Nancy Grace Roman Space Telescope: Core Community and General Astrophysics Surveys

Saurabh W. Jha (Rutgers University) & Gail Zasowski (University of Utah)
NAS CAA, 6 Oct 2025

ROTAC Co-Chairs



Roman WFI Observations: A Community Driven Process



 $Figure\ credit:\ Margaret\ Carruthers,\ Landsat/Copernicus,\ IBCAO,\ USGS,\ SIO,\ NOAA,\ U.S.\ Navy,\ NGA,\ GEBCO,\ Google$

STScI Newsletter Volume 42:

More than 1,000 scientists from over 350 institutions contributed to this <u>community process</u>. Nearly 200 science pitches and white papers were submitted, describing science investigations that could be enabled with one or more of the surveys using Roman's imaging and slitless spectroscopy capabilities from 0.5 to 2.3 microns.

NANCY GRACE

ROMAN

Roman Mission & Observing Strategy Timeline





2022	NAS CAA Report on Roman Observing Time Allocation Principles released	I T			
2023	2023 White papers solicited from the community				
Core Community Survey (CCS) Definition Committees (DCs) design survey options - Galactic Bulge Time Domain Survey (GBTDS) - High Latitude Time Domain Survey (HLTDS) - High Latitude Wide Area Survey (HLWAS)					
2024 November	November ROTAC convened to evaluate CCS survey options				
2025 January	CCS DC reports & presentations released: GBTDS HLTDS HLWAS				



2025 April	ROTAC report and presentation

2025 September

Early-Definition Galactic Plane Survey (GPS) report released

2025 November (planned) Cycle 1 General Investigator (GI) Call for Proposals, due March 2026

2026 June–August First GI programs and General Astrophysics Surveys (GAS) selected

Roman WFI Observations: A Community Driven Process



Members of the Roman Observations Time Allocation Committee

Co-Chairs: Gail Zasowski (University of Utah), Saurabh W. Jha (Rutgers University)

Members: Laura Chomiuk (Michigan State University), Xiaohui Fan (University of Arizona), Ryan Hickox (Dartmouth College), Dan Huber (University of Hawaii, Manoa), Eamonn Kerins (University of Manchester), Chip Kobulnicky (University of Wyoming), Tod Lauer (NOIRLab), Masao Sako (University of Pennsylvania), Alice Shapley (University of California, Los Angeles), Denise Stephens (Brigham Young University), David Weinberg (Ohio State University), Ben Williams (University of Washington)

Galactic Bulge Time Domain Survey

Chairs: Jessie Christiansen (NExScI/Caltech), Daniel Huber (UH/USyd)

Members: Annalisa Calamida (STScI), Jessica Lu (UC Berkeley), Eduardo Martin (IAC), Benjamin Montet (UNSW), Kris Pardo (USC), Matthew Penny (LSU), Hans Walter Rix (MPIA), Jennifer Sobeck (IPAC)

High Latitude Wide Area Survey

Chairs: Ryan Hickox (Dartmouth), Risa Wechsler (Stanford/KIPAC)

Members: Micaela Bagley (UT Austin), Keith Bechtol (UWisc), Michael Blanton (NYU), Chris Hirata (OSU), Elisabeth Krause (U of Arizona), Nikhil Padmanabhan (Yale), Ismael Tereno (IA), Anja von der Linden (Stony Brook), David Weinberg (OSU), L. Y. Aaron Yung (STScI)

High Latitude Time Domain Survey

Chairs: Brad Cenko (NASA/GSFC), Masao Sako (UPenn)

Members: Alessandra Corsi (Texas Tech), Michael Fasnaugh (Texas Tech), Sebastian Gomez (STScI), Rebekah Hounsell (UMBC/GSFC), Takashi Moriya (NAOJ), Gordon Richards (Drexel), Russell Ryan (STScI), Schuyler Van Dyk (IPAC/Caltech), Ashley Villar (CfA/Harvard)

