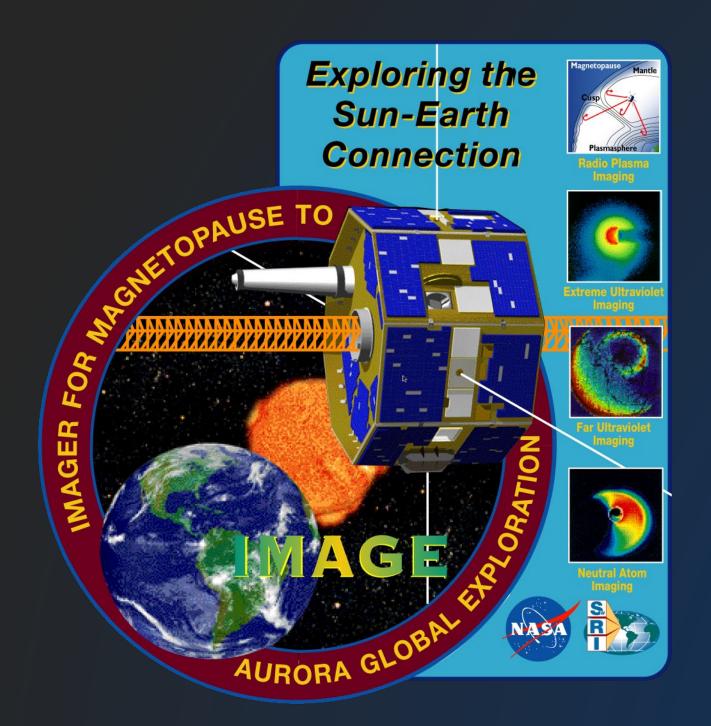
Mapping the Magnetosphere with Energetic Atoms

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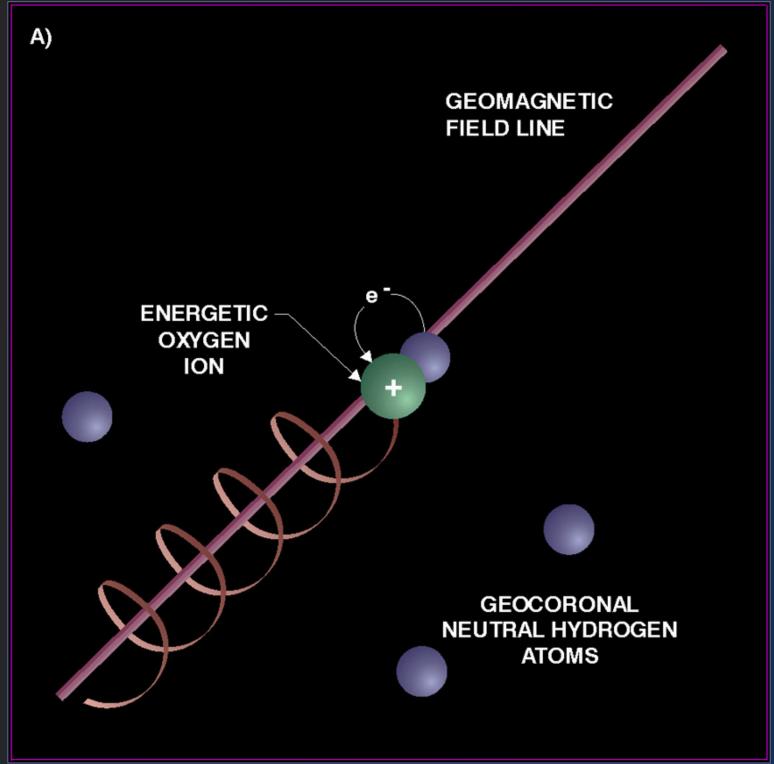


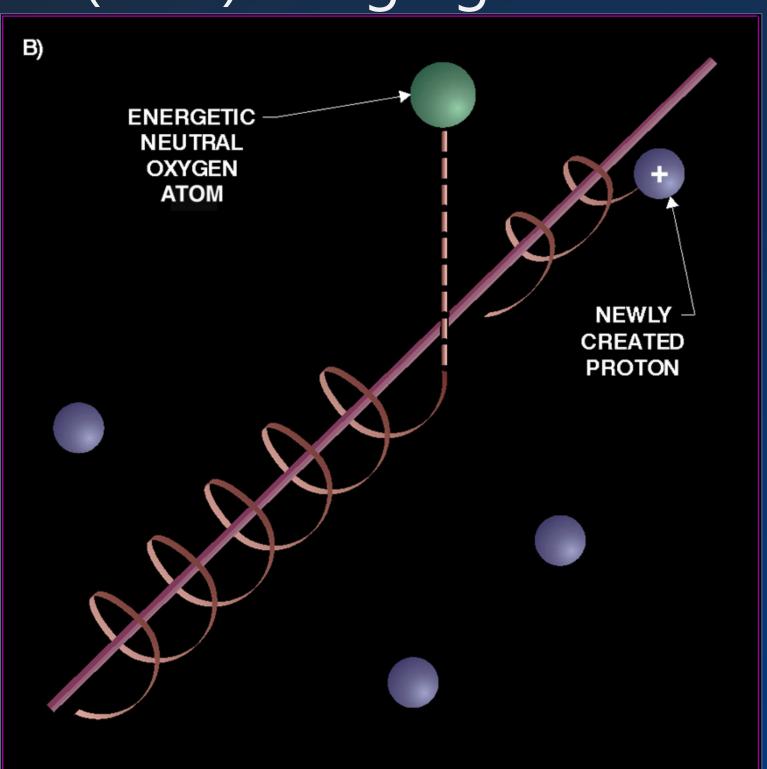






What is Energetic Neutral Atom (ENA) Imaging?

