

# NAS Space Studies Board

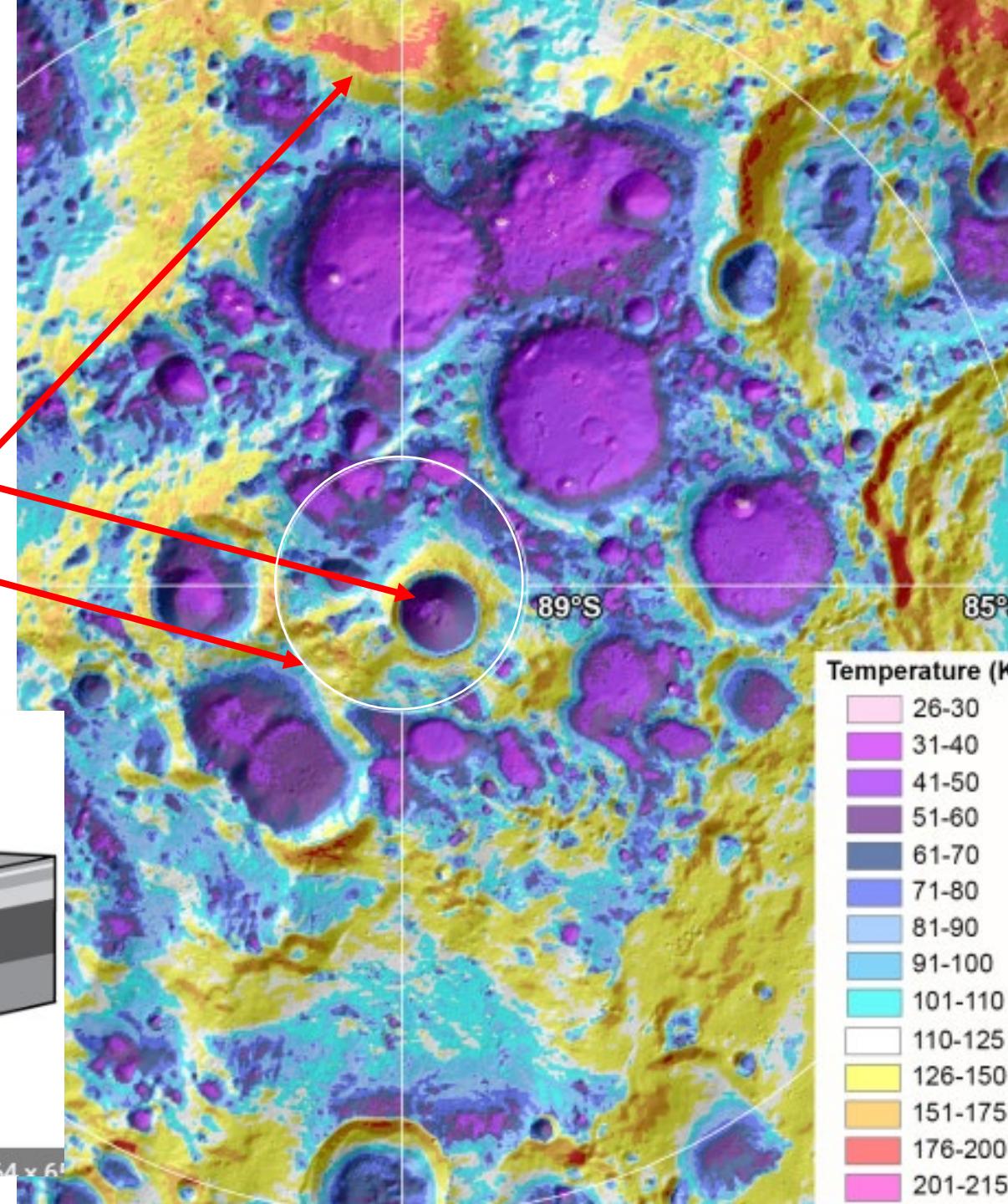
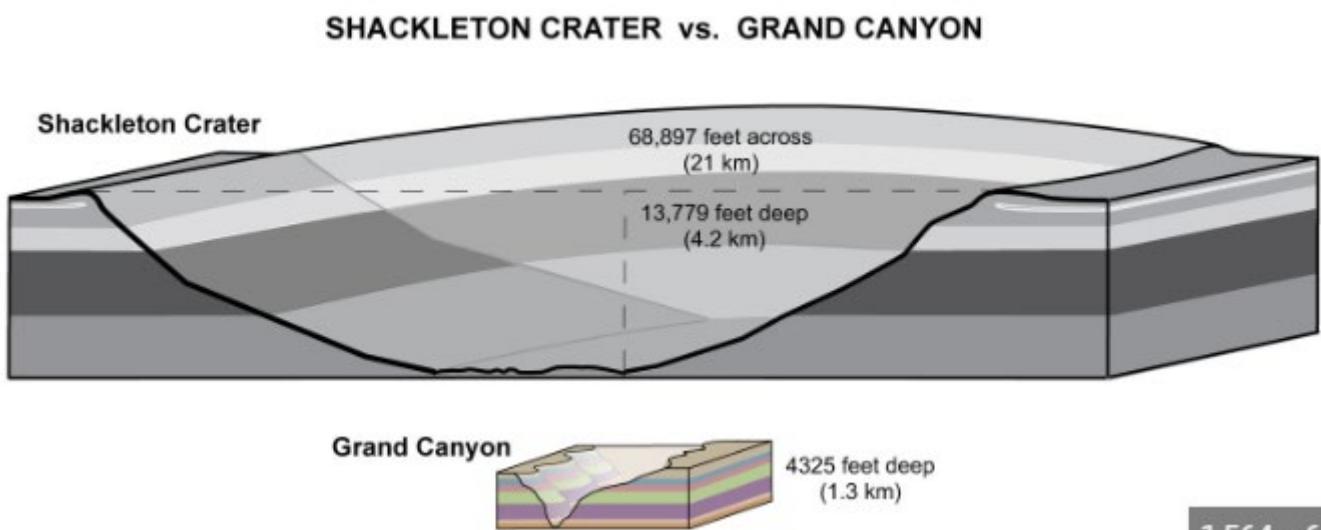
## The lunar pole - a uniquely powerful site for UVOIR and cryo telescopes