Solar System Liaisons

Susan Benecchi (Planetary Science Institute), Rosemary Pike (CFA)

Roman Space Telescope WFI's Prime Mission

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SPACE TELESCOPE

- High Latitude Wide Area Survey (520 days)
- High Latitude Time Domain Survey (180 days)
- Galactic Bulge Time Domain Survey (438 days)
- General Astrophysics Surveys (389 days)
 - 25.5% of prime mission science operations time, including the early-definition Galactic Plane Survey + any remaining margin (e.g., unused calibration time)

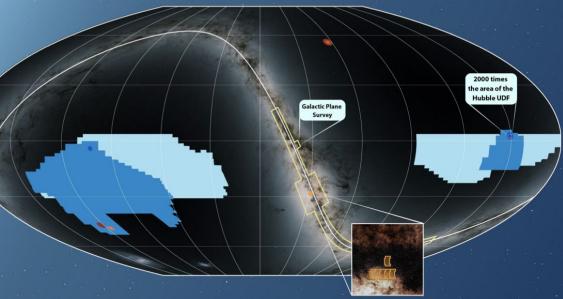
Science for All ROMAN'S EXTENSIVE SURVEYS **High Latitude Time Domain Survey** High Latitude Wide Area Survey (9 5-10 days ■ Deep [□] + ■ Wide N [Ö]^{×5} S [Ö]^{×5}■ △ ■ Medium 👸 + 📥 Ultra Deep 👸 ×3 ■ Deep N 👸 *5 S 👸 *5 ■ 🛆 520 days throughout the primary 5-year mission Mapping the universe in high resolution 180 days executed primarily in the middle of the mission Roman Effective Area SCA01 Roman Effective Area Evolution and Energy Z Y J H W F087 F106 F129 F158 F146 o.s area [m²] 0.5 1.5 600 million galaxies with shape measurements over 12% of the sky 100,000 transient light curves 2000 times the area of the 1.5 0.50 0.75 1.00 1.25 1.50 1.75 2.00 2.25 **Hubble UDF** Wavelenath (microns) **Galactic Plane Galactic Bulge Time Domain Survey** Five fields + Galactic Center ~5 days [6] 438 days over high and low cadence bulge seasons 100,000 exoplanet discoveries

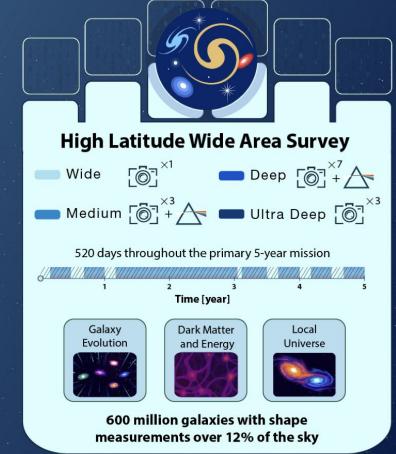
ROMAN SPACE TELESCOPE

Science for All

ROMAN'S EXTENSIVE SURVEYS

Precise measurements of cosmic expansion and the growth of cosmic structure





Via weak gravitational lensing, galaxy clustering, baryon acoustic oscillations, and redshift-space distortions

Multi-band imaging and spectroscopy enabling studies from solar system objects to galaxy and quasar evolution out to the epoch of reionization SPACE TELESCOPE

High Latitude Wide Area Survey (HLWAS)

ROMAN

SPACE TELESCOPE

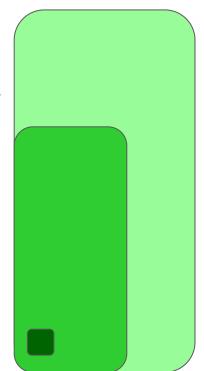
NANCY GRACE

Three tiers: **Medium**, **Wide**, and **Deep**

Medium (2400 deg²) tier includes three bands of imaging (YJH) and grism spectroscopy. This allows detailed calibration of weak lensing measurements and a wide-area redshift survey, plus many astrophysical applications.