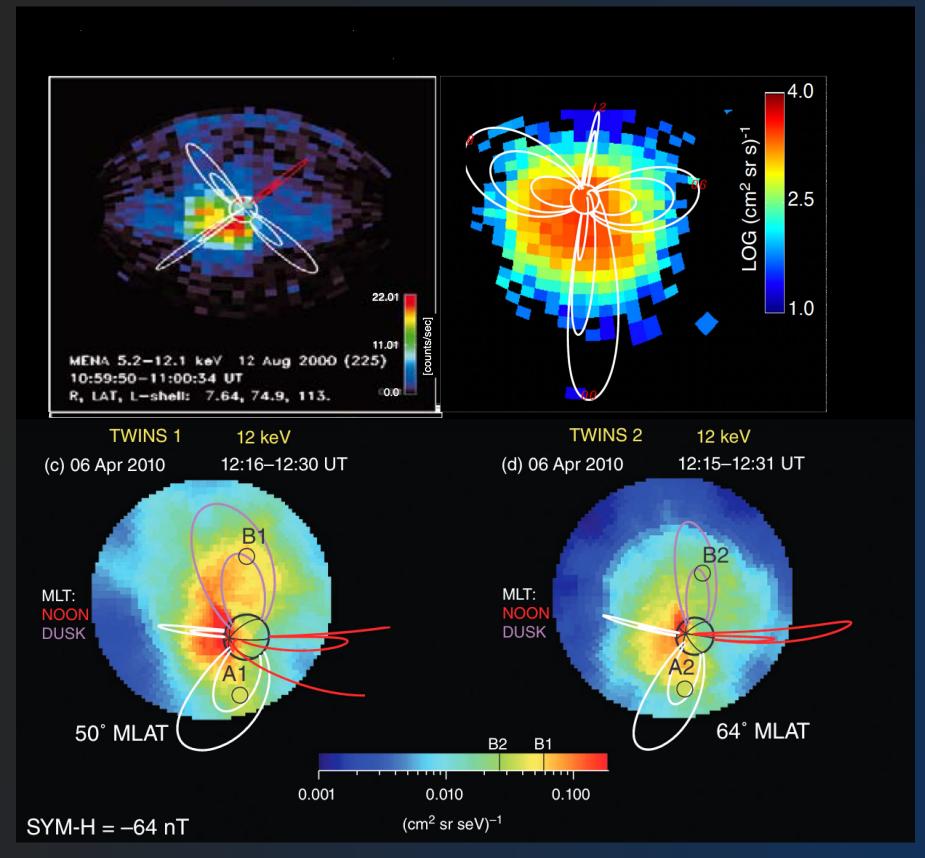




ENA imaging takes advantage of charge-exchange collisions



What are the benefits of ENA Imaging?

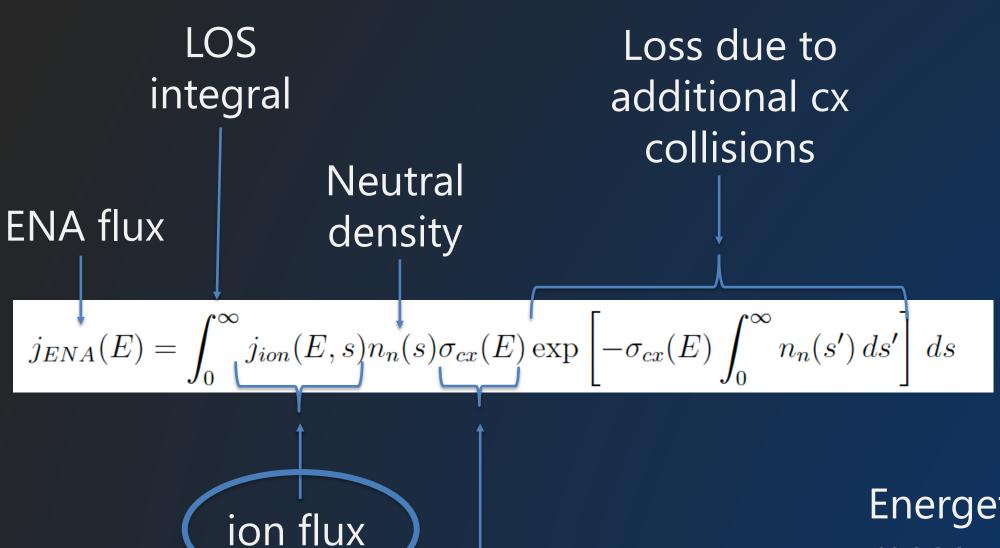


Energetic neutral atom (ENA) imaging is a remote measurement that can provide a global view of the ions in the magnetosphere