Roger Angel,  
University of Arizona

9/24/2025

# Cryo IR telescope at S. Pole

- Susante et al 2002 proposed 8 m cryo telescope in Shackleton crater, 35K
- $89^{\circ}$  south circle , radius 29 km
  - communication from Malapert mountain, always facing Earth
  - But Shackleton cold and hard to navigate:

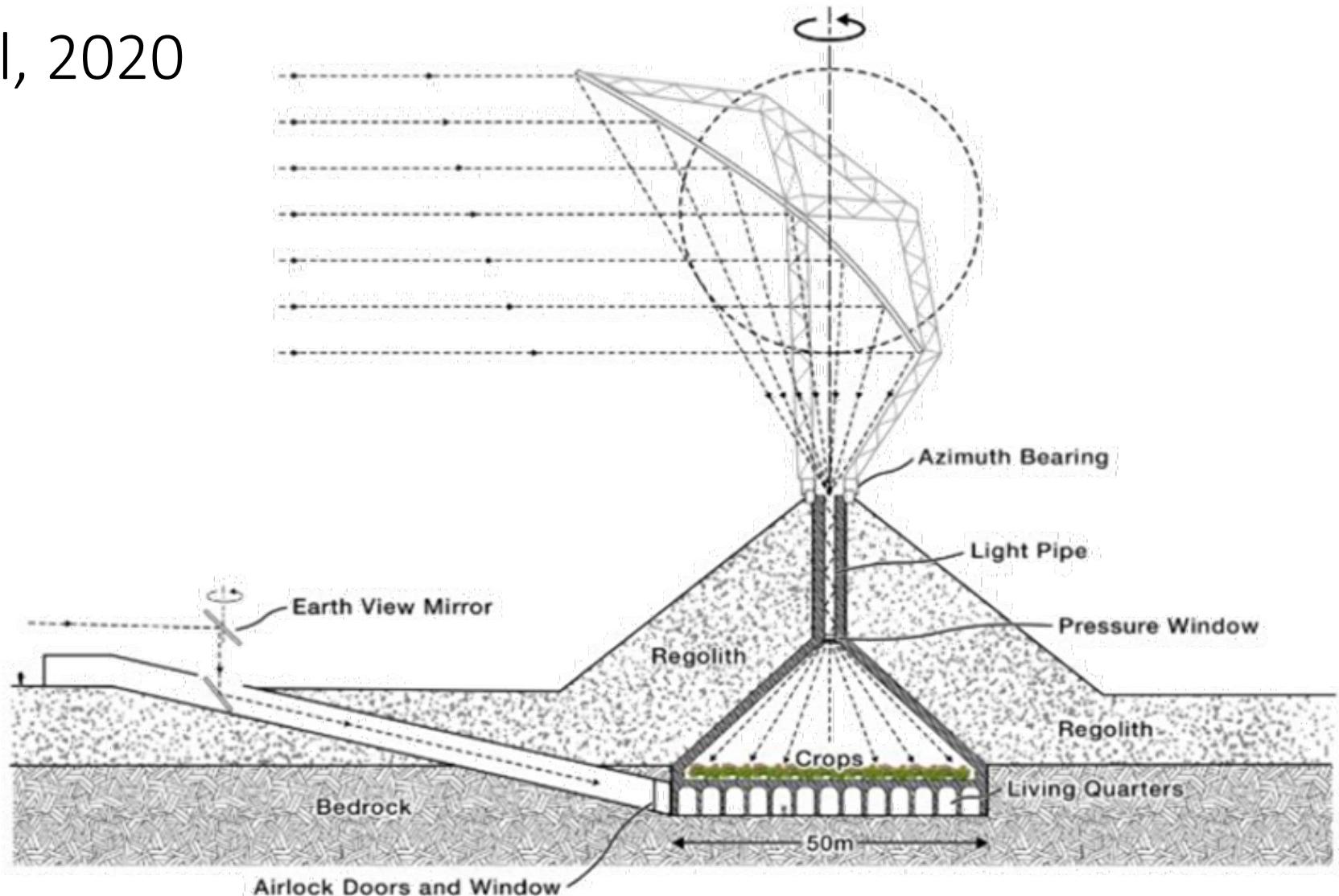


Suppose a polar base is built – for example:

*Pantheon habitat made from regolith, with a focusing solar reflector*

Woolf and Angel, 2020

- 25 m thick regolith cone balances 14 psi atmosphere inside
- Growing area 50 m diameter
- Feeds 40 inhabitants
- Some could work with infrastructure to build telescopes



# Roman Pantheon

Dome structures:

Let in light  
stable and long-lived



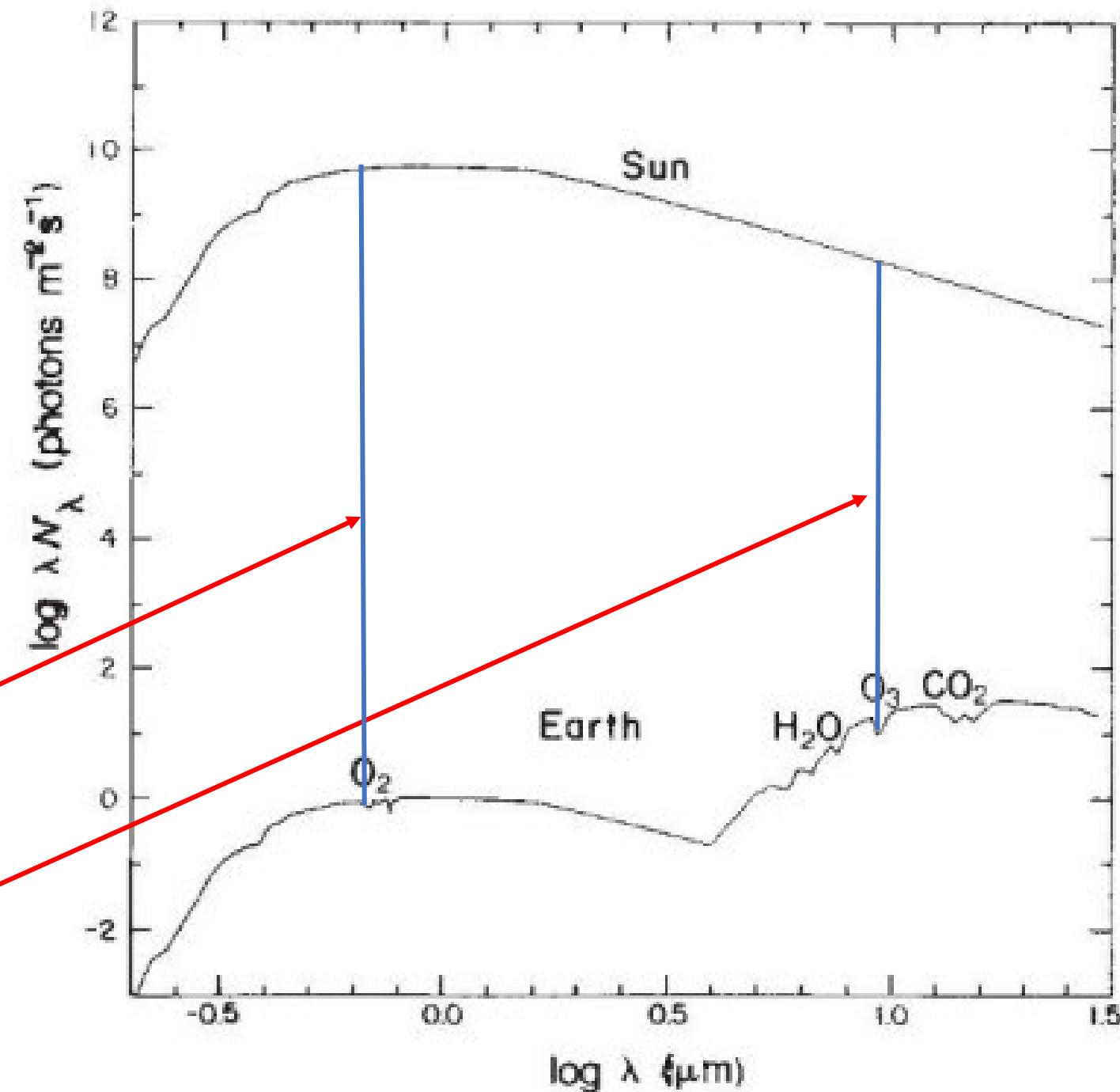
# Ozone as biosignature

*“A space telescope for infrared spectroscopy of Earth-like planets”*

Angel, Cheng & Woolf, Nature 1986

Earth 10 billion times fainter than sun in visible light

- but only 10 million for infrared, ozone absorption.



# Model of 4 x 7 m mirror telescope

- 40 m baseline
- used as Bracewell double nulling interferometer to interfere out the central star
- At 10  $\mu\text{m}$ ,  $10^{-8}$  null at 0.1 arcsec – 1 AU at 10 pc
- Also 5x resolution of JWST