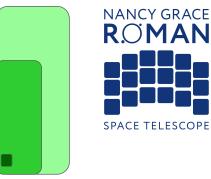
Wide (2700 deg²) tier provides additional imaging area in a single band (H) for both cosmology and general astrophysics. Numerous community inputs advocated for a very wide-field survey.

Deep (20 deg²) tier provides essential calibration of the **Medium** and **Wide** tiers. Includes multiple bands (WZYJHFK) and grism spectra, plus a **5 deg²** region with **ultra-deep** YJH imaging.



High Latitude Wide Area Survey (HLWAS)

	In-Guide Survey Breakdown							
	Name	Area & Filters	Exposure times	Depth	Survey yields			
$\frac{1}{\sqrt{2}}$	Wide	2702 deg ² H	2x3x107 s (H)	26.2 (H)	n _{eff} = 26.7 arcmin ⁻² N _{shape} = 260M			
\	Medium	2415 deg ² YJH + Grism	2x3x107 s (YJH) 4x2x190 s (Grism)	26.5 (Y) / 26.4 (JH) 1.5x10 ⁻¹⁶ (Grism)	$\begin{aligned} & n_{\text{eff}} = 41.3 \text{ arcmin}^{-2} \\ & N_{\text{shape}} = 360 \text{M} \\ & N_{\text{spec}} = 19.1 \text{M (total)} \\ & 9.3 \text{M (z=1-2)} \\ & 1.7 \text{M (z=2-3)} \\ & n_{\text{spec}} = 9.8 \times 10^{-4} h^3 \text{Mpc}^{-3} \end{aligned}$			
\	Deep	19.2 deg ² WZYJHFK + Grism	+5x3x295 s (WZYJHFK) +36x2x190 s (Grism)	27.7 (ZY) / 27.6 (J) / 27.5 (H) / 27.0 (F) / 25.9 (K) / 28.3 (W) 5.8x10 ⁻¹⁷ (Grism)	(for calibration)			
\	Ultra Deep	5 deg ² YJH	+10x3x295 s (YJH)	28.2 (YJ) / 28.1 (H)	(for calibration)			



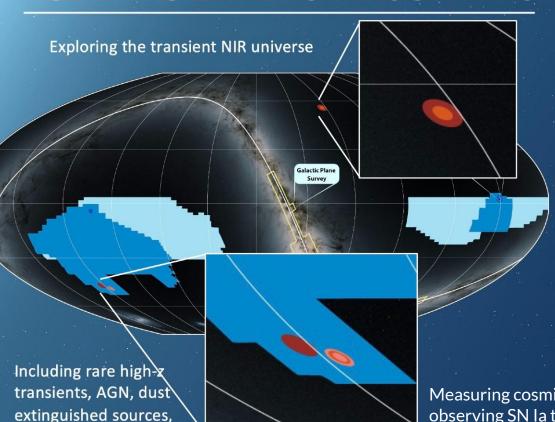
Survey yields:

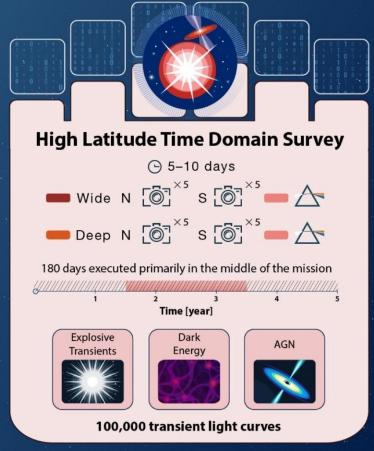
- > 600M galaxies with shape measurements
- > 19M galaxies with spectroscopic redshifts

