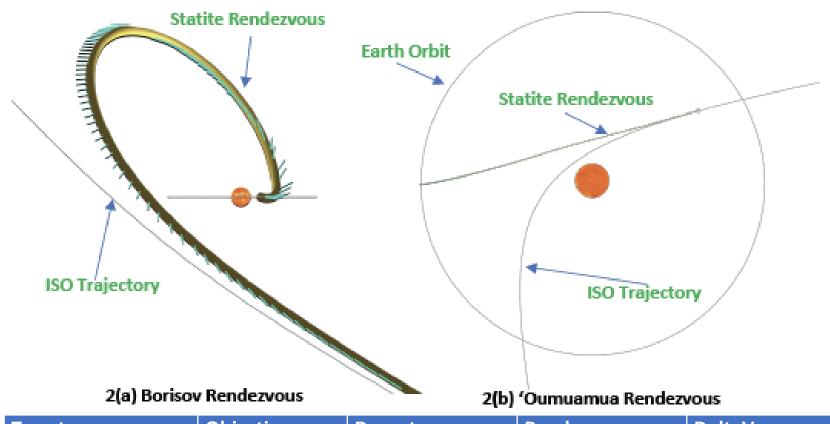
$$E = -\frac{\mu}{r} + \frac{v^2}{2} = -\frac{2\mu_{sun}}{a} \to a = \frac{r}{2} \to \frac{T}{2} = \frac{1}{2} 2\pi \sqrt{\frac{r^3}{8\mu_{sun}}} = \pi \sqrt{\frac{(1\,AU)^2}{8*4\pi^2}} = 0.1768\,\,years$$
Energy
Orbital Period
Orbital Period

- This relationship shows that a statite at 1 AU has a free-fall trajectory of about 64 days.
- This fast response time to a potential ISO can be thought of as a slingshot effect, since the solar sail is used to "store energy" that is released when

Release Trajectories



Initial Results

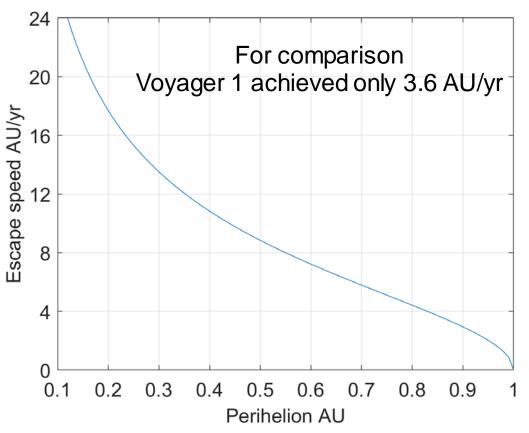


Target	Objective	Depart	Rendezvous	DeltaV
ISO1/'Oumuamua	Max. departure	June/25/2017	Oct/01/2017	104 km/s
ISO2/ Borisov	Min. flight time	Dec/29/2016	Feb/24/2024	145 km/s

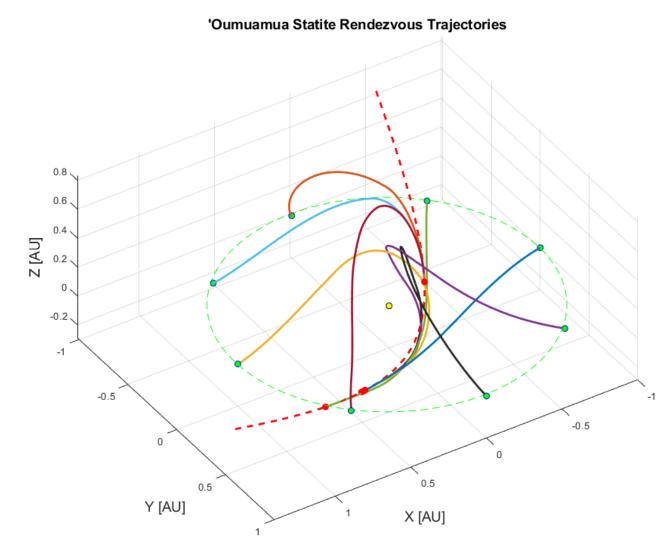


ISO Coverage

Figure shows "back of the envelope" calculation



Solar System Escape trajectories for Statites.





Concl usi on

- •The existence of ISOs offer us a great scientific opportunity to develop a detail understanding of the formation and history of other star systems.
- •However, due to their high heliocentric velocities and relatively short lead time, this may be extremely difficult with current satellite propulsion systems.
- The proposed approach can respond to a newly detected ISO within 2-3 month to perform a flyby.
- The proposed approach also allows for rapid response rendezvous missions.