# Example from close mirror array - LBT image of Io

Two phased 8.4 m mirrors, 6 m separation

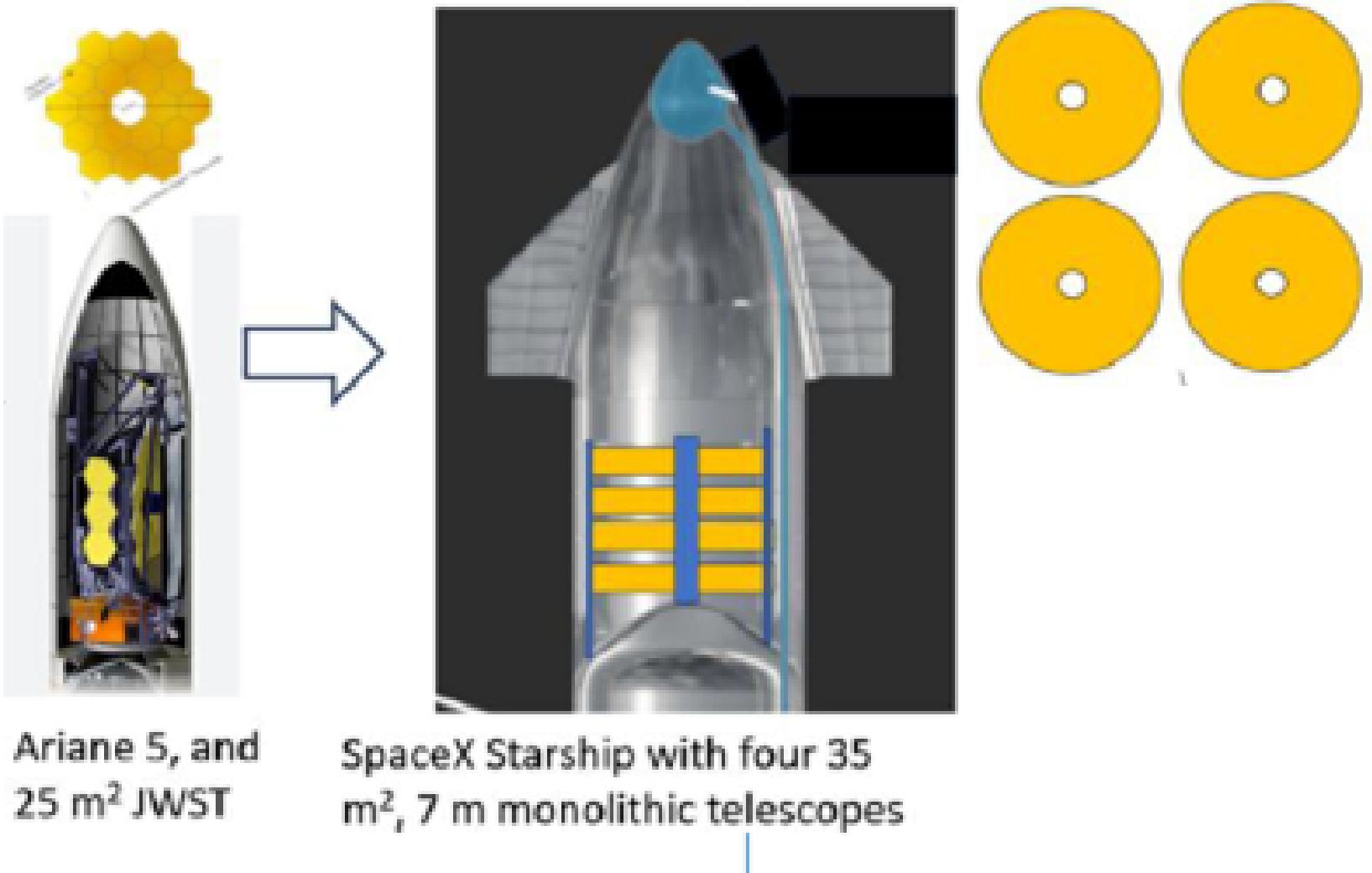
Earth rotation for full cover of Fourier plane.

With adaptive secondary mirrors, obtained optical diffraction limit at 23 m baseline



SHARK-VIS tri-color image of Io, processed with the Kraken deconvolution code.