Science for All

and cool, red objects

ROMAN'S EXTENSIVE SURVEYS





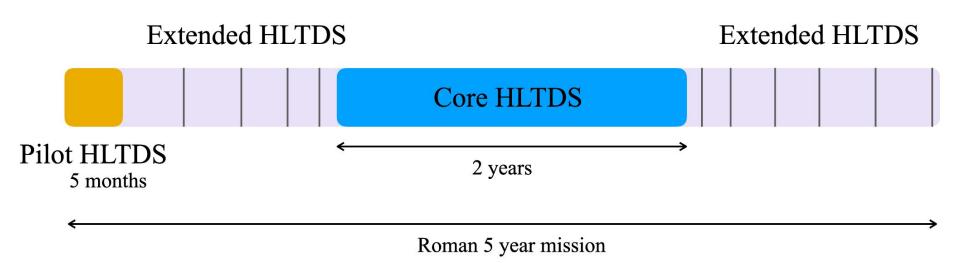
Measuring cosmic expansion and dark energy by observing SN Ia to z > 2.5, with color information and spectra at 5-day cadence over two years



High Latitude Time Domain Survey (HLTDS)

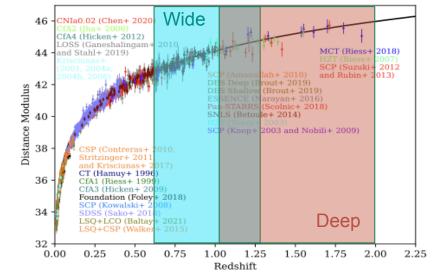
NANCY GRACE ROMAN SPACE TELESCOPE

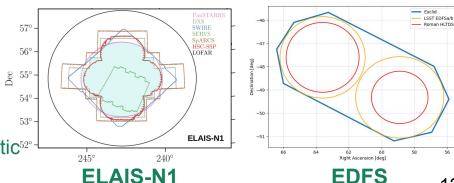
3 Components:



High Latitude Time Domain Survey (HLTDS)

- Core Component (158 days)
 - Imaging (78%)
 - SN Ia: z ~ 0.9 (Wide), z ~ 1.7 (Deep)
 - 10-day interlaced cadence
 - RZJ / RYH Wide
 - ZYH / ZJF Deep
 - Prism spectroscopy (22%)
 - 5-day cadence in Deep Imaging
 - 900 sec, 3600 sec exposures
- Pilot Component (15 days)
 - 8 visits to all imaging and prism fields in first ~5 months of Roman.
 - Templates & other Core Component preparation, and early science.
- Extended Component (7 days)
 - 8 visits to Deep Imaging fields before and after Core Component.
 - Long-term monitoring of high-redshift, exotic^{52°} transients.

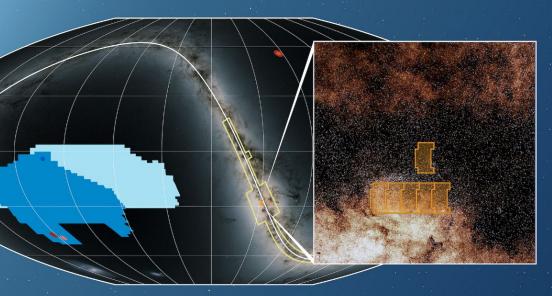


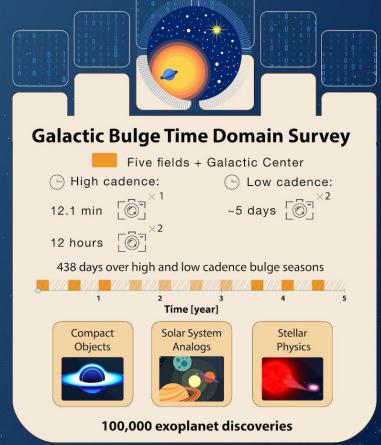


Science for All

ROMAN'S EXTENSIVE SURVEYS

High plus low cadence seasons maximize scientific yield





Enabling microlensing detections from scales of solar system analogues to stellar mass black holes, and detection > 100,000 transiting exoplanets

Advancing stellar and Galactic astrophysics via red giant asteroseismology, monitoring Sgr A*, and studies of compact objects