Skymap images of ENA flux from MENA (upper left) (Fok et al, 2003), HENA (upper right) (Brandt et al, 2002), and both TWINS (bottom) (Maggiolo et al, 2021)

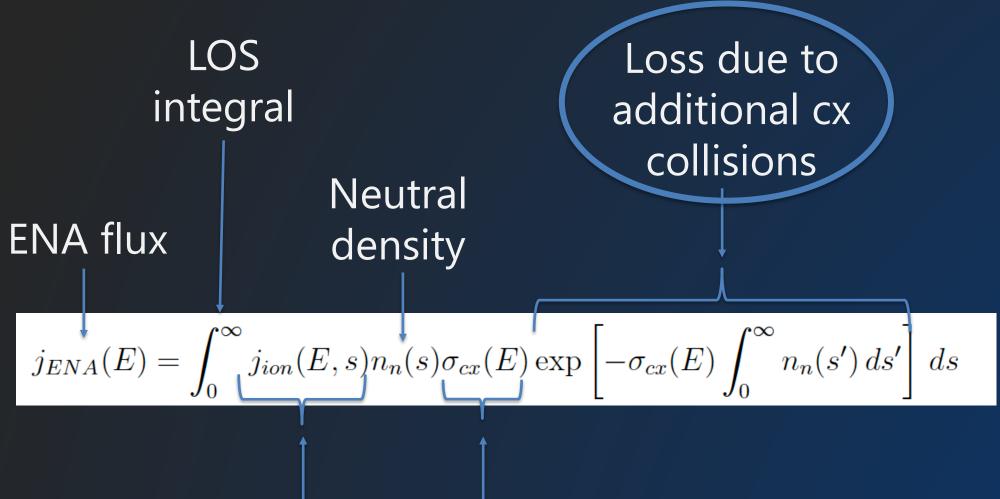


What do ENA instruments measure?



Charge exchange cross section

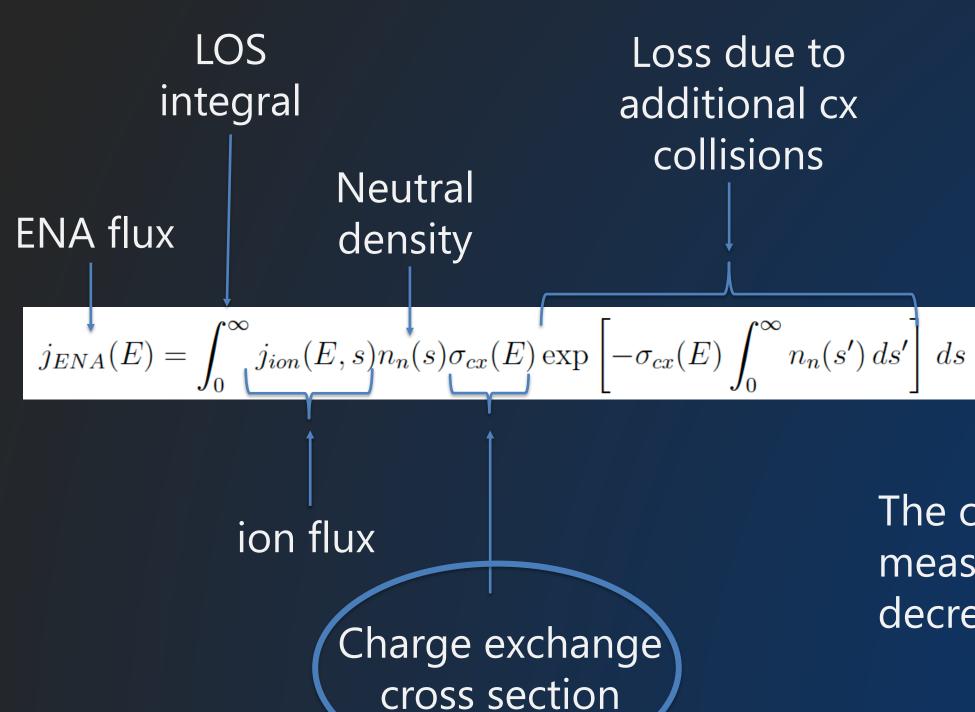
Energetic neutral atom (ENA)
measurements help us get
information about the ion population



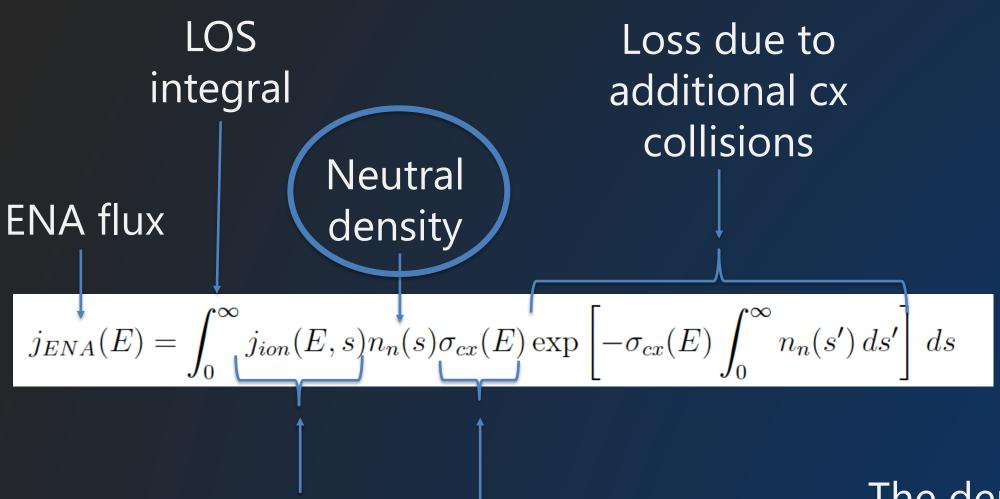
ion flux

Charge exchange cross section

Additional charge exchange collisions need to be considered for low altitude emissions (LAE)



