



Near-Earth Object Surveyor and Planetary Defense

Prof. Amy Mainzer, Survey Director



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NEOs: The Critical Questions



- Need to know **when** impacts could occur and **how bad** they will be
 - **When**: Comes from finding objects & determining good orbits for them
 - **How bad**: Comes from measuring the impact energy (KE)
- Impact energy scales as $KE = 1/2 \text{ mass} \times \text{velocity}^2$
 - Velocity comes from orbit
 - Mass = density x volume = density x diameter³
 - → Impact energy depends strongly on diameter (size)



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The Future of Planetary Defense



- **Present events demonstrate that rare disasters can sometimes happen**
 - A modest effort to be prepared can significantly improve outcomes
- **We can provide best chance of mitigating an asteroid or comet impact if we know about it well in advance: ideally years to decades**
- **Surveys that combine great sensitivity with large field of view offer best chance of comprehensive coverage**
 - Planetary defense is a “team sport”: experience has shown it takes multiple facilities with complementary capabilities to provide thorough coverage
 - Ideally want basic physical property measurements for large numbers of objects to quickly quantify any potential hazard
 - NEOs can become unobservable for long periods of time, and are often brightest at the discovery apparition



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Priorities for Planetary Defense



- **Finish survey of 100m+ NEAs & calculate basic physical properties for them (diameter, albedo, spin state, thermal inertia)**
 - Evaluate future survey strategies based on population statistics found in the course of performing this survey
 - Follow up targets of concern: radar, spectroscopy, dense lightcurves, etc.
- **Provide an all-sky monitoring capability for comets**
- **Strengthen and expand international collaborations**
- **Ensure robust lab/modeling capabilities for predicting impact effects**
- **Demonstrate reconnaissance missions for rapid in situ characterization of NEAs/comets**
- **Demonstrate additional mitigation technologies e.g. gravity tractor**



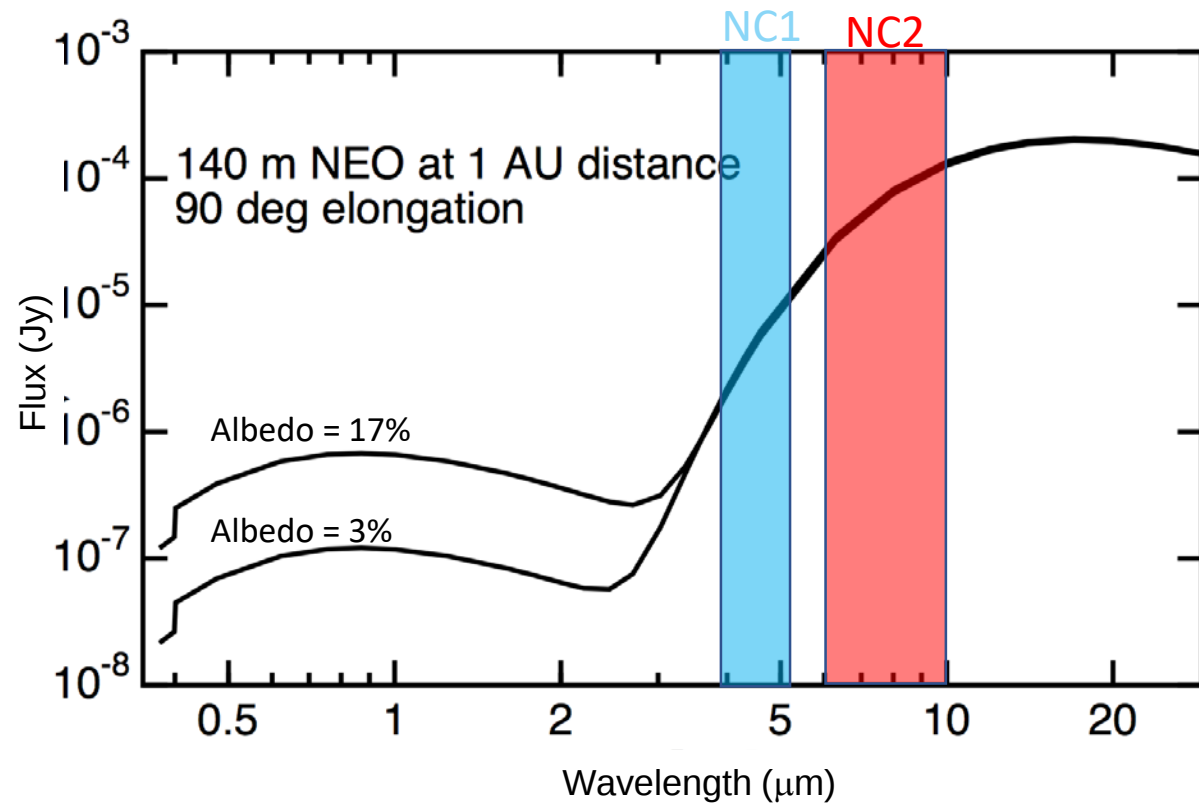
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Asteroid Spectral Energy Distributions Peak in the Infrared

