Composition of the Venusian atmosphere

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- What we know of the composition of the atmosphere
 - Neutral atmosphere
 - Clouds structure & composition
 - Based on previous observations from space, from Earth
- What is missing, what we do not understand

What we know

Vertical information (solar/stellar occultations, GB sub-mm)

Top of the clouds (GB & space nadir; IR, UV)

Below the clouds (GB & Space, IR)



Below the clouds

- Need to look through the clouds
 - Infrared
 - Only during the night
 - Only through 'Atmospheric transparent windows'





- Limited number of species
- Limited vertical information
- Sparse information close to the surface
- Need temperature profiles

(Bézard and De Bergh, JGR 2007)

Above the clouds

- Observations
 - Solar/stellar occultation, limb: high vertical resolution, high sensitivity, different species can be detected, but at terminator (solar) or night side (stellar)
 - sub-mm GB: vertical information from absorption line profile, 70-100 km; require a temperature contrast between sounded altitude and deeper layers; Day-Night
 - GB & space nadir, UV & IR dayside: ~top of . the clouds, depending on the wavelength; shortterm spatial and temporal variations
- High variability
 - In time and space on short-term/scale
 - On long-term



Clancy et al (2012)

Above the clouds

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Encrenaz et al (2020)

Example: mesospheric SO₂



Courtesy of David Grinspoon and Carter Emmart



Venus Express found episodic injection of SO₂ into the mesosphere

- Is this an connected to volcanic activity ? (like Pinatubo)
- Or is it dynamical variability ?

(like El Niño / La Niña)

Example: mesospheric SO₂



- Mesospheric SO₂ is highly variable spatially and temporally.
- Local mesospheric SO₂ enrichment = enhanced upwelling of tropospheric air.
- What causes these upwellings?



Encrenaz et al (2013)

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Link between troposphere and above the clouds SO₂

- Decrease of 3 orders of magnitude between 30-40 km and top of the clouds
- Models cannot reproduce the strong gradient through the clouds



Missing species

- Example Sulfur cycle
- SO₃, (SO)₂, Sn
- Liquid H₂SO₄
- O₂
- Also missing laboratory supporting data



Vandaele et al, ORE (2020)



- Impact on the radiative balance impact on retrieval below the clouds
- Vertical structure Different 'modes'
- Refractive index compatible with 75-85% H_2SO_4
 - but Glory observations compatible with higher RI
 - Descent probes found traces of P, S, Cl, Fe

→UV absorber in the upper clouds still unidentified

Structure/composition highly variable wrt latitude, spatial & temporal scales



- Need in situ observations of the composition & structure
- Long-lived aerial platforms sounding different altitudes, latitudes
- High resolution spectra to identify UV absorber

Final words

- Venus surface-atmosphere should be investigated as one system:
 - Need to identify the role of surface's changes, eventually volcanism, in supplying the atmosphere
 - Need to understand the role of dynamics in the variability of trace gases and exchanges between the low atmosphere and that above the clouds
 - Need to better understand the chemistry & composition of the clouds → need also to know the gaseous composition & chemistry below the clouds
 - Below the clouds:
 - we need to characterize the spatiotemporal variability, BUT remote sensing is restricted to the night
 - We need more vertical information, BUT remote sensing restricted to specific IR windows, i.e. altitudes

→ Need of holistic approach, including different assets/platforms/instruments

- ➔ in situ observations of the composition & structure within and below the clouds
- → Orbiter investigation of the surface and atmosphere above the clouds