The charge exchange cross section, measured in laboratory experiments, decreases with increasing energy

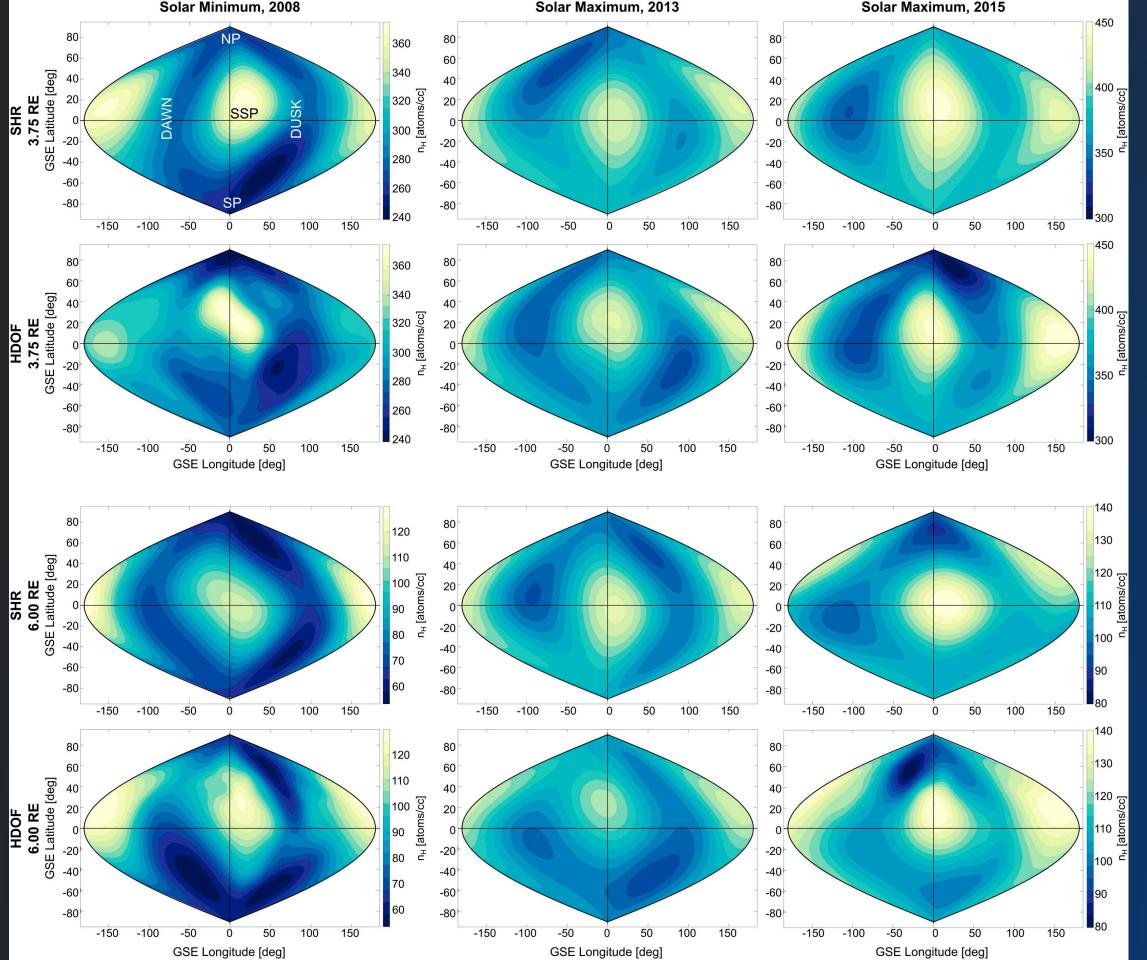


Charge exchange cross section

ion flux

The density of neutrals in the exosphere decreases with distance from the Earth, but has been shown to be dynamic