Four 7 m mirrors could be taken as a single Starship payload

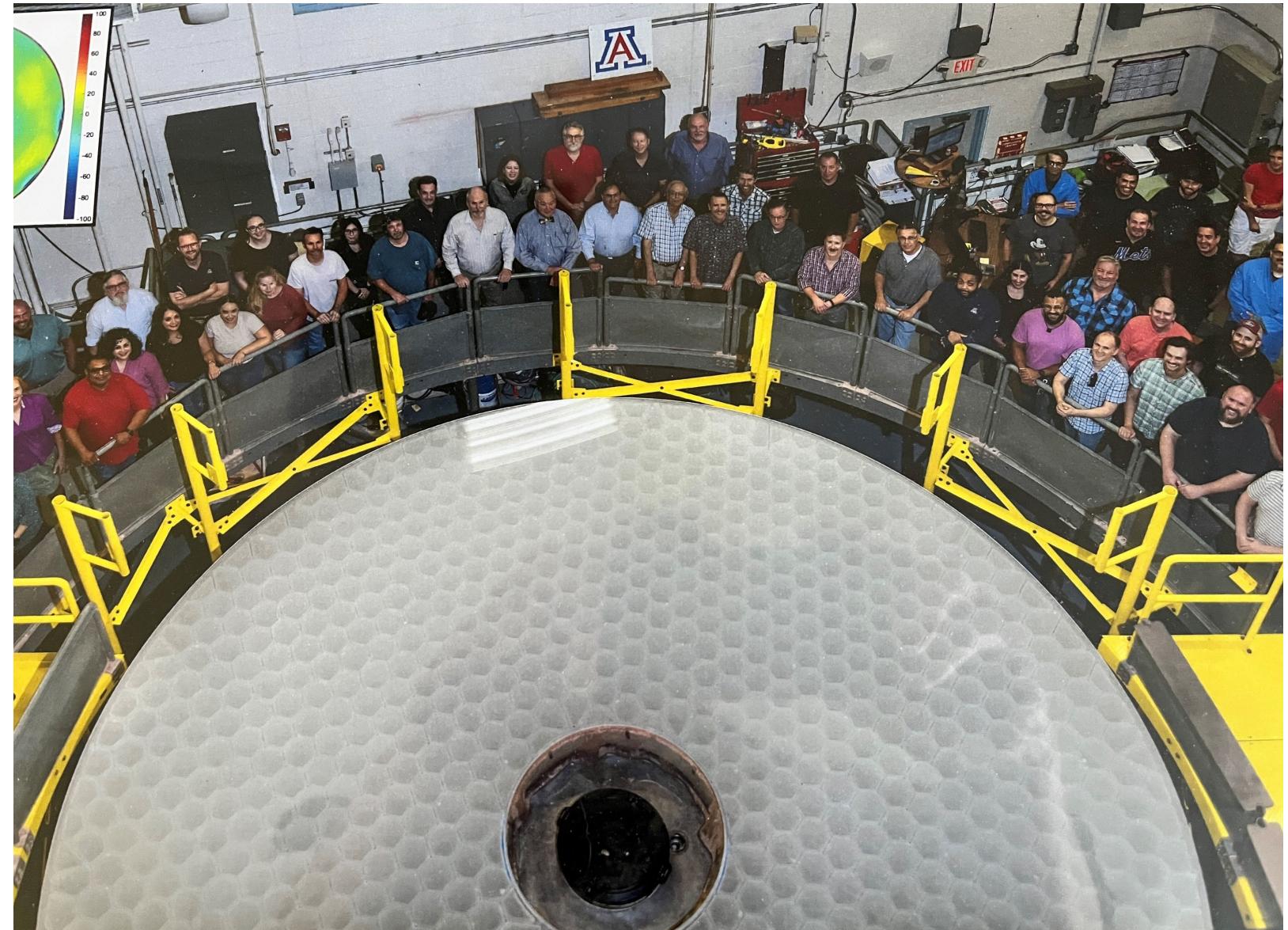


# 6.5 m mirror at the University of Arizona Caris Mirror Lab

Finished in 2024 to  
better than diffraction  
limited

Borosilicate  
honeycomb glass

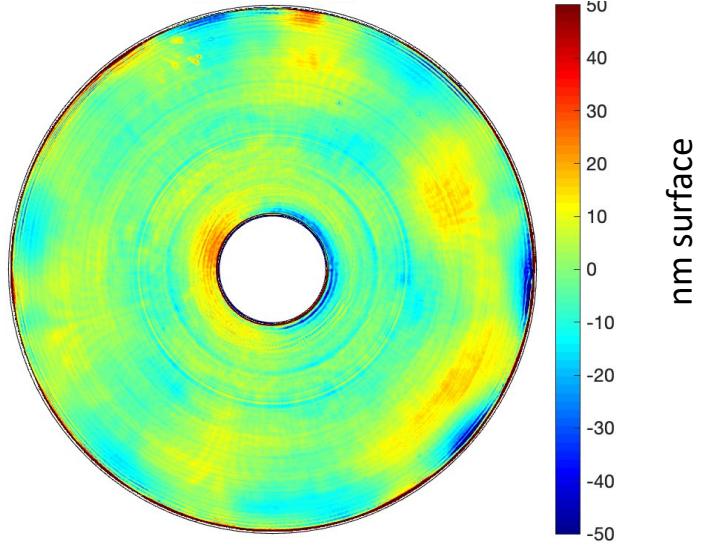
Dr Buddy Martin,  
Project Scientist for  
Mirror Polishing



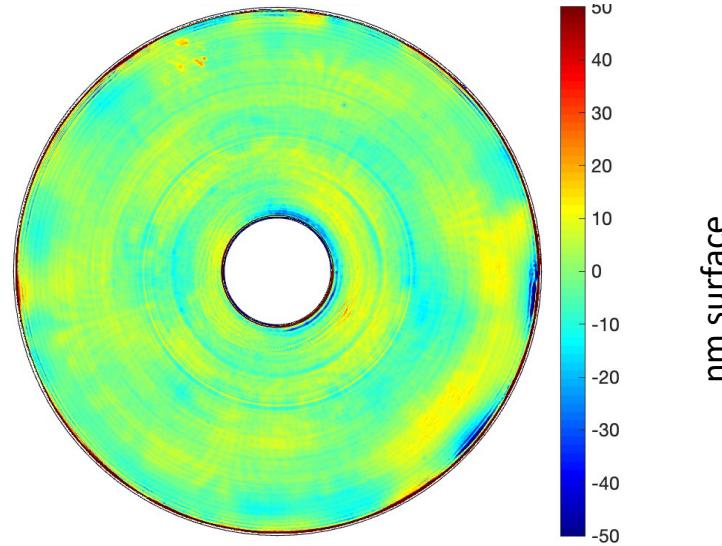
# Surface error after bending correction

Same rms waveform error as HST after its correction,  
but for 2.5 x larger aperture and thus sharper images