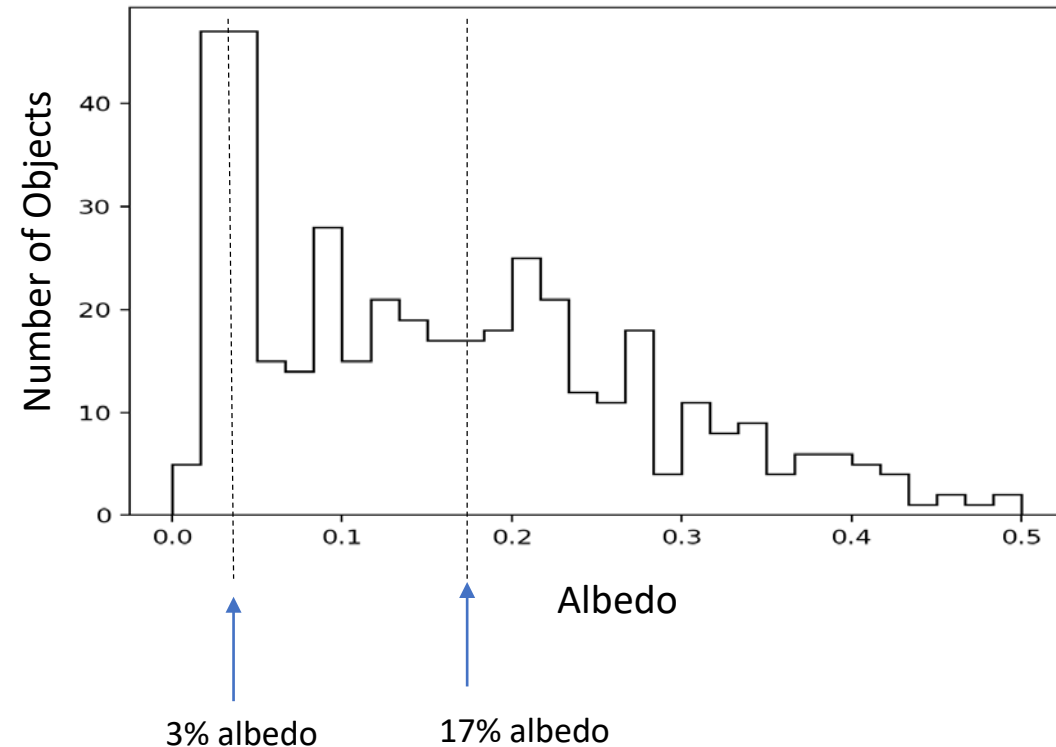


NEO Albedo Distribution



- There is an important population of very dark NEOs (3% albedo), and a population of less dark ones (17% albedo)
- Visible surveys can find bright NEOs, but struggle to find dark ones.
 - Finding >90% of NEOs >140m requires reaching 90% completeness at $H < 23$ mag (Wright et al. 2016)