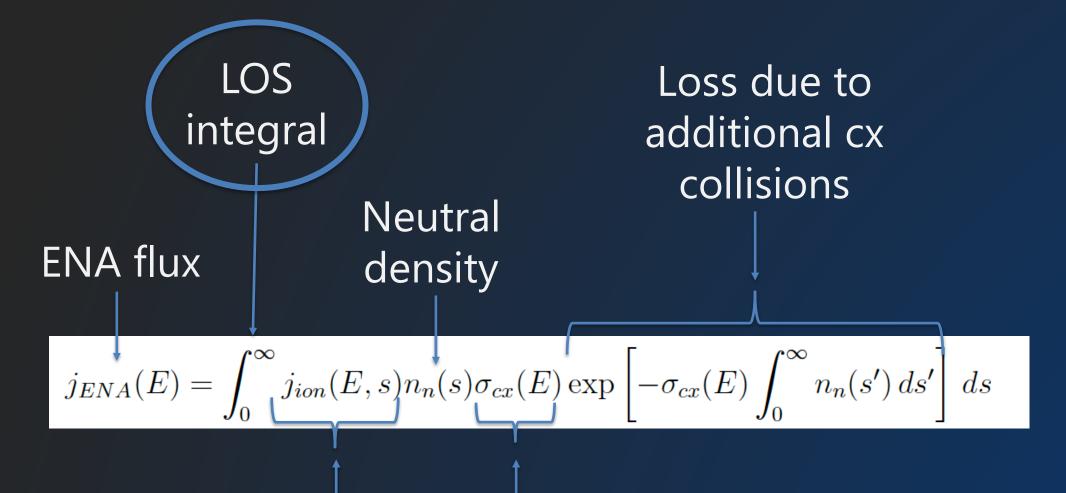




Lyman-α measurements of Hydrogen geocorona using TWINS show spatial and temporal variations

Zoennchen JH et al. (2024) doi: 10.3389/fspas.2024.1409744



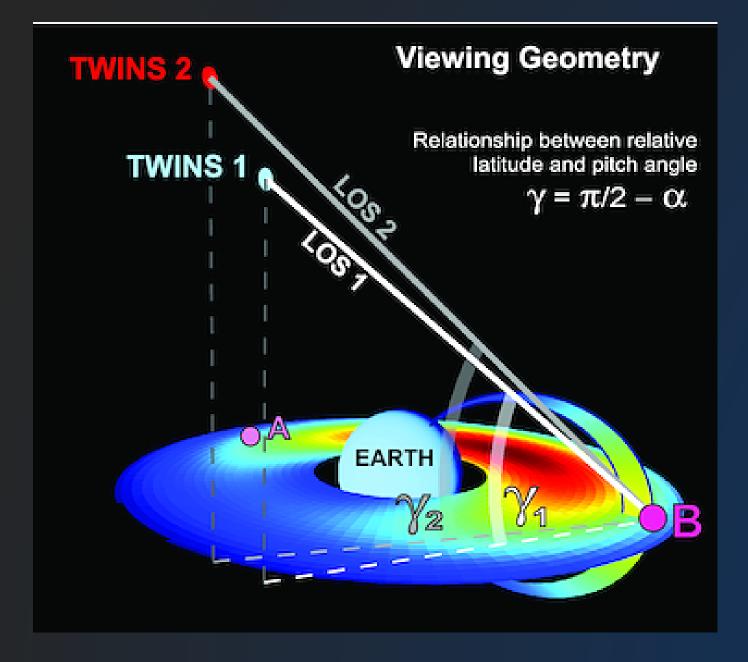


ion flux

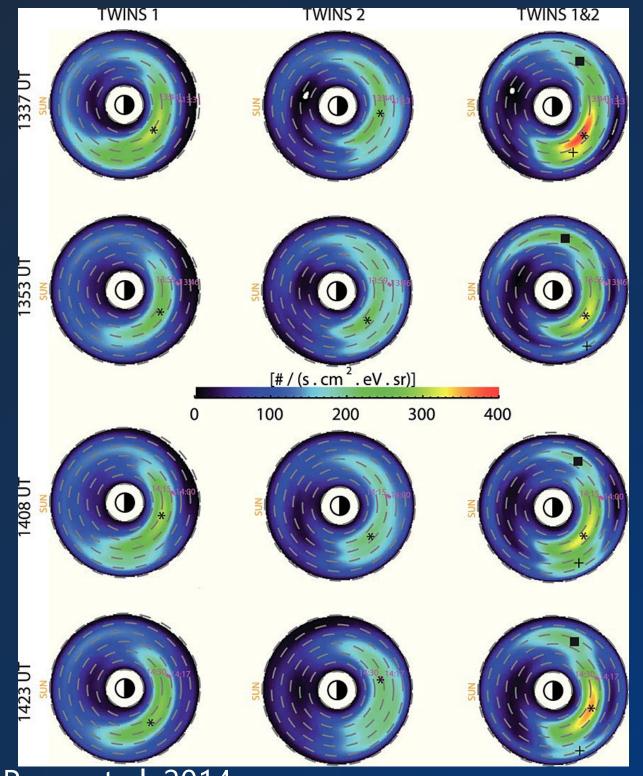
Charge exchange cross section

The measurement combines everything along the line of sight (LOS)

