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SISCOE'S IDENTIFICATION OF A CULTURE GAP

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A Culture of Improving Forecasts: Lessons From Meteorology

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Space weather forecasting, like all forecasting, is challenging. Nevertheless, it has the advantage that as a young field it can look to the older field of terrestrial weather forecasting for guidance. Terrestrial weather forecasting enjoys enviable success from the perspective of space weather forecasters, who can learn from the experience of terrestrial weather forecasters as far as circumstances allow.

There is, however, a lesson from forecast meteorology that could have a bigger impact on space weather forecasting if learned not by practicing forecasters but by the general space weather research community: A significant benefit to both research and practice can result when the research community in general is familiar with the problems and the methods of the forecast community. This is a comment about the culture of the community in which space weather research and space weather operations are embedded.

TRACKING PROGRESS

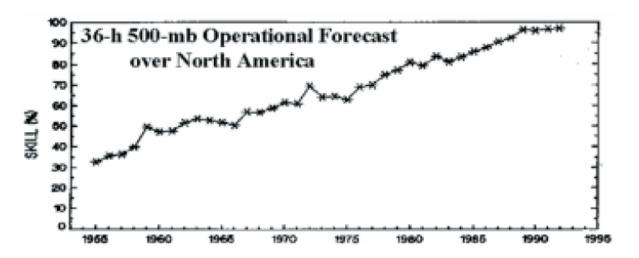


Figure 1. Forecast skill from 1955 to 1992 in the 36-hour predicted height of the 500-mbar surface over the United States, showing the escalator effect on skill that occurs once forecasting bases predictions on numerical integrations of the equations of motion (stage 9) [*McPherson*, 1994].

FORECASTING: HOW DO WE KNOW WHERE IMPROVEMENTS ARE NEEDED

With no benchmarks of skill scores communicated across the community, how can we make ANY quantitative assessment about where improvements are needed



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Center for integrated space weather modeling metrics plan and initial model validation results

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			Danelina Madala	Skill Score	Physics
	_	1. 10.44.0	Baseline Models	Data Sets	Models
		rational SW Community			
	1	Shocks and CMEs at L1	Augmented Vrsnack-	ACE	MAS+ENLIL
		a Speed	Gopalswamv ^a		
		b Arrival time	n n		
		c Bz			
		d Duration	m m		
	2	SEP Properties		GOES	UCB
S	-	a Event/No Event	PROTONS ^b	u u	OOD
Operational Metrics		b Rise Time	"		
Ž		c Peak Flux	n		
<u>a</u>		d Duration	"		
. <u>ē</u>		e Cutoff	Shea-Smart®	POES	
rat	3	Magnetic Indices			
be		a Dst	Temerin-Li d	NGDC	LFM+RCM
0		b Ap/K	ARX-McPherron		
	4	Regional Ground dB/dt	Weigel-Baker *	IMAGE (mag)	LFM+TING
	5	Radiation Belt EP fluxes			
		a GEO	Li'	LANL	RBM
	١.	b MEO and LEO	Vassiliadis ⁹	SAMPEX	
	6	lonosphere/Neutral Atmosphere			
		a "State" of ionosphere	IRI h	Digisondes	TING
	-	ntific SW Community			
	1	Solar/Coronal			
		a Coronal Hole Index	PFSS/Wang-Sheeley i	SOHO UV maps	MAS+ENLIL
		 White-light Streamer Belt Index 	PFSS/Yi-Ming Wang ^I	SOHO LASCO	"
	2	Solar Wind/IMF at L1			
		a Density	WSAk + nv = constant	ACE	MAS+ENLIL
		b Velocity	WSA	"	
(C)		c IMF - vector	WSA + IBI		
흔					
4.0	3	GEO/MEO Environment		l .	
Neti	3	a Magnetic field	Tsyganenko ^r	GOES	LFM+RCM
e Meti	3		Tsyganenko ¹ MSM ^{III} , CRRESELE ^{II}	GOES GOES/LANL	LFM+RCM LFM+RCM,RBM
nce Met	3	a Magnetic field	, ,		
cience Met	4	 a Magnetic field b Particle fluxes (ring current/rad belt) c M'pause crossing 	MSM ", CRRESELE "		LFM+RCM,RBM
Science Metrics		a Magnetic fieldb Particle fluxes (ring current/rad belt)	MSM ", CRRESELE "		LFM+RCM,RBM
Science Meti		Magnetic field Particle fluxes (ring current/rad belt) M'pause crossing MI Coupling Polar Cap Potential	MSM ^m , CRRESELE ⁿ Shue °	GOES/LANL "	LFM+RCM,RBM LFM+RCM
Science Met		a Magnetic field b Particle fluxes (ring current/rad belt) c M'pause crossing MI Coupling a Polar Cap Potential b Polar Cap Boundary	MSM ^m , CRRESELE ⁿ Shue ^o Weimer ^p	GOES/LANL "	LFM+RCM,RBM LFM+RCM
Science Met		a Magnetic field b Particle fluxes (ring current/rad belt) c M'pause crossing MI Coupling a Polar Cap Potential b Polar Cap Boundary c Field Aligned Currents (2D)	MSM ^m , CRRESELE ⁿ Shue ^o Weimer ^p Weimer	GOES/LANL "	LFM+RCM,RBM LFM+RCM LFM+TING "
Science Met	4	a Magnetic field b Particle fluxes (ring current/rad belt) c M'pause crossing MI Coupling a Polar Cap Potential b Polar Cap Boundary c Field Aligned Currents (2D) d Particle precipitation	MSM ", CRRESELE " Shue " Weimer P Weimer Weimer Weimer	GOES/LANL "	LFM+RCM,RBM LFM+RCM LFM+TING " LFM+TING+MIC
Science Met		a Magnetic field b Particle fluxes (ring current/rad belt) c M'pause crossing MI Coupling a Polar Cap Potential b Polar Cap Boundary c Field Aligned Currents (2D)	MSM ", CRRESELE " Shue " Weimer P Weimer Weimer Weimer	GOES/LANL "	LFM+RCM,RBM LFM+RCM LFM+TING " LFM+TING+MIC

^a Vrsnak and Gopalswamy (2002), Gopalswamy et al. (2001), Bothmer and Rust (1997), and Owen and Cargill (2002).

b Balch (1999).

c Shea and Smart (1990)

d Temerin and Li (2002).

⁹ Weigel and Baker (2003).

f Li et al. (2003).

⁹ Vassiliadis et al. (2004). ^hBilitza (2001, 2003).

Wang and Sheeley (1992).

Wang et al. (1997).

Wang and Sheeley (1992); Arge and Pizzo (2000); Arge et al. (2004). Tsyganenko (1995, 2003).

Magnetospheric Specication Model from AF-GEOSpace. CRRESradiation belt electron model from AF-GEOSpace.

Shue et al. (1997,1998).

P Weimer (1996), Weimer (2001a, b).

^q Air Force Statistical Auroral Models from AF-GEOSpace.

CULTURE ISSUES

- Would forecasting capabilities improve if we moved toward open source models?
 - Improves collaborative approach
- What are the metrics tracked by the community and who is responsible for tracking these metrics over the long term?
- What are the user-specified requirements for space weather forecasting?
- Can funding lines be created with <u>the PRIMARY</u>
 <u>OBJECTIVE</u> of improving forecasts?