12-um Selected Sample of NEOs Detected by NEOWISE



Mission Summary:

NEO Surveyor Is a Planetary Defense Mission



- NEO Surveyor is designed to respond to the objectives of NASA's Planetary Defense Coordination Office by **detecting, cataloging, and characterizing NEOs**
- Mission is **optimized** for the task of finding and characterizing the impact risks posed by potentially hazardous objects (PHOs), both as individual objects and as populations – quantifies the risk
- NEO Surveyor provides critical **decision support** for stakeholders who must assess the risks of NEO impacts to Earth and must identify potential mitigation strategies



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Mission Objectives – Introduction



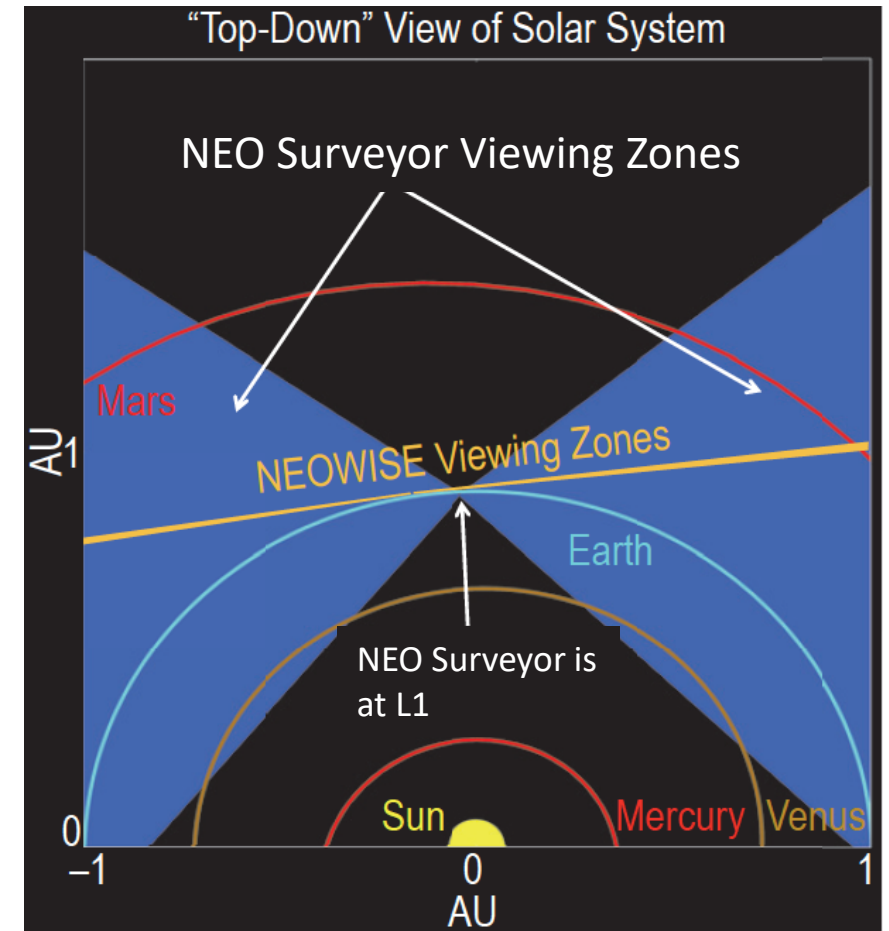
- **Mission objectives break into groups:**
 - Finding potentially hazardous asteroids (PHAs) and comets
 - PHA = asteroid with orbit that comes within 0.05 AU of Earth's orbit
 - Informing us about the nature of NEOs, including their subpopulations and related source regions
 - Providing the capability for detailed follow-up of individual objects with unusually high impact risks