TWINS studies utilized dual vantage points



Goldstein et al, 2012 https://doi.org/10.1029/2012JA017804

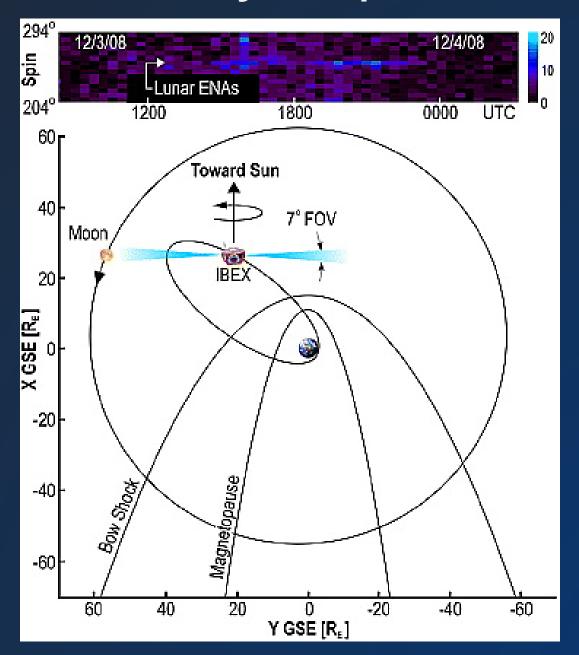


Perez et al, 2014 https://doi.org/10.1002/2014JA020662

University of New Hampshire

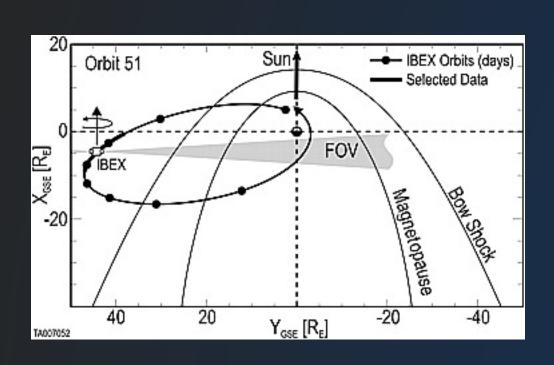
What if we put an ENA imager on the moon?

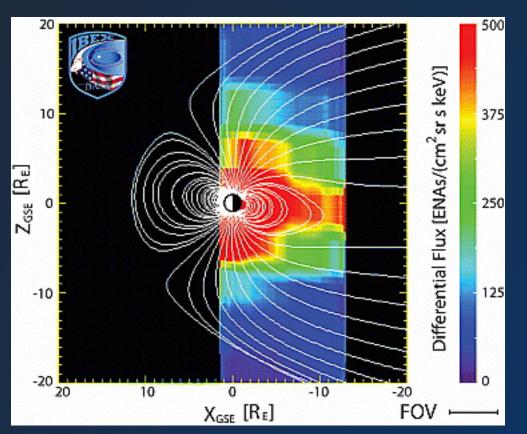
- Designed to observe ENAs from ~100 AU
- Observed ENAs of lunar origin
- Relevant to solar wind-moon interactions throughout solar system and other space plasma/small body interactions
- A consideration as background signal in moon-based measurements



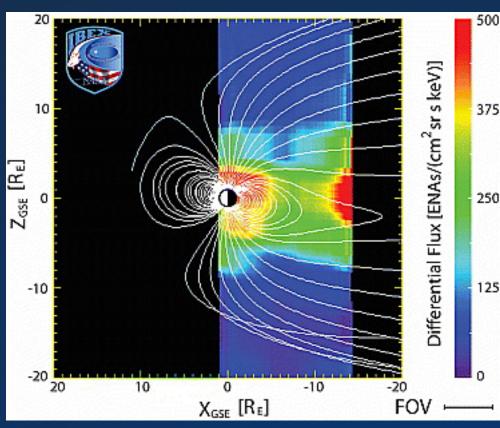
McComas et al, 2009 https://doi.org/10.1029/2009GL038794

Observations of the magnetotail/plasma sheet: Observations of downtail intensification





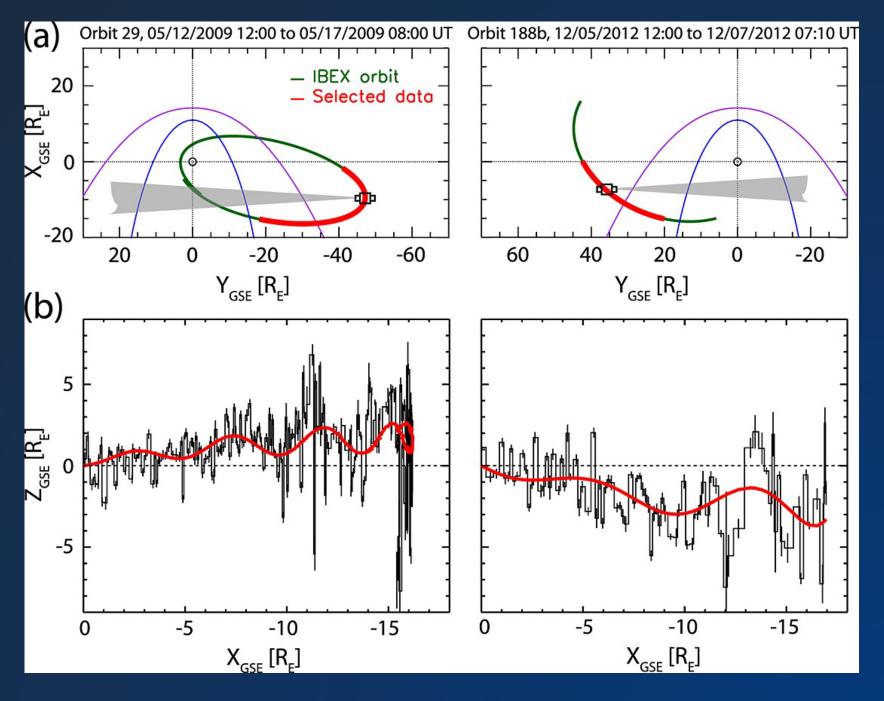
1048 UT on 5 November - 0223 UT on 7 November 2009