9.4 nm rms over clear aperture

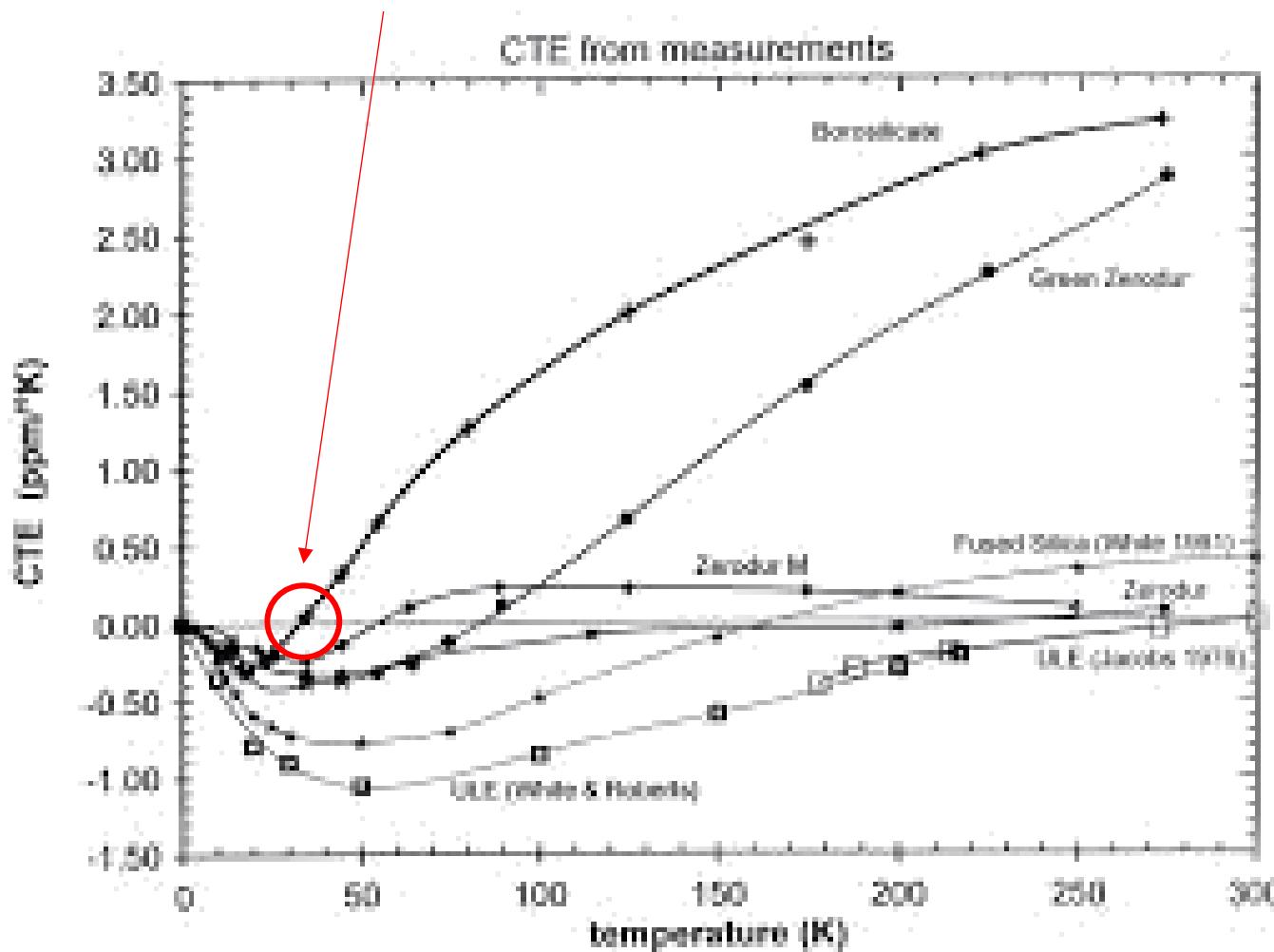


8.1 nm rms over clear aperture

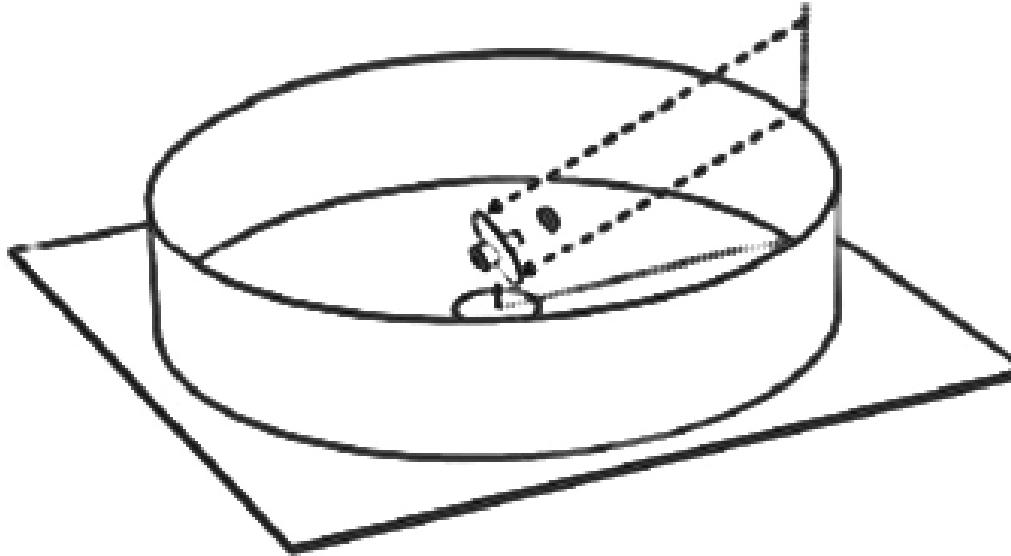


- End-to-end simulations show the as-figured surface would allow a contrast of  $<10^{-8}$  on Alpha Centauri A.

Borosilicate that we use is good to hold high accuracy, only glass type with zero expansion coefficient at  $\sim 40^{\circ}\text{K}$



Rather than trek down into Shackleton, build the telescope on flat land, near the rim



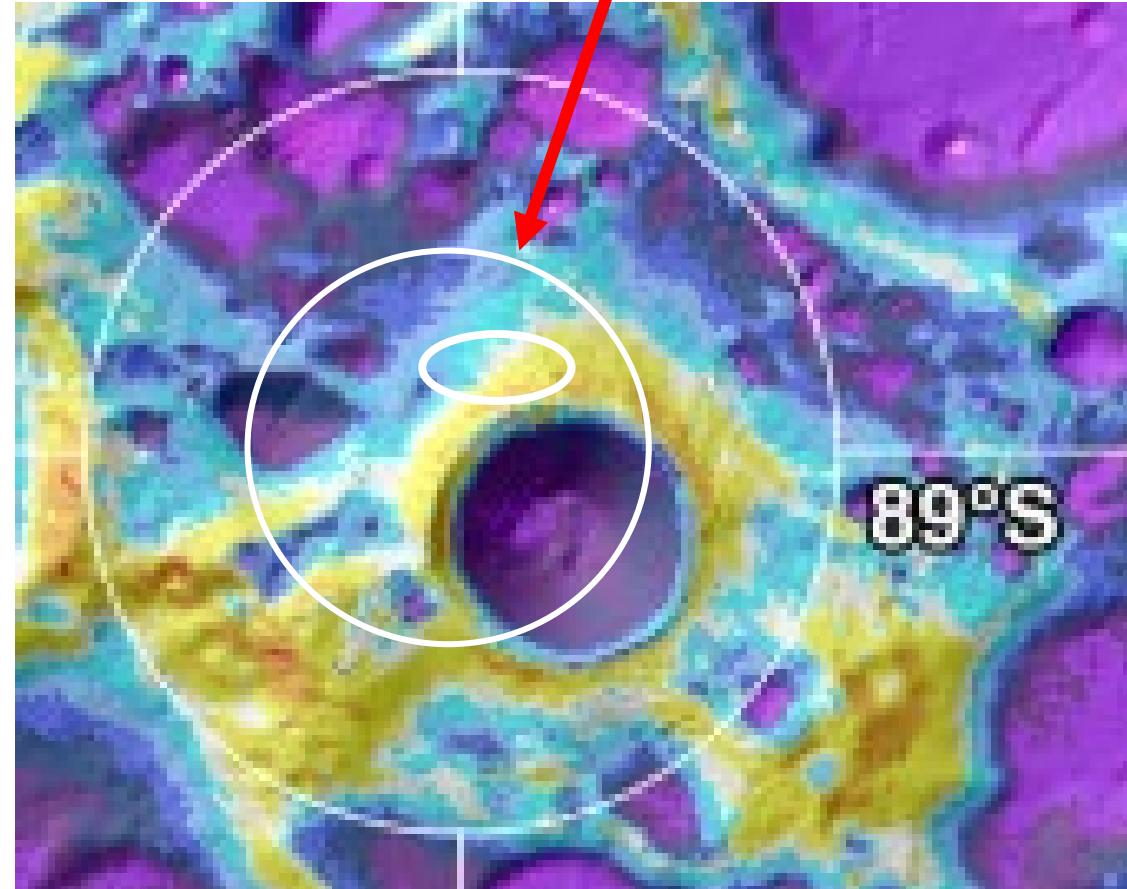
- After construction is complete, cool to 35°K, by erecting a lightweight plastic sunshield
- For a start, test with a single 6.5 m mirror here, then at the lunar pole

Need to be within 15 km ( $1/2^\circ$ ) of pole

- This is because the Moon's spin axis tilted at only  $1\frac{1}{2}^\circ$  to the ecliptic plane, so a low screen then does it.
- It can shade for dark sky at all times, while allowing access to all S sky down to 15 degree elevation

Need to protect this  $\frac{1}{2}$  degree, 15 km radius area, so telescopes won't be compromised

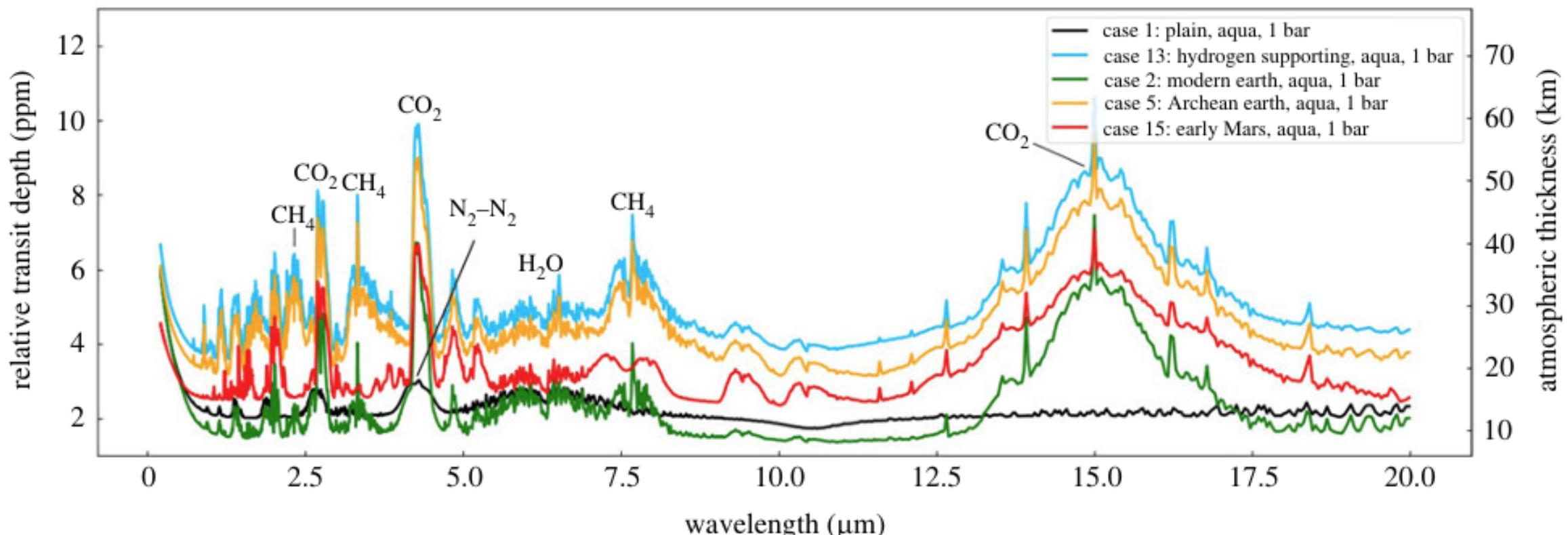
- The site could easily be spoilt if:
  - launch facilities or water mining operations were built with line of sight of the telescopes
  - Could blast regolith or debris at the telescopes.
- Plenty of water in bigger, equally cold craters at -88 or -87 latitude
- Planning must take into account the very long shadows cast when sun is only at one degree elevation
  - 100 m habitat casts 6 km long shadow!