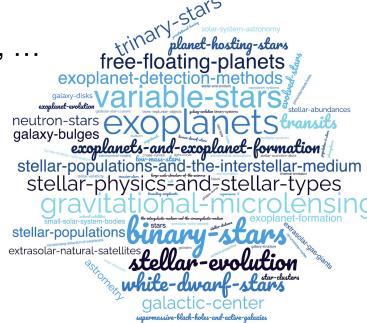


Galactic Bulge Time Domain Survey (GBTDS)

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ROMAN

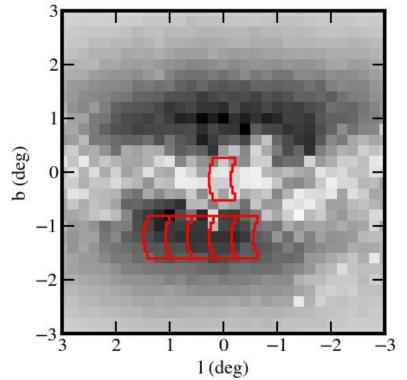
SPACE TELESCOPE

- High-cadence seasons: exoplanet microlensing (Level 1 requirement), transiting exoplanets, asteroseismology, Galactic Center populations, stellar flares, ...
- Low-cadence seasons: black hole microlensing, long-period variables
- Photometric snapshots: stellar colors, extinction
- Grism snapshots: stellar properties, kinematics, binary stars



Galactic Bulge Time Domain Survey (GBTDS)

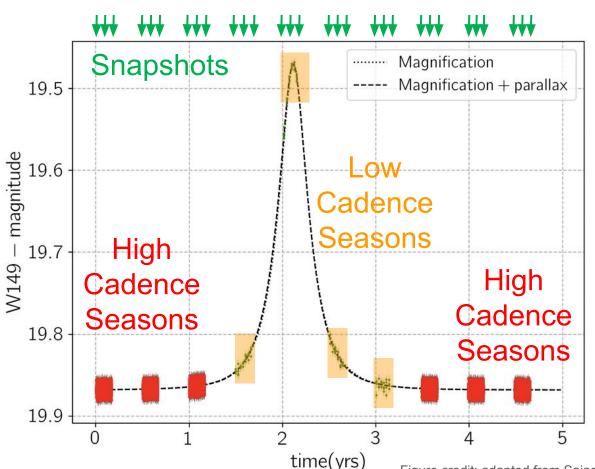




4 Components

- 6 × 70.5 day **high-cadence seasons**
 - first & last 3 bulge seasons
 - 5 fields + Galactic Center
 - 12.1 min cadence
 - wide F146 filter, 67 sec exptime
 - F087 and F213 at 12 hr cadence
- 4 low-cadence seasons
 - 5 day cadence (to accommodate HLTDS scheduling constraints)
- 30 photometric snapshots (5 filters)
- 30 grism snapshots

Galactic Bulge Time Domain Survey (GBTDS)





Total Survey
Duration: 438 days

~423 days

~6 days

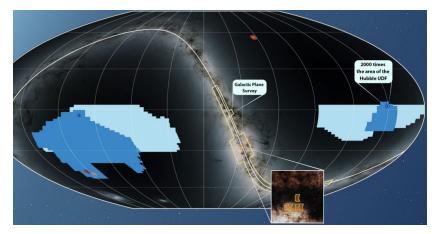
~9 days

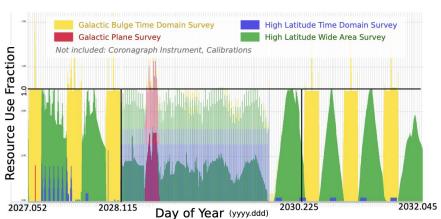
2 days during first season reserved for GPS

General Astrophysics Surveys

- ROTAC saw its role as stewards of the highly-valued GAS time (in accord with CAA report)
- 389 days available for General Astrophysics Surveys (25.5% of prime mission science time)
 - □ Including the ~30 day early-definition Galactic Plane Survey