2121 UT on 27 October - 1340 UT on 29 October 2009

McComas et al, 2011 https://doi.org/10.1029/2010JA016138

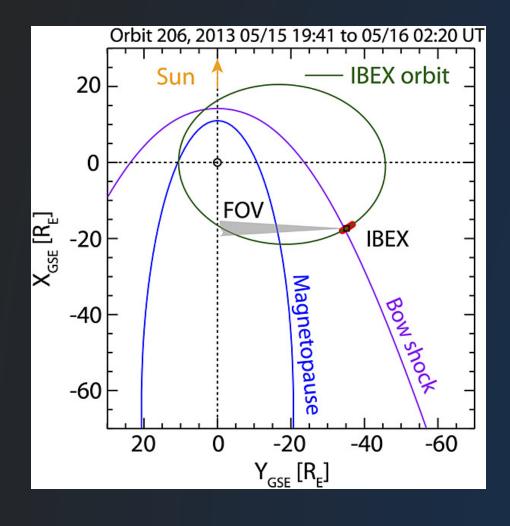
Observations of the magnetotail/plasma sheet:
Observations of plasma sheet diurnal motion and short timescale flapping

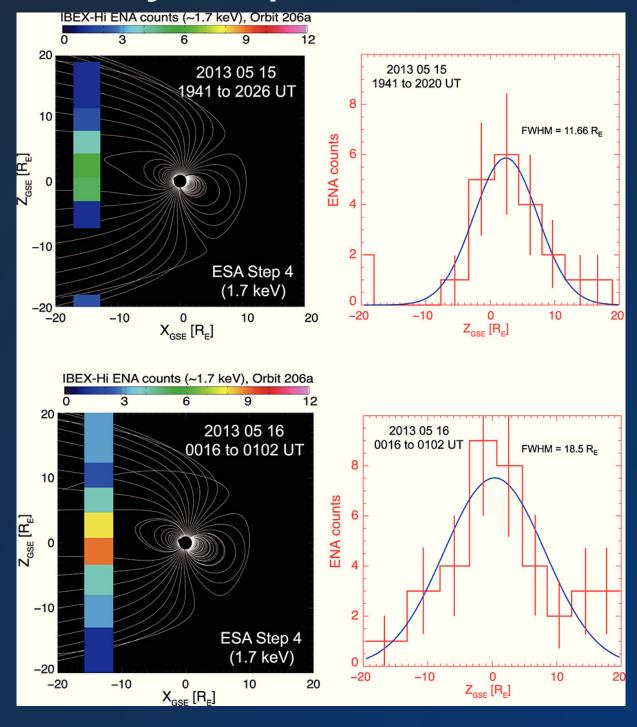


Dayeh et al, 2015 https://doi.org/10.1002/2015GL063682



Observations of the magnetotail/plasma sheet:
Development of cold dense plasma sheet during northward IMF