NEO Surveyor Is Optimized for NEOs



- The single instrument consists of a 50-cm infrared (IR) telescope that collects images simultaneously at 4-5 and 6-10 μm
- It surveys a large fraction of the entire sky, focusing on the regions where PHAs spend most of their time
- Spacecraft is stationed at Sun-Earth L1 Lagrange point: nominal 5-year lifetime allows for discovery of large fraction of PHAs >140 m (goal: 12-year lifetime)



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Space Dynamics
LABORATORY
Utah State University



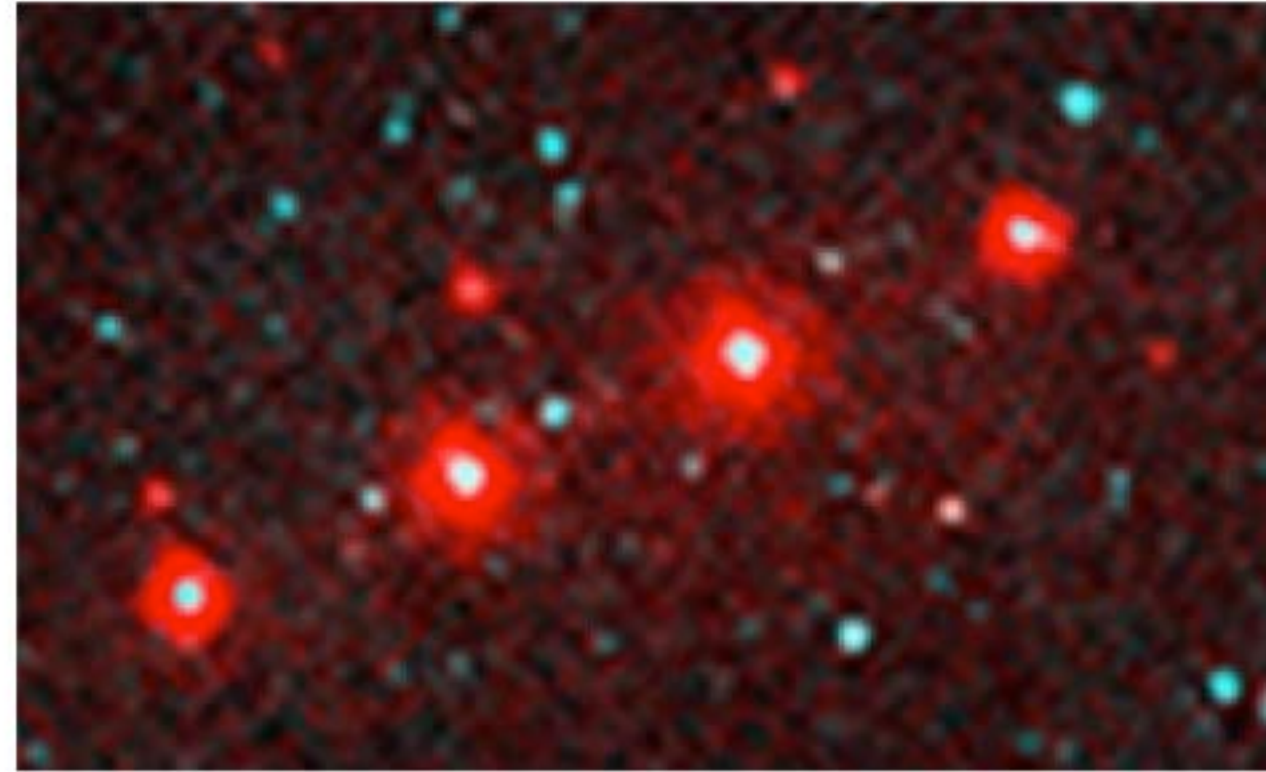
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Mission Design



- L1 is close enough to Earth to allow full images to be downlinked and searched using optimal algorithms
- Single instrument
- Yet far enough from Earth to support passive cooling, thus avoiding costly complex cryocoolers and life-limiting cryostats
- Survey pattern is fixed, highly repeatable, and optimized for potentially hazardous asteroid and comet discovery: single operating m



Asteroid detections from NEOWISE

- NEOs >140m only