# Exoplanet transit spectroscopy

Has no need for interferometry or big field of view

Needs huge aperture to measure features of few ppm depth  
and cold telescopes to observe 2 – 20 micron spectra

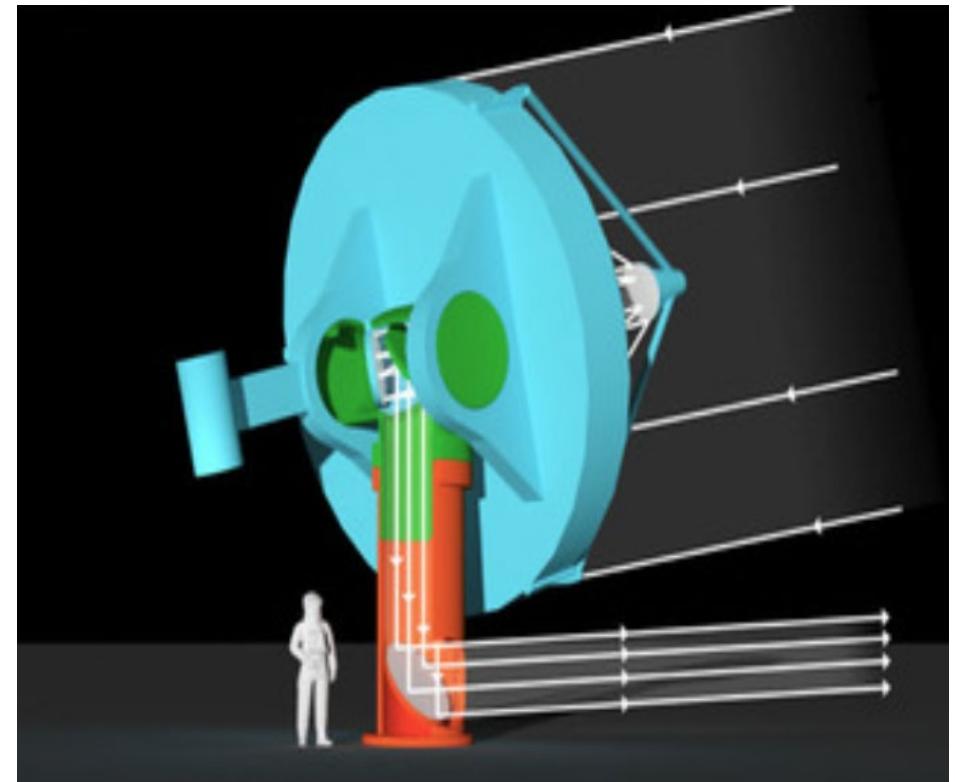


Transit spectra modelled for TOI-700 d for different atmospheric archetypes - Suissa et al

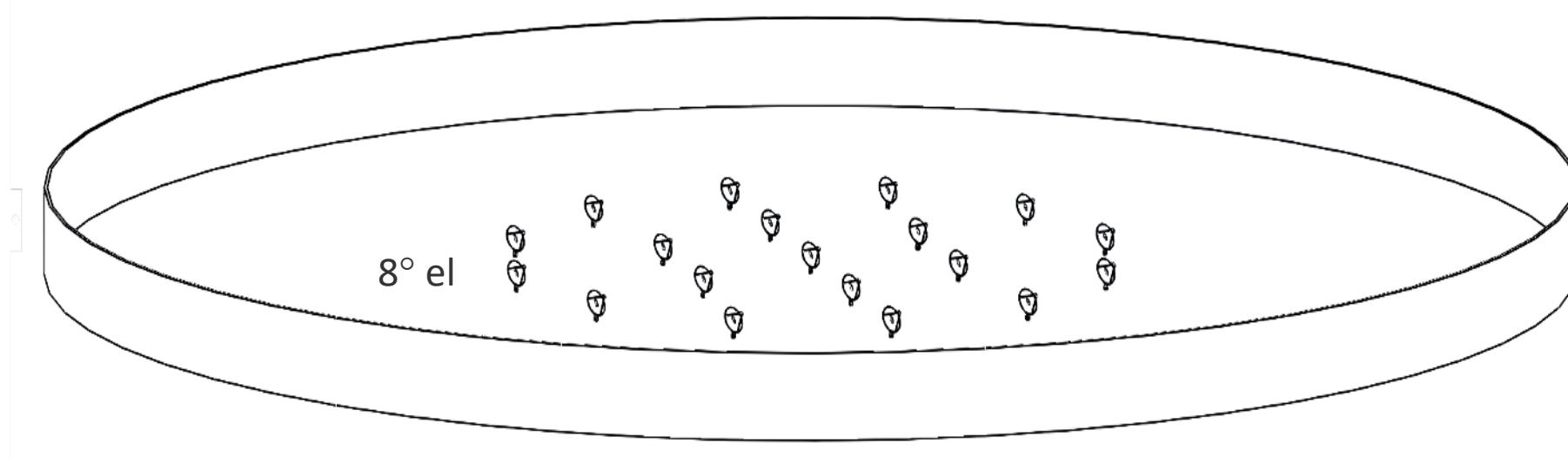
# A 600 m<sup>2</sup> array of 6.5 m telescopes at the lunar pole

Angel, Phil Trans Royal Society, 2024

- 18 telescopes, 6.5 m aperture, in 200m dia. array
- within 1/2° (15km) of lunar pole
- Perimeter screen for shade and cooling to cryo temp
- 500m diameter for unobscured access down to 8°el.
- Central beam combiner for single image from 0.4μm to 10μm.
- High res. and multi-object spectrographs to cover combined 2' field with the diffraction limited resolution of 6.5 m aperture.
- Apps. e.g. search for molecular biosignatures in transiting exoplanets, and galaxy evolution to  $> z=10$
- Cost, including delivery to the Moon by SpaceX Starship for installation using lunar base infrastructure, around \$10 billion



200 m diameter array surrounded by a cylindrical thermal shield 500 m in diameter



For location within  $\frac{1}{2}$  degree from the pole, the sun doesn't illuminate the most distant telescope of the array, and all telescopes have an unobscured view for  $\text{el.} > 8^\circ$

# Conclusions

- Something great for polar habitat occupants to build
- Get started by putting one cryo telescope, shaded, within  $\frac{1}{2}$  degree of the pole
- Make clear to lunar infrastructure builders why the closest 15 km to the poles should be protected and un-shadowed, - for astronomy.