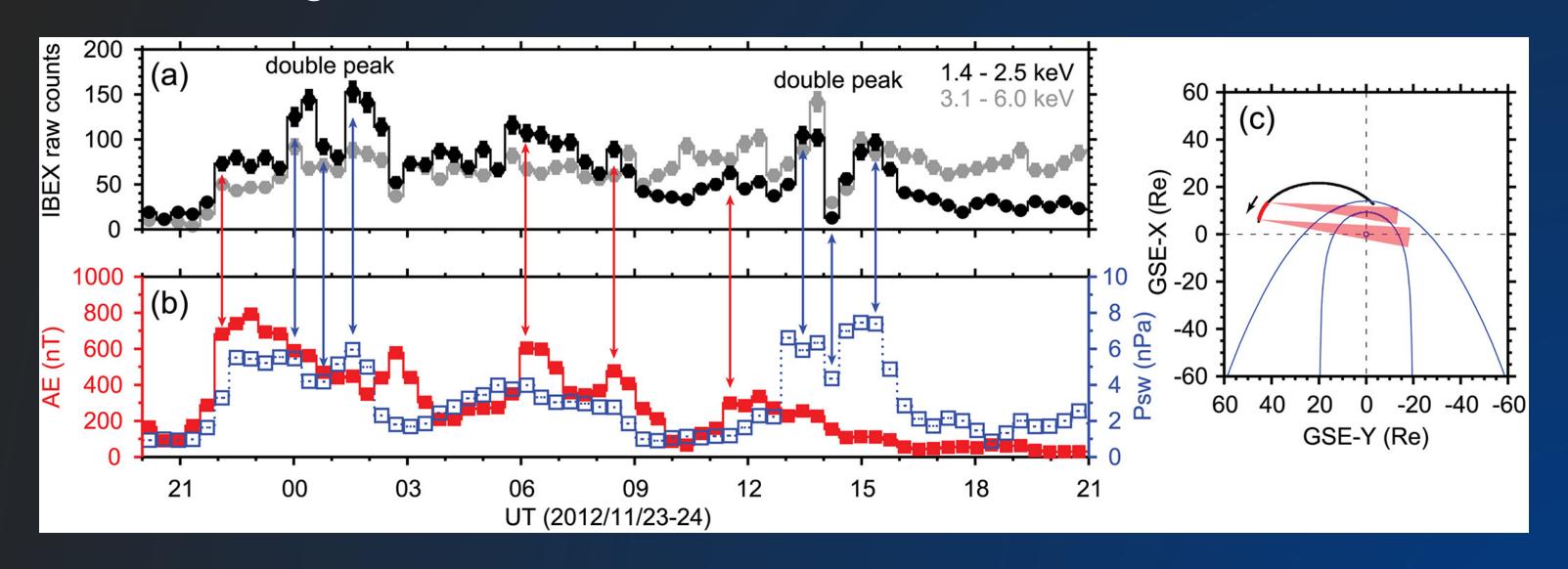




Fuselier et al, 2015 https://doi.org/10.1002/2015GL065716

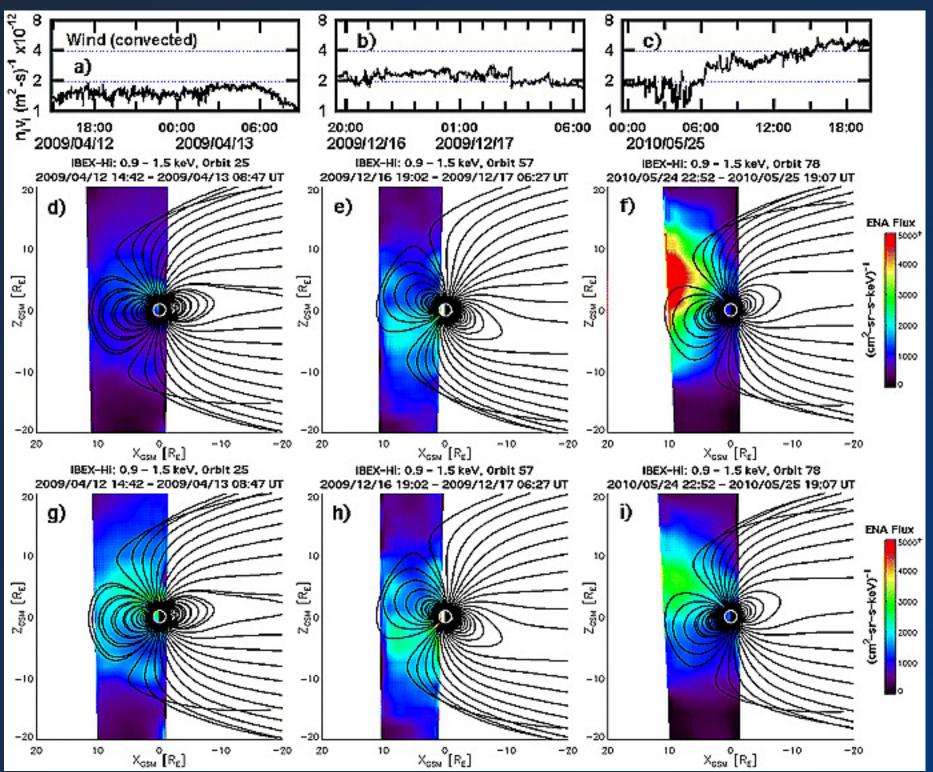


Observations of the inner magnetosphere: Correlations with the AE index indicating strong emission during FAC formation



Observations of the dayside cusps:

Cusp ENA emission related to magnetopause reconnection locations



Petrinec et al, 2011 https://doi.org/10.1029/2010JA016357

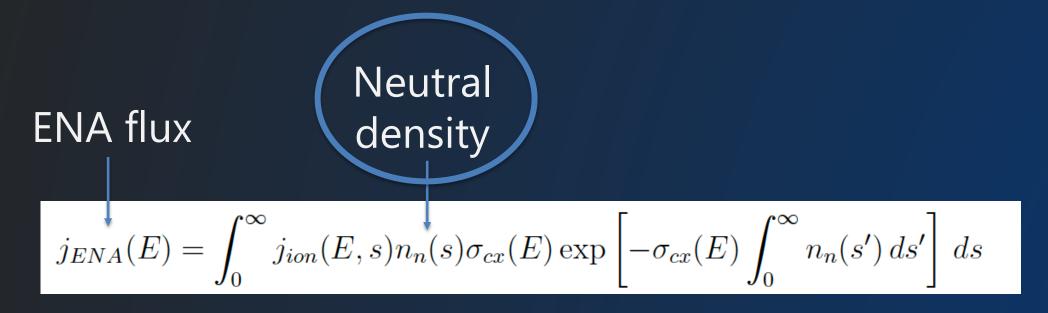


What should be considered for a lunar ENA imager?

- Spatial and temporal resolution
- Energy range
- Time delays

Increase counts with active ENA experiments

- Artificially increase neutral density to increase ENA flux
- Release could be at a modulated frequency for improved detection
- Releases of unused fuel from rockets could be used for such experiments





What should be considered for a lunar ENA imager?

- Spatial and temporal resolution
- Energy range
- Time delays

Combine with measurements from other sources