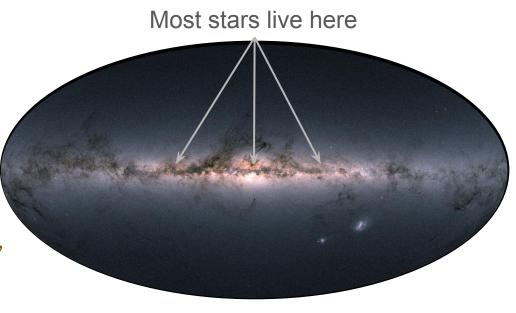
GAS #1: Galactic Plane Survey

<u>Galaxy studies</u>: Milky Way structure, Galactic dynamics, far side of the Galaxy

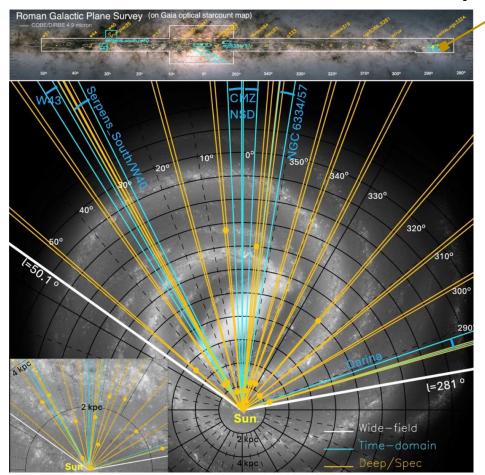
Stars & stellar physics: Low-mass stars, star clusters, brown dwarfs, eruptive stars, variable stars of all ages, compact objects

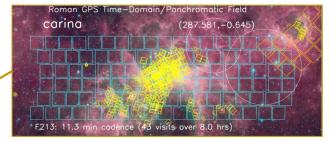
<u>Stellar/planetary systems</u>: Star formation, microlensing, solar system demographics, free-floating planets

<u>Interstellar medium</u>: dust, extinction curves, stellar nurseries



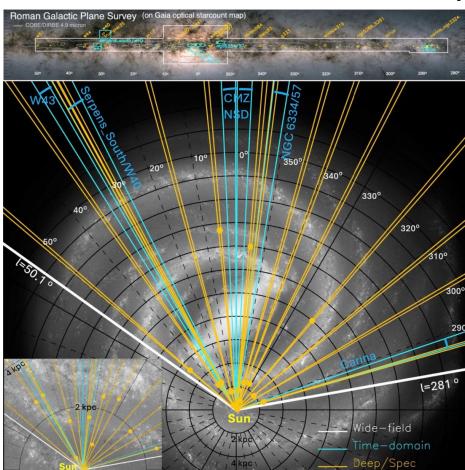
GAS #1: Galactic Plane Survey





- Wide-field: 3-4 filter imaging of a contiguous 691 deg² of the disk and bar/bulge
- Time-domain: Repeat
 observations in 1+ filters with
 cadences from 11 min to several
 weeks, in 19 deg² of carefully
 selected fields
- Deep-field/spectra: Deep,
 7-filter imaging and grism+prism
 spectra in 4 deg²

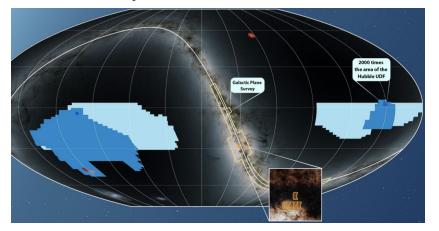
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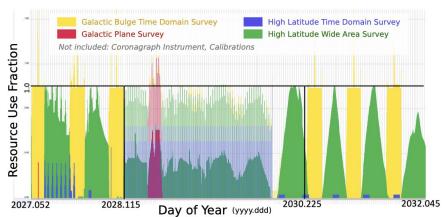


	Hours	Percent	Filters/Cadence*
Wide-Field Science	540.65	77.2%	
Disk	382.18	54.6%	F129,F158,F184,F213
Disk Carina	44.43	6.3%	F129,F158,F184,F213
Bulge_Bpos	52.27	7.5%	F129,F158,F213
Bulge_Bneg	51.94	7.4%	F129,F158,F213
Serpens_South	9.83	1.4%	F106, F129 ,F158,F184, F213
Time-Domain Science	129.52	18.5%	
TDS_Carina	11.89	1.7%	F213 (min)+F062.F087,F106,F129
TDS_NGC6334_6357	11.89	1.7%	F213 (min)+F062,F087,F106,F129
TDS Galactic Center Q4	39.04	5.6%	F213 (min/hrs/wks)+F129 (wks) + F062,F087,F106,F129
		12.50.00000	F213 (min/hrs/wks)+F129 (wks) +
TDS_Galactic_Center_Q1	39.04	5.6%	F062,F087,F106,F129
TDS_Serpens_South_W40	14.86	2.1%	F213 (hrs) (+F106,F129,F158,F184,F213 from Wide Field Science)
			F184 (min)+
TDS_W43	12.80	1.8%	F062,F087,F106,F129,F213
Deep/Spectroscopic Science	30.80	4.4%	
Deep_NGC_3324_Carina	1.79		All filters deep, grism,prism
Deep_Acrux	1.79		All filters deep, grism,prism
Deep_NGC_5269_5281	1.79		All filters deep, grism,prism
Deep_Window_319.5-0.2	1.79		All filters deep, grism,prism
Deep_G333	1.79		All filters deep, grism,prism
Deep_ASCC_85	1.79		All filters deep, grism,prism
Deep_Teutsch_84	1.79		All filters deep, grism,prism
Deep_NGC_6357_Lobster	1.79		All filters deep, grism,prism
Deep_Window_355.0-0.3	1.79		All filters deep, grism,prism
Deep_VVV-CL001_UKS_1	1.79		All filters deep, grism,prism
Deep_M17_Omega	1.79		All filters deep, grism,prism
Deep_Trumpler_35	1.79	0.3%	All filters deep, grism,prism
Deep_W40	5.74		Very Deep F129,F158,F213, prism
Deep_W44	1.79		All filters deep, grism,prism
Deep_W51	1.79	0.3%	All filters deep, grism,prism
TOTAL	700.97	100.1%	

General Astrophysics Surveys

- ROTAC saw its role as stewards of the highly-valued GAS time (in accord with CAA report)
- 389 days available for General Astrophysics Surveys (25.5% of prime mission science time)
 - Including the ~30 day early-definition Galactic Plane Survey
- Up to 30 GAS selected over several proposal calls (first proposals anticipated due March 2026)
 - o GAS proposals likely in Cycles 1, 3, and 4; selection will be weighted towards later calls, due to high demand for early-mission time, as well as learning from observatory performance and data
- The community is encouraged to organize and design large GAS proposals with broad science goals, including those complementary to or building on the CCS
- We anticipate GAS to comprise a larger fraction of the schedule in an extended mission, but the community should evaluate this as the science returns from early observations become known

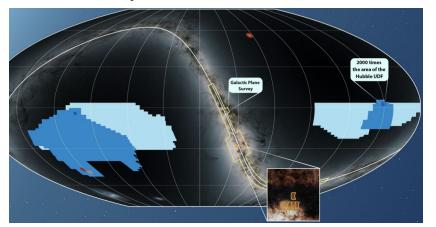






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Roman Cycle 1:

- Funding for
 - Analysis of Roman data
 - Analysis of relevant ground-based data
 - Theory studies
 - Limited number of new GAS



Roman Prime Mission: Summary

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The community-driven CCS+GAS design allows Roman to address the most compelling questions in exoplanet demographics & cosmology while enabling a broad range of general astrophysics investigations, featuring fully public data. The science portfolio of NASA's next flagship observatory will be altogether